

Astro 350
Lecture 37
Nov. 28, 2011

Announcements:

- HW11 due Friday
- Discussion Question 11 – please give your vote
- Check syllabus: lowest HW and Discussion score dropped but you are still responsible for all of the material
- **ICES** available online – please do it!

I do read and use comments!

Before break: primordial nucleosynthesis

- Theory: nuclear reactions in the early universe at times $t = 1$ sec to 3 minutes

make lightest elements:

76% hydrogen, 24% ^4He , traces of D, ^3He , ^7Li
amounts each depend on cosmic density of

baryons = anything made of atoms

and nothing heavier! all other elements made by stars

- Observation: measure light elements in real universe
not easy to do: have to find samples
with minimum “pollution” by element production in stars

Result

Combine observations (+ errors!)

observed light elements agree with theory (and each other!)

if baryon density $\Omega_B = \rho_B/\rho_{\text{crit}}$ in range

$$0.040 \lesssim \Omega_B \lesssim 0.050 \quad (1)$$

recap: extrapolated big bang to $t = 1$ s, predicted lite elts

kinda amazing: not only qualitative agreement (“lotsa helium”)

but even detailed quantitative agreement with observations!

Cosmo bragging rights: BBN is earliest probe!

BBN: Implications

Qualitatively

extrapolated big bang to $t = 1$ s

predicted light elements \rightarrow agreement with observations

big bang working well back to 1 sec!

Quantitatively

observed light elements measure cosmic baryons

i.e., total amount of matter in form of atoms

$$\Rightarrow 0.040 \lesssim \Omega_B \lesssim 0.050$$

1. $\Omega_B \ll 1$: baryons don't close the U.

‡ 2. $\Omega_{lum} \sim 0.007 \ll \Omega_B$

baryonic dark matter hot (10^{6-7} K) intergalactic gas?

3. $\Omega_{\text{matter}} \approx 0.3 \gg \Omega_{\text{B}}$:

non-baryonic dark matter

confirms: **most dark matter** is **not**
made of atoms of any kind in any arrangement!
→ must be exotic form of matter!

known matter = anything on the periodic table
is a tiny fraction of the makeup of the cosmos!

Early Universe Cosmology Scorecard

Recall strategy:

- inventory universe today
- **extrapolate** back to early epochs
- apply known laws of nature to expanding, cooling U
- identify observable consequences (“fossils”) persisting today
- measure fossils → learn about early U!

Cosmo Report Card

Epoch	Recombination	Big Bang Nuke
cosmic time t	$\sim 400,000$ yr	~ 1 sec–3 min
micro-processes	nuclei + $e \rightarrow$ atoms	$p + n \rightarrow$ nuclei
predicted fossils	thermal radiation	baryons \rightarrow H, He, Li
observed?	Yes! \rightarrow CMB	Yes! \rightarrow primordial abundances
grade	A	A

iClicker Poll: ICES

Vote your conscience!

Did you get one or more ICES notification emails for this course?

- A** Yes, and I promise to fill out the form if I have not already
- B** No, but I promise to go to <https://ices.cte.uiuc.edu/> and fill out the form if I have not already

The Very Early Universe

CMB success \Rightarrow understand Univ at $t \sim 400,000$ yr
 $z \sim 1100$ and $T \sim 1$ MeV

BBN success \Rightarrow understand Universe at $t \sim 1$ s
 $z \sim 10^{10}$ and $T \sim 1$ MeV

success gives confidence:
boldly extrapolate to $t \ll 1$ s
and $T \gg 1$ MeV

Q: what are conditions like?

∞ *Q: what physics needed to describe?*

A Brief History of Time

The Very Early Universe & Ultra-High-Energy Physics

Planck Epoch: $t \lesssim 10^{-43}$ s

general relativity invalid – quantum effects large

⇒ need quantum GR theory: quantum gravity

i.e., we don't know what happens at 10^{-43} s

which means the one thing we can be sure of is that

we aren't yet “qualified” to go back earlier

to the big bang itself $t = 0$ sec!

