Astro 350 Lecture 14 Sept. 26, 2011

Announcements:

- HW4 available, due Friday
- Discussion Question 4 due *Wednesday*
- Hour Exam 1 grading underway

### **Breaking News! Report of faster-than-light motion!**

result of a neutrino experiment Q: what's a neutrino? how are they made?

- $\nu$ s produced in Switzerland www: CERN
- travel about about 700 km  $\approx$  400 miles *under* the Alps *Q: what?! how?*
- measured in Italy www: Gran Sasso

You can understand the calculation!

- ▶ precisely measure distance *d* travelled *Q*: how?
- $\triangleright$  precisely measure time t travelled Q: how?
- $\triangleright$  then solve  $v_{\nu} = d/t$

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 $\Rightarrow$  reported result:  $v_{\nu} > c$  !  $\nu$ s faster than light!?! www: actual research paper reporting result

Q: why this this a very Big Deal?

# Neutrinos

neutrinos (symbol  $\nu$ ) are subatomic particles

- electric charge  $Q(\nu) = 0 \rightarrow$  neutral (hence the name)
- mass: we know  $m(\nu) > 0$ , and we know  $m(\nu) \ll m_{\rm electron}$  but as yet we do not know neutrino masses
- only made in *some* nuclear reactions ('Weak Interactions'') for example: in Sun, first step in reaction chain for helium production is:  $p + p \rightarrow d + e^+ + \nu$  $\Rightarrow$  the Sun is a huge source of neutrinos!
- very weakly interacting most neutrinos born in Sun's core pass entirely through Sun to effectively stop neutrinos from the Sun, would need a column of lead about 1 light-year long!

 We will see: neutrinos are among the most interesting and possibly the most important particles in the Universe

## iClicker Poll: Faster-Than-Light Neutrinos

Vote your conscience! Lay your bets! What is your gut reaction to this result?

- A Woo hoo! Sounds very likely, and I never liked speed limits anyway!
- B Meh. Sounds iffy, but more likely correct than false
- С
- Meh. Sounds possible, but more likely false than true
- D

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No way! Sounds like it has gotta be wrong!

New result, nobody (including instructor!) knows where it will lead

Instructor's vote: *Gotta be wrong!* Why? Because I am deeply convinced of relativity Will try to give an idea why starting Friday

To make things interesting: *two wagers* 

- by this Friday, new papers will appear with new proposed theories to explain results if this is correct: candy for everyone!
- before the last ASTR350 class, new paper(s) will appear explaining the result as an error or a misinterpretation if this is correct: *pizza for everyone!*
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# Our Milky Way Galaxy: Where are we?

recall MW structure: disk shape note similarity with planar concentration of planets in SS

where are we in the disk-near middle or edge?
www: MW mosaic

on MW band in sky, stars  $\approx$  evenly distributed *Q: simplest interpretation?* www: Herschel model (1700's) *Q: loophole in the argument?*  clue: dark strips in MW dust: absorbs light  $\rightarrow$  only see small part of MW disk this fact only verified in 20th century

But then: How to determine MW structure and size?

H. Shapley (1910's): **globular clusters** of stars most lie **out** of disk plane  $\rightarrow$  we have unobscured view

*Q:* how does sky pattern of GC's tell where we are?

If we are at MW center:

 $\rightarrow$  see GC's evenly spread around the sky

If we are off-center:

 $\rightarrow$  see GC's more on one side of sky

 $\rightarrow$  that's Galactic "downtown"

www: observed GC sky distribution
\* we are not at the Milky Way center!

high-tech update:

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dust obscures *visible* light, but not longer wavelengths dust "invisible" if  $\lambda \gg$  dust size so infrared, radio telescopes *can* see all of MW will see: these confirm we are off-center

### **Revolution Revisited**

Cosmologist Y. Berra: It's déjà vu all over again!

#### Copernican Revolution I (17th Century):

we're one typical planet among many not center of solar system

#### Copernican Revolution II (earth 20th Century):

we're one typical star among many not center of Milky Way Galaxy

... stay tuned for more...

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## **Observed Milky Way Structure**

Milky Way contains about  $10^{11} = 100$  billion stars

**I. Disk Components**: most of luminous matter radius  $R \sim 15,000 \text{ pc} = 15 \text{ kpc}$  (kpc = kiloparsec = 1000 pc) thickness  $h \sim 200 \text{ pc}$  at our location: thin! www: IRAS full sky: dust. False color, Galactic coords www: DIRBE near-IR image: cool stars note-confirms our suburban location! 1. disk contains most stars 2. also dust, gas  $\rightarrow$  fuel for star formation

Disk Structure

- disk thickest in center, tapers off outward
- disk shows evidence for spiral arms

 $\rightarrow$  we are spiral galaxy! (as in www: M104)

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#### **II. Spherical Components**

- 1. bulge at center (old stars, can see in DIRBE image)
- 2. globular clusters
- 3. "halo" of old stars

#### Milky Way Dynamics

in MW, all objects exert gravity on all others

- $\rightarrow$  everything accelerating
- $\rightarrow$  everything is in motion

# Milky Way Rotation

measure speeds of stars, gas via Doppler effect complication: we are moving too

stars orbit MW center disk stars: ~ circular orbit  $\rightarrow$  disk rotates but disk stars *don't* spin like frisbee (i.e., a solid object) *Demo*: frisbee: rigid rotation in time Sun goes around once stars closer to center go around more than once stars further out — less than once  $\rightarrow$  "differential rotation"

how measure rotation speeds? use halo stars, globular clusters (don't rotate)

Sun orbit speed:  $v_{\odot} = 220 \text{ km/s}$ at our location  $R_{\odot} \simeq 8 \text{ kpc}$  (about halfway out!)