

Astro 350
Lecture 29
Nov. 2, 2011

Announcements:

- Discussion Question 8 due today
- HW 8 due Friday

Office hours: Instructor – today after class

TA: tomorrow 2-3 pm

Hints for Q2: in *static* U, $a(t) = \text{constant} = 1$ always!
and $\rho = \text{constant}$ always and everywhere

Last time:

- matter-only universe: $a(t) \propto t^{2/3}$

Q: what is the final fate of such a universe?

how do you tell?

matter-only prediction: cosmic age $t_0 = \frac{2}{3} \frac{1}{H_0} = 9.2$ billion yr

typo fixed from last time

but oldest stars have age > 12 billion years

Q: and so?

fate of Universe: what happens in distant future
mathematically, far future: $t \rightarrow \infty$

matter-only: $a \propto t^{2/3}$

so as $t \rightarrow \infty$, then $a \rightarrow \infty$

distances between objects grow infinitely large!
universe expands forever!

matter-only prediction: $t = 9.2$ billion years

actual data: $t > 12$ billion years

\Rightarrow contradiction!

We do not live in the “what you see is what you get”

ω matter-only universe!

also last time:

cosmic expansion history and cosmic acceleration

key ideas:

- at any time t , everywhere $v(t) = H(t) r(t)$
- when looking at galaxies far, far away
seeing light emitted a long time ago
- so if measure v (from redshift) and r (from std candle)
then can find $H(t) = v(t)/r(t) =$ expansion at past time $t!$
- even better: can find $H(t)$ for many times t *Q: how?*

Bottom line:

- ‡
- ★ can measure *change* in expansion rate
 - ⇒ can measure cosmic deceleration/acceleration

Acceleration vs Deceleration

consider to objects today, say two galaxies
presently at some distance r_0 , say 100 Mpc

right now moving apart at some speed given by $v_0 = H_0 r_0$
due to cosmic expansion

imagine times the past, when closer together

*Q: if “turn off” gravity (really, remove all matter),
what are speeds at earlier times?*

*Q: now turn on (ordinary, attractive) gravity: what are speeds
at earlier times?*

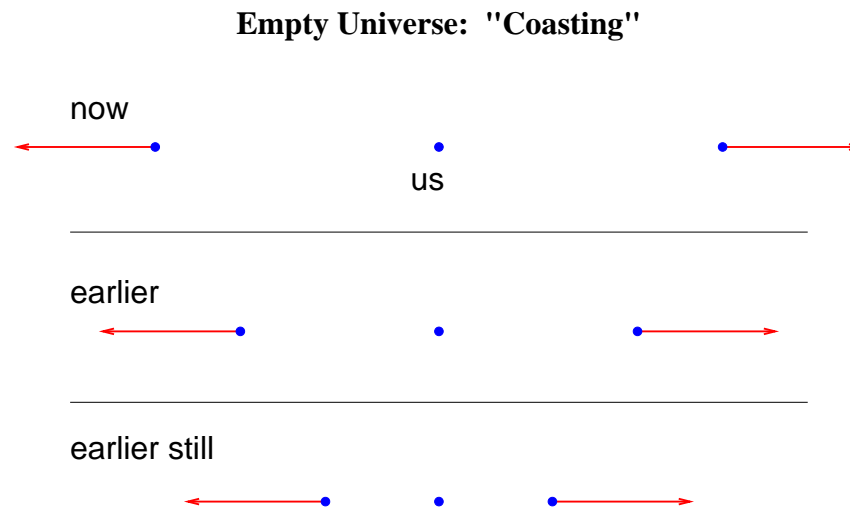
*Q: but now follow supernova results—what are speeds at earlier
times?*

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Q: what could explain these results?