→ the nature of the big bang itself intimately tied

to the unification of gravity and quantum mechanics

the ultimate inner space/outer space connection!

Of course, there are ideas:

- maybe universe described by string theory?
- maybe spacetime infected w/ quantum fuzziness (?)
- quantum black holes created and evaporated (?)
- or maybe the U is a **braneworld**...

Extra Dimensions on the Brane

Braneworld Scenario for Quantum Cosmology

proposes that our (expanding) 3-dimensional space is just a “surface” / “membrane” in a much larger 4-dimensional(!) “bulk” space!

with particles (i.e., us!) confined to brane, but gravity extending into the “bulk”

one suggestion: our 3-D “brane” has another “parallel” brane very nearby (in a 4th dimension!)

with side-effect: gravity from matter in sibling brane appears to us as DM (and we are DM for them!)

braneworld ideas currently being in the lab!

would show up as departure from $F_{\text{grav}} \propto r^{-2}$

so far: inverse square holds down to ~ 1 mm

→ “sibling” brane has to be at least this close!

Particle Physics Today: Success and Its Discontents

Current theory of elementary particles:
“the Standard Model of Particle Physics”

all known particles explained in terms of

- matter particles in “families” of quarks and “leptons” (e , ν and cousins)
- interacting with four fundamental forces: gravity, electromagnetism, and the nuke and weak forces
- with forces “carried” by another set of particles i.e., photons and cousins

How does this stack up against experiment?

- extremely (annoyingly!) successful

⇒ no known disagreement with experiment!

e.g., e^- magnetic moment ($g - 2$) measurement agrees with theory to 1 part in 10^{10} !

- But: Standard Model only tested in lab to Fermilab energies $E \sim 1 \text{ TeV} = 10^{12} \text{ eV}$
- And: Standard Model begs the questions:
 - why the patterns of particles we see?
 - why four forces are they unified (like E&M are)?
 - where does mass come from?
 - why is matter one class of particles (fermions) and force carriers another (bosons)?

Standard Model a “victim of its own success”
carries the seeds of its destruction/supplanting

To address these questions: new particle theories
give possible answers to these questions
as a by-product, forced to invent new particles:

- almost always high-mass ($m \gtrsim 1 \text{ TeV} = 1000 m_{\text{proton}}$)
- almost always weakly interacting
(at “low” energies = Fermilab/CERN)
- note: invented to fix particle problems,
not with cosmology in mind (no ulterior motive!)

Today: new particles hard to make

But in early U: created everywhere!

13 Q: *possible fossils today? what conditions needed?*

The Heavenly Accelerator and Dark Matter

If exotic massive particles exist

→ created in early universe

If stable: remain today

→ natural candidates for **dark matter**

bonus: naturally weakly interacting

“just what the doctor ordered”

Weakly Interacting Massive Particles: WIMPs!

key point: not invented for cosmology

but for particle physics reasons

So: if particle theorists are right:

can't *avoid* a U filled with crazy WIMPs

assume they are the DM:

Q: how detect them in the lab?

Direct Detection of WIMPs

Difficult! ...but not impossible

weakly interacting \rightarrow experiments similar to ν detection

- go underground
- expect small count rate (\lesssim few events/month)

www: WIMP experiments

WIMP-nucleus collisions: nucleus recoils with ~ 1 keV
measure recoil energy: cryogenic detectors

strategy: look for annual variations

$$\vec{v}_{\text{WIMP}} = \vec{v}_{\odot} + \vec{v}_{\text{Earth,orbit}}$$

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\rightarrow velocity has time change due to earth orbit

\rightarrow modulation in 1-year period, amplitude $v_{\text{Earth}} \sim 10\%v_{\odot}$

WIMP Search Results

1998: Italian experiment (DAMA) claims evidence!

by now: claim evidence is strong

- very controversial result!
- most competing groups don't see signal
- could be different WIMP interactions for different nuclei
- ...or could be false alarm

How to resolve dispute? Better experiments

- will be coming online
- either will find WIMPs, or rule out favorite theories
- stay tuned!