

Astro 350  
Lecture 4  
Aug 28, 2011

Announcements:

- Discussion Question 1 due Wednesday nite
- HW1 due at start of class Friday  
turn in paper copy, but can & should upload on Compass  
online submission gives record if question of HW loss
- **register** your iClicker by next time!  
follow link on course page

Last time:

- geocentric cosmology  
*Q: which is what?*  
*Q: why would anyone believe this?*  
*Q: who was Ptolemy? What is the Ptolemaic system?*
- Copernicus & heliocentric cosmology  
*Q: what is an AU?*

## Copernicus: What's New and What's Not

- planets still on spheres
- Copernicus still used epicycles!
- predictions *not* better than in Ptolemy's model  
→ geometrically equivalent Q: *meaning?*
- Copernicus' model not generally accepted  
and Ptolemaic–Copernican disagreement though to be metaphysical, *unanswerable* question

Q: *so how do we decide which is right?*

## Tycho Brahe 1546-1601: Danish Astronomy Extraordinaire

in youth: observed “nova stella” (supernova) www: Tycho sketch  
→ change observed in heavens → corruptible!  
observed Sun, Moon, planets for 20 years: careful, accurate data  
but not a good number cruncher  
→ like any good professor: made grad student do the work!

## Johannes Kepler 1571–1630: Harmony of the Worlds

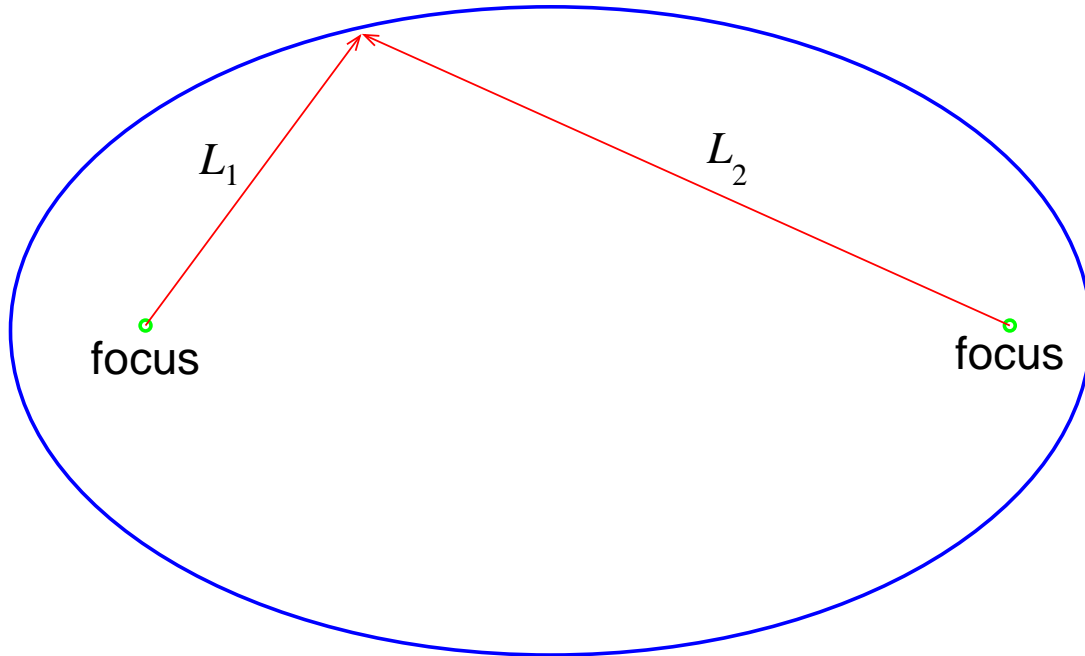
Analyzed Tycho’s data for **20 years**(!), especially Mars motions  
used heliocentric model with circles  
but observations didn’t quite agree  
a small error (few arc min!) remained...took seriously  
→ after trial & error:

