

Astro 210  
Lecture 6  
Sept 3, 2010

Announcements

- HW1 due now
- HW2 available, due in class next Friday
- **register** your iClicker; link on course webpage

Last time: a tale of two cosmologies

- Geocentric
  - Q what's that? how does it explain sunrise? retrograde?*
- Heliocentric
  - Q what's that? how does it explain sunrise? retrograde?*

└

Today: geocentric vs heliocentric cagematch!

# Tycho Brahe 1546-1601: Danish Astronomy Extraordinaire

in youth: observed “nova stella” (supernova) [www: Tycho sketch](#)

→ heavens corruptible!

observed Sun, Moon, planets for 20 years: careful, accurate data

but not a good number cruncher

→ like any good professor: made grad student do the work!

## Johannes Kepler 1571–1630: Harmony of the Worlds

Analyzed Tycho’s data for **20 years**(!), especially Mars motions

used heliocentric model with circles

but observations didn’t quite agree

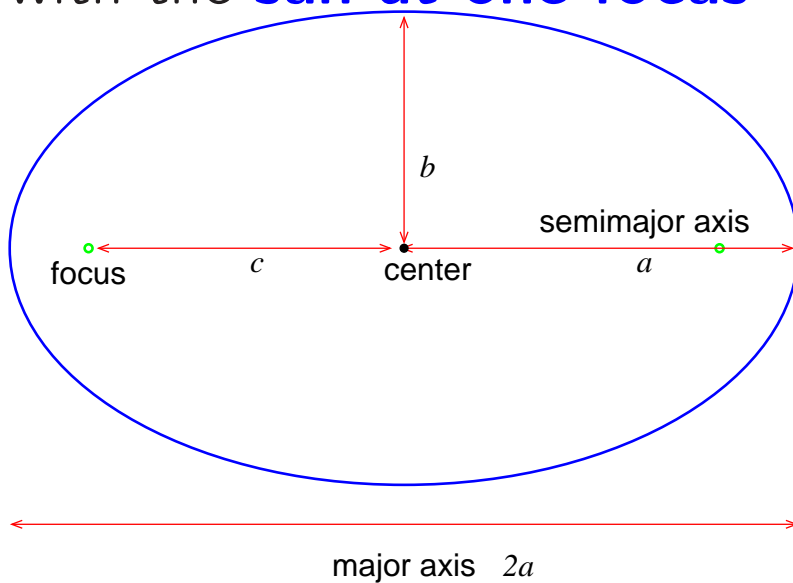
a small error (few arc min!) remained...took seriously

↷ → after years of trial & error:

completely & accurately described planet orbits

# Kepler I: Law of Ellipses

each planet's orbit is an **ellipse**  
with the **sun at one focus**

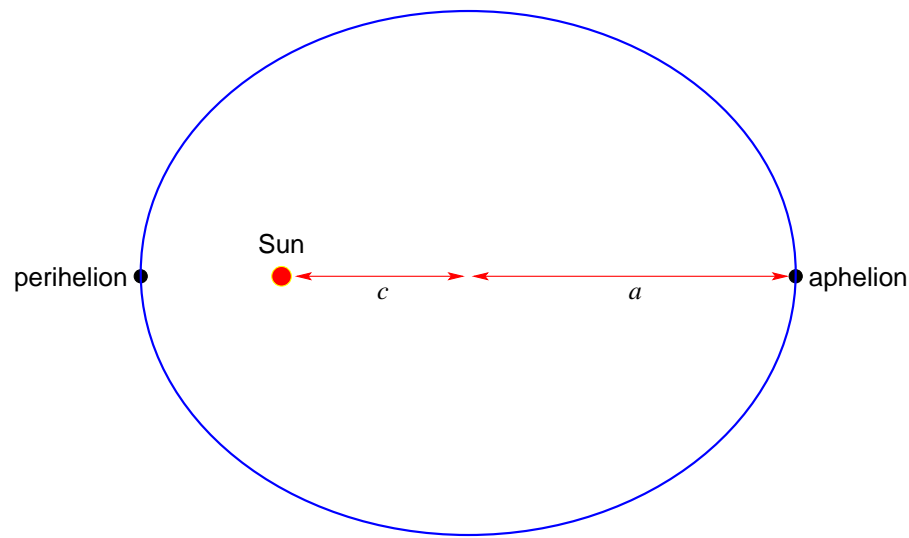


anatomy: foci, **semi-major axis**  $a$ , focal length  $c$

ω can be characterized by  $a$  and **eccentricity**  $e = c/a$

Q: what is  $e = 0$ ?  $e = 1$ ?

