Astro 596/496 NPA Lecture 5 Sept. 2, 2009

Announcements:

- PF 1 due Friday ≤12noon see tips on next slide
- World-famous cosmologist in the house!
 Paul Steinhardt, Princeton
 "The Endless Universe"
 Physics Colloquium tomorrow, 4pm Loomis 141

Last time: nuclear structure and decays

Q: decay modes?

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Q: radioactive decay time behavior?

Tips for PF1 Optional Reading (Asplund et al 2005) Abundance Notation

very commonly used

$$[A/B] \equiv \log_{10} \left[\frac{(A/B)_{\text{observed}}}{(A/B)_{\text{solar system}}} \right]$$
(1)

e.g.,: if a star has $Fe/H_* = 0.01Fe/H_{\odot}$, then $[Fe/H]_* = -2$ note: [A/B] is a *logarithmic* measure of abundance i.e., [A/B] is a "decimal exponent \rightarrow really dimensionless, but "units" sometimes called "dex"

also used

$$[A] = 12 + \log_{10}(A/H) = \log_{10}(10^{12}A/H)$$
(2)

s e.g., since $(Fe/H)_{\odot} = 3 \times 10^{-5}$, then [Fe] = 7.5also dimensionless logarithmic measure, "units" sometimes also called "dex"

Radioactive Decay Rate

const decay probability P per unit time:

$$\frac{dP}{dt} = \lambda = const \tag{3}$$

and so

$$\frac{dn}{dt} = -n \ \frac{dP}{dt} = -\lambda n \tag{4}$$

decay lifetime ("mean life") $\tau = 1/\lambda$ solution: $n = n_0 e^{-t/\tau}$ also write $n = n_0 2^{-t/t_{1/2}} = n_0 (e^{\ln 2})^{-t/t_{1/2}} = n_0 e^{-t \ln 2/t_{1/2}}$ "half-life" $t_{1/2} = \tau \ln 2$

ω www: supernova 1987A brightness vs time = ''lightcurve''

Astrophysical Gamma Decays

famous example: ²⁶Al decay chain

$$\overset{26}{\operatorname{AI}} \rightarrow \overset{26}{\operatorname{Mg}}^{*} + e + \nu$$

$$\overset{26}{\operatorname{Mg}}^{*} \rightarrow \overset{26}{\operatorname{Mg}}^{g.s.} + \gamma$$

$$(5)$$

www: ²⁶Al decay scheme $t_{1/2}({}^{26}Al) = 0.7 \text{ Myr}; t_{1/2}({}^{26}Mg^*) = 0.5 \text{ ps}$ Q: decay modes? overall timescale?

search sky for 1.8 MeV decay γ rays www: Galactic coordinates www: COMPTEL 1.8 MeV sky

decay line seen!

^{*} Q: basic, model-independent implications? Q: $t_{1/2}$ (²⁶AI) vs relevant astrophysical timescales-implications?

²⁶Al Gamma Lines: Smoking Gun of Nucleosynthesis

But we can push further:

- radioactive decay is exponential in time: radionuclides lost after a few $t_{1/2}$
- we know $t_{1/2}(^{26}\text{AI}) = 0.7 \text{ Myr}$
- age of Earth/Sun ~ 5 Gyr, age of Galaxy (i.e., Milky Way disk) ~ 10 Gyr
 ⇒ the ²⁶Al we see was born "yesterday"

nucleosynthesis is ongoing in the Galaxy today!

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Q: implications of ²⁶Al signal morphology (sky pattern)?

Implications of ²⁶Al Gamma-Line Sky Morphology

- ²⁶Al signal traces Galactic plane this is where star birth (and death) occur
- ⇒ (at least some) nucleosynthesis is associated with stars
- ²⁶Al signal is *diffuse* not collection of points (tricky to show because angular resolution bad)
 ⇒ ²⁶Al is in the *interstellar medium* (ISM) and since ²⁶Al short-lived compared to stars
- stars eject fresh nucleosynthesis products which mix into the ISM

INTEGRAL can measure line shape, position well www: ²⁶Al line shift vs Galactic longitude *Q: summarize result? implications? possible interpretation?*

Nuclear Reactions

$$\begin{array}{l} \hline \text{Notation} \\ A+b \rightarrow c+D = A(b,c)D \\ \text{usually } b,c \in p,n, \alpha = {}^{4}\text{He}, d = {}^{2}\text{H}, \gamma \\ \text{e.g., } {}^{12}\text{C} + {}^{4}\text{He} \rightarrow {}^{16}\text{O} + \gamma \text{ is } {}^{12}\text{C}(\alpha,\gamma){}^{16}\text{O} \end{array}$$

Kinematics

preferred frame: center of mass Diagram $m_1 \vec{v}_1 + m_2 \vec{v}_1 = (m_1 + m_2) \vec{V}_{CM}$ (non-rel) relative vel: $\vec{v} = \vec{v}_1 - \vec{v}_1$

can show:

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$$KE = \frac{1}{2}m_1v^2 + \frac{1}{2}m_2v^2$$

= $\frac{1}{2}\mu v^2 + \frac{1}{2}(m_1 + m_2)V_{CM}^2 = (KE)_{in CM} + (KE)_{of CM}$

where reduced mass $\mu = m_1 m_2 / (m_1 + m_2)$

Energy conservation

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for
$$a + b \rightarrow c + d$$

 $E_f = E_i$ including rest mass energy!
 $(m_c + m_d)c^2 + (KE)_f = (m_a + m_b)c^2 + (KE)_i$
 $(KE)_f = (KE)_i + (m_a + m_b - m_c - m_d)c^2 = (KE)_i + Q$
with reaction "Q" value:

$$Q = \text{initial masses} - \text{final masses}$$
(6)
$$= [m_a + m_b - (m_c + m_d)]c^2$$
(7)
$$= \Delta_a + \Delta_b - \Delta_c - \Delta_c$$
(8)

where last line uses mass defects & baryon conservation

exothermic reactions: Q > 0 (mass energy released) can occur for any $(KE)_i$ endothermic reactions: Q < 0need $(KE)_i > |Q|$ to go: \Rightarrow there is a "threshold" energy

Reaction Physics

classify reactions according to timescale τ :

Direct: $\tau \simeq t_{cross} = r_{nuke}/v$ retain "memory" of initial states "single step" formation of product

Compound: $\tau \gg t_{\text{cross}}$

"forget" initial state form intermediate QM state (resonance) "compound" nucleus $A + b \rightarrow W \rightarrow c + D$ decays, emitting particles of final state

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