ω completely & accurately described planet orbits

*Q: Kepler’s Laws?*

# Kepler I: Law of Ellipses

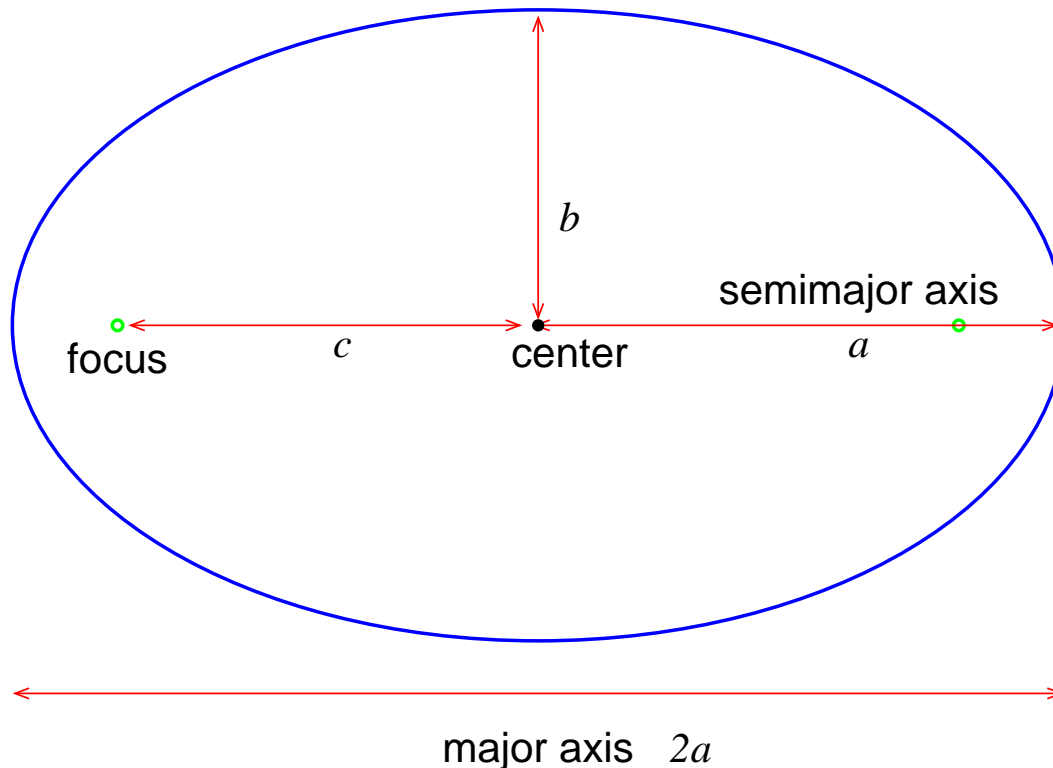
each planet's orbit is an **ellipse**  
with the **sun at one focus**



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$$L_1 + L_2 = \text{constant}$$

# Ellipse Anatomy



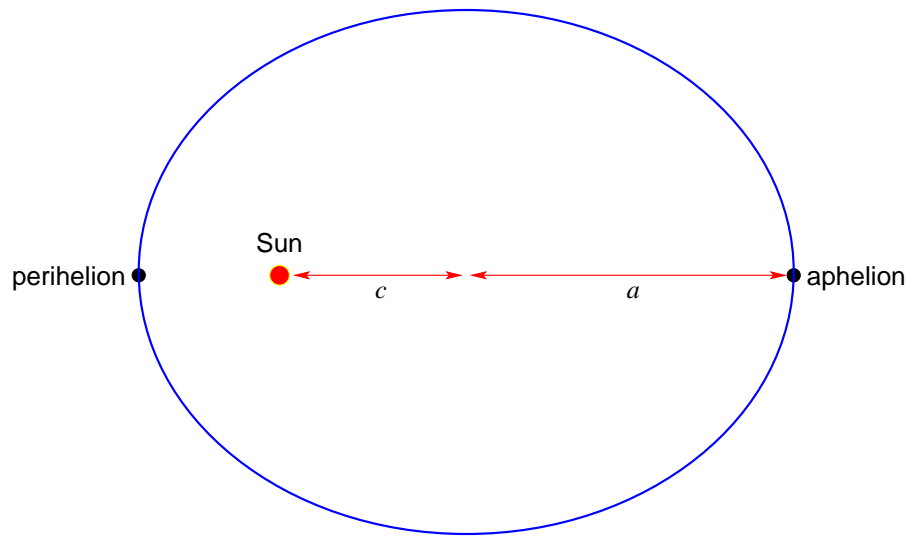
- two foci
- semi-major axis  $a$
- focal length  $c$
- semi-minor axis  
 $b = \sqrt{a^2 - c^2}$