Kepler I: orbit is **ellipse** with **sun at one focus**



Orbit anatomy

*aphelion*: farthest point from Sun

*perihelion*: closest point to Sun

*Q: what is aphelion distance in terms of  $a$  and  $e$ ?*

$$r_{\text{ap}} = a + c = a + a\frac{c}{a} = (1 + e)a \quad (1)$$

⌞ *Q: If the Sun's at one focus, what's in the other focus?*

*Q: What does Kepler I not say about orbits?*

At the other focus: nothing! (sorry!)

Note: Kepler I only gives orbit *shape*  
but says *nothing* about how orbit evolves in time  
→ need more info to fully describe orbit, hence...

## Kepler II: Law of Equal Areas

a straight line from the planet to the sun  
sweeps out equal areas in equal times

*diagram: sketch areas*

www: area animation

note that this amounts to telling about speed of planet

## iClicker Poll: Kepler II and Planet Speed

When does a planet move the *fastest*?

- A** When it is closest to the Sun
  - B** When it is farthest from the Sun
  - C** Trick question! In vacuum of space, planet speeds must be constant
- 

↘ Q: *This still doesn't fully characterize an orbits—why not?*

Kepler I gives orbit shape in space

Kepler II gives orbit evolution over time

but haven't yet connected the two:

how does spatial character (e.g., semimajor axis  $a$ )  
relate to time character (e.g., period  $P$ )?

Need one last law...



## Kepler III: The Mighty Equation

period  $P$  and  $a$  are related:

$$P^2 \propto a^3$$

$\Rightarrow P^2/a^3 = \text{const}$ , and since must hold for Earth:

$$P_{\text{yrs}}^2 = a_{\text{AU}}^3 \quad (2)$$

*Q: ok for earth?*

where  $P$  written in years,  $a$  in AU

Very powerful! e.g.:

Asteroids exist with orbits inside 1 AU (and some cross 1 AU!!)

☉ www: inner solar system objects--in real time!

## iClicker Poll: Kepler III

Kepler III:  $P_{\text{yrs}}^2 = a_{\text{AU}}^3$

Consider an asteroid with an orbit entirely inside 1 AU

Is its period longer or shorter than a year?

- A**  $P > 1$  yr, no matter eccentricity  $e$
- B**  $P < 1$  yr, no matter what  $e$
- C** can't answer without knowing  $e$

## Kudos to Kepler

Several points worth noting...

★ An amazing discovery—mathematics underlies the workings of the cosmos!

★ Keplers laws remain accurate to this day—indeed, in slightly generalized form will show up in many (most!) situations where motions are controlled by gravity

★ Yet note what we still don't have:

an understanding of *why* Kepler's laws hold

→ that is, what is the *mechanism* that makes

⇐ planets move this way

...for that, need to wait for Kepler's successors...

# Galileo Galilei

First to use telescope in Astronomy

www: Galileo shows scope to Duke

contributions:

- mountains on the moon
- moons of Jupiter
- sunspots
- phases of Venus

www: Venus phase animation

observations contradicted Aristotle

supported Copernicus

“paradigm shift” (Kuhn)

12 radical change in outlook/conceptual framework

Galileo brilliant but also arrogant and politically naive  
→ offended powerful people, including the Pope, a former ally  
tried in Inquisition and forced to recant geocentric view

- his work, Copernicus, Kepler banned until 1832
- official semi-apology ( “mistakes were made” ) 1992(!)

complex situation: crackdown as much political as theological

Note:

1. really not at all obvious to people that Earth orbits Sun
2. the paradigm shift was difficult and threatening

With earth removed from center of universe,  
Aristotle's division of terrestrial and heavenly  
no longer made sense as physics

- ⇒ need to re-examine “natural motion”
- ⇒ search for force that keeps planets in place
- ⇒ Galileo's **experiments**

# Dynamics & Gravity

Galileo not only great astronomer  
but also a great physicist  
paved way for Newton's dynamics by study of  
two special cases of motion

1. **“free body”** – *no* external influences  
natural motion: coast in straight line, const speed  
→ retain current state of motion  
→ bodies have **inertia**

2. **“free fall”** – when only influence is *gravity*  
Galileo recognized another key motion

*Demo:* Tower of Pisa expt

→ constant acceleration **indep of mass!**

$$a = g, g = 9.8 \text{ m/s}^2$$

Galilean free fall: constant acceleration  $a = g$

So speeds change linearly with time

$v = v_0 + gt$ ; if  $v_0 = 0$ ,  $v = gt$

Distance traveled is quadratic in time:

$$d = \int_0^{t_f} dt v(t) = \int_0^{t_f} dt gt = \frac{1}{2}gt_f^2 \quad (3)$$

**Ex** how long does it take to drop from table to floor?

$$d \sim 1\text{m} \Rightarrow t^2 = 2d/g = 2 \times 1\text{m}/9.8\text{m/s}^2 \sim 0.2\text{s}^2 \Rightarrow \boxed{t \sim 0.45\text{ s}}$$