Matter-Free “Empty” Universe

- No matter → no attraction between galaxies
- nothing to change galaxy speeds
- galaxies “coast” keeping constant velocity
- ⇒ same speeds in past



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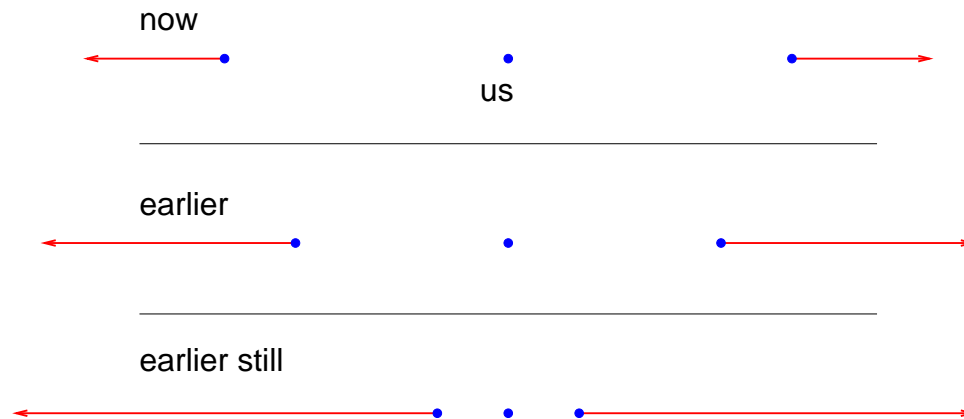
expansion rate:

neither accelerated **nor** decelerated

Ordinary Gravity and Matter

- The real universe has galaxies with mass → attract each other
- inward gravity slows expansion
- speeds constantly *decreasing*, galaxies *decelerating*
- ⇒ to achieve observed speed today, had to be *faster* in past!

Normal Gravity and Matter: Decelerating Universe

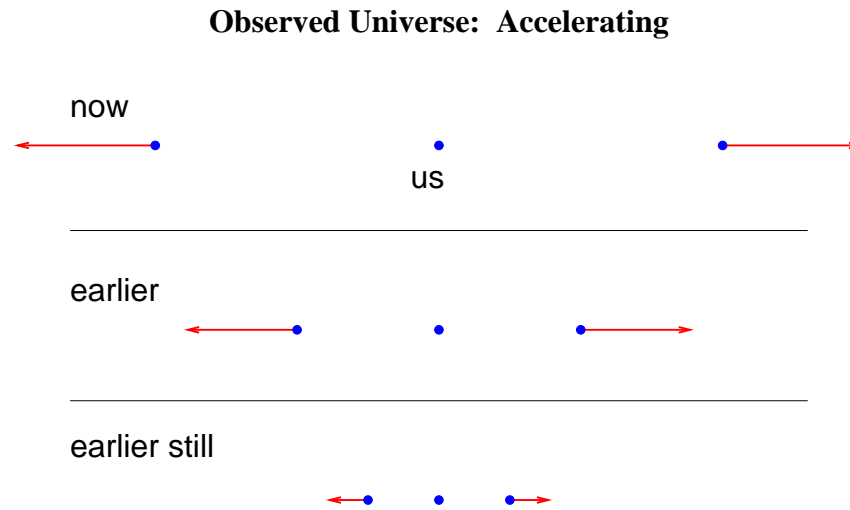


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expansion rate: **decelerated**

Observed Universe

Our actual observed universe:
galaxies *slower* in past!



∞ expansion rate: **accelerated!**

Cosmic Expansion History: the Full Story

in fact, observations show that cosmic expansion had *two* phases

★ *today and in the recent past*

expansion is **accelerating**

opposite of prediction from matter + General Relativity

★ *in the more distant past*

at redshifts $z > 0.3$, times before $t < 10$ billion years

that is, more than about 3 billion years ago

expansion was **decelerating**

◦ agrees with prediction from matter + General Relativity!

Cosmic Acceleration: Who Ordered That?!

ordinary matter and ordinary gravity: **attractive**
gravity acts to draw galaxies together,
⇒ slows outward expansion
should give cosmic **deceleration**

but we observe cosmic **acceleration!**
two known options:

1. Universe contains something bizzare that pushes objects apart!
in fact, this repulsion has to be so strong that it overcomes gravity attraction from matter!

2. *Q: what's the other option?*

Einstein Overthrown?

cosmic acceleration seen when looking at the Universe
over vast distances

result is surprising because our gravity theory
= Einstein's General Relativity
predicts gravity makes matter attractive

but note: while GR very well tested on Earth
and in Solar System
not tested on cosmic lengthscales

perhaps acceleration tells us:
General Relativity is incomplete/wrong

→ if so, need new gravity theory!