any ellipse fully characterized by:

$a$  and eccentricity  $e = c/a$

Q: what do we get for  $e = 0$ ?  $e = 1$ ?

Kepler I: orbit is **ellipse** with **sun at one focus**



Orbit anatomy

*aphelion*: farthest point from Sun

*perihelion*: closest point to Sun

*Q: what is aphelion distance in terms of  $a$  and  $e$ ?*

$$r_{\text{ap}} = a + c = a + a\frac{c}{a} = (1 + e)a \quad (1)$$

o *Q: If the Sun's at one focus, what's in the other focus?*

*Q: What does Kepler I not say about orbits?*

At the other focus: nothing! (sorry!)

Note: Kepler I only gives orbit *shape*  
but says *nothing* about how orbit evolves in time  
→ need more info to fully describe orbit, hence...

## Kepler II: Law of Equal Areas

a straight line from the planet to the sun  
sweeps out equal areas in equal times

*diagram: sketch areas*

note that this amounts to telling about speed of planet

*Q: where fastest? slowest?*

www: area animation

*Q: This still doesn't fully characterize an orbits—why not?*



Kepler I gives orbit shape in space

Kepler II gives orbit evolution over time

but haven't yet connected the two:

how does spatial character (e.g., semimajor axis  $a$ )

relate to time character (e.g., period  $P$ )?

Need one last law...

## Kepler III: The Mighty Equation

period  $P$  and semi-major axis  $a$  are related:

$$P^2 \propto a^3$$

$\Rightarrow P^2/a^3 = \text{const}$ , holds for all planets, with same constant and since must hold for Earth:

$$P_{\text{yrs}}^2 = a_{\text{AU}}^3 \quad (2)$$

*Q: ok for earth?*

where  $P$  written in years,  $a$  in AU

Very powerful! e.g.:

Asteroids exist with orbits inside 1 AU (and some cross 1 AU!!)

www: inner solar system objects--in real time!

## iClicker Poll: Kepler III

Kepler III:  $P_{\text{yrs}}^2 = a_{\text{AU}}^3$

Consider an asteroid with an orbit entirely outside 1 AU

Is its period longer or shorter than a year?

- A**  $P > 1$  yr, no matter eccentricity  $e$
- B**  $P < 1$  yr, no matter what  $e$
- C** can't answer without knowing  $e$

## Kudos to Kepler

Several points worth noting...

★ An amazing discovery—mathematics underlies the workings of the cosmos!

★ Orbits have a simple geometry  
...but *not* simplest: ellipse not circle

★ Kepler's laws remain (almost) perfectly accurate to this day—indeed, in slightly generalized form will show up in many (most!) situations where motions are controlled by gravity

★ Yet note what we still don't have:  
an understanding of *why* Kepler's laws hold  
→ that is, what is the *mechanism* that makes  
planets move this way  
...for that, need to wait for Kepler's successors...

## Galileo Galilei: Astronomer

First to use telescope in Astronomy

www: Galileo shows scope to Duke

contributions:

- mountains on the moon
- moons of Jupiter
- sunspots

These are bad for Ptolemy (but maybe not deadly) Q: *how?*