Q: requirements for such a theory

Improving on Einstein?

if new gravity theory:
still have to explain *all* data

so: any new theory has to

- give same answers as GR
on Earth, solar system scales
- and keep other successful GR features:
redshifting, lensing, time dilation
- yet also give different *answers on cosmic scales*

iClicker Poll: The Reason for Cosmic Acceleration

Vote your conscience!

Of these two basic explanations for cosmic acceleration

Which do you think is right?

- A** General Relativity *correct*, but the Universe contains something bizarre that makes it accelerate
- B** General Relativity *incorrect*, and the Universe only contains matter

Explaining Cosmic Acceleration

Cosmologists are working hard on *both* avenues

- a new cosmic “accelerant”: “dark energy”
- alternatives to General Relativity: “modified gravity”

Simplest solution: Einstein “cosmological constant” Λ

originally invented by Big Al (1917):
“fudge factor” in General Relativity

changes gravity force law:

mass m at distance r from mass M feels force

$$F_{\text{gravity}} = -\frac{GMm}{r^2} + \frac{1}{3}\Lambda mr \quad (1)$$

Q: what if $\Lambda = 0$ and $M = 0$? what does this mean?

Q: what do we get if $\Lambda = 0$ but $M > 0$? why the $-$ sign?

what happens to particle released from rest?

Q: what if $M = 0$ and $\Lambda > 0$?

what happens to particle released from rest?

Q: Λ invented to prevent cosmic expansion—how?

In cosmological context, Λ changes acceleration to

$$\frac{\ddot{a}}{a} = -\frac{4\pi}{3}G\rho + \frac{1}{3}\Lambda \quad (2)$$

can “fine-tune” Λ to prevent expansion/collapse by exactly compensating for normal gravity and make the universe not accelerate or decelerate

Q: what fine tuning needed?

Einstein did this

but: after Hubble sees expansion in 1929,

Al allegedly sez Λ his “greatest blunder”

Q: what if Λ exist but not fine tuned?

However: if Λ *not* fine-tuned
can “overcompensate” for gravity attraction
→ lead to acceleration!
so: in wake of SNIa results, Λ rebirth!

→ even AI’s blunders turn into gold!

in detail: Λ alters Friedmann:

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi}{3}G\rho - \frac{K}{a^2} + \frac{\Lambda c^2}{3} = \frac{8\pi}{3}G\rho_{\text{tot}} - \frac{K}{a^2} \quad (3)$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi}{3}G\left(\rho + 3\frac{P}{c^2}\right) - \frac{\Lambda c^2}{3} = -\frac{4\pi}{3}G\left(\rho_{\text{tot}} + 3\frac{P_{\text{tot}}}{c^2}\right) \quad (4)$$

where ρ_{tot} includes $\rho_{\Lambda} = \Lambda c^2/8\pi G$ “vacuum energy density”

P_{tot} includes $P_{\Lambda} = -\rho_{\Lambda}c^2$

In other words:

- $\Lambda \neq 0$ gives energy content to empty space!
(“vacuum energy”)
 - and vacuum energy has negative pressure!
- \Rightarrow fills the bill for acceleration!

Λ and its Discontents

Recall that expansion accelerated only “recently”
seems we live in a special time – right when acceleration begins
Why Λ only important right now?

Try another approach: look for more general solutions
acceleration does requires new strange cosmic ingredient
but not required to be cosmo constant

Acceleration \rightarrow U Contains Negative Pressure Substance

but: can prove that matter of any kind

and radiation of any kind must have $P \geq 0$

\Rightarrow only known substance that has $P < 0$ is

new kinds of energy fields (“scalar fields”)

generically—negative pressure substances: **dark energy**

- Λ simplest, does not evolve with time or z
- generally, dark energy *does* evolve:
can be larger or smaller in past

In particular: some dark energy models (“quintessence”)
give DE evolution that always keeps

ρ_{DE} close to ρ_{matter}

so not unusual that the two are close today+

Warning!

some dark energy models have ρ_{DE} increasing with time!

i.e., expansion \rightarrow *larger* density of dark energy!?!

leads to scale factor growth

$$a_{\text{future}}(t) \propto \frac{1}{t_{\text{rip}} - t} \quad (5)$$

20 where $t_{\text{rip}} > t_0$ is a future time

Q: *what happens when $t = t_{\text{rip}}$? why is this bad?*