

Astro 350
Lecture 7
Sept. 7, 2011

Announcements:

- HW2 due next time
- Discussion Question 2 due today

Please take and test-drive a diffraction grating slide

Last time:

Newton's laws and Newton's gravity

→ combined to give a complete theory of solar system motions

↳ (first) Copernican revolution

Cosmology with Newton

Great progress, many aspects quite modern:

- solar system motions understood exquisitely accurately
- masses of Earth, Sun, planets known
- stars understood to be other suns at great distances
scale/size of cosmos huge and \gg than Ptolemy imagined!

but still: at the time “Universe” = few thousand stars):
needed more powerful telescopes, and
better understanding of local “building blocks” – stars
 \Rightarrow need technology to understand star emission (light)
and star interiors (atoms)

Cosmologist's Toolbox: Outer Space

Most of what we will learn about cosmology comes from studying object beyond earth

Q: take a minute, meet your neighbor, and make a list:

From earth,

What can we **directly** measure about planets/stars/galaxies?

What can we **directly** measure about planets/stars/galaxies?

Hard-nosed list:

- position on sky
- color/spectrum
- intensity/brightness
- time changes in the above (if any)

Other information (all the really interesting stuff!) is *indirect*!

Lesson:

⇒ can only measure light!

⇒ need to understand light and its interactions with matter
to decode info

Last time:

- ⊢ ● light = electromagnetic wave *Q: what does this mean?*
- characteristic properties

Light and Atoms: the Quantum Microworld

Light

Light deeply connected to electric charge, electric and magnetic forces

Experiments show:

changing E force generates M force

changing M force generates E force

1. E&M linked: “electromagnetic force”

2. EM disturbances can travel through space:

each regenerates the other: $E \rightarrow M \rightarrow E \rightarrow M \rightarrow \dots$

electromagnetic waves = “EM radiation”

light is a type of **wave**:

a wave is oscillating disturbance in a medium

wave can travel, medium does not

Demo: the wave!

snapshot in space, taken at one instant of time:

diagram

wavelength λ size of one cycle

→ wave “ID number”

intensity I

“strength” of wave = “height of peaks”

Demo: slinky: same wavelength, diff’t intensities

∨

Q: Sound waves: how do we experience λ ? I ?

Q: Light waves: how do we experience λ ? I ?

Sound:

$\lambda \leftrightarrow$ pitch

high pitch (treble): small λ

low pitch (bass): large λ

intensity = loudness

Light:

$\lambda \leftrightarrow$ color

visible light: larger λ : more red

smaller λ : more blue intensity = brightness

The Speed of Light

Light is very fast!

So fast that it was a feat to measure the speed in lab
now known quite well

$$\begin{aligned}c &= \text{constant} = 299,792,458 \text{ m/s} \\ &= 3.0 \times 10^8 \text{ m/s} = 186,000 \text{ miles/s} = 6.7 \times 10^8 \text{ miles/hr}\end{aligned}$$

enormous—but **not infinite!**

→ finite speed of light hugely important for astronomy

→ telescopes are time machines *Q: how?*

note: light speed c is *same* for all λ

◦ *Q: what would happen if this were not true?*

The Electromagnetic Spectrum

EM waves can have λ outside of visible range

www: EM spectrum

IN general, light is combination of pure waves with different λ

distribution of intensities:

different brightness at diff. λ : **spectrum**

diagram: sketch spectrum axes

Q: spectrum of laser pointer?

Q: spectrum of white light?

Technology Tim's Terminology Tip: "Radiation"

Warning!

meaning of "radiation" in Physics, Astronomy, Cosmology
≠ "radiation" in everyday parlance!

In Physics, Astronomy, Cosmology...and more importantly...
In this course and on the exams:

radiation = movement of energy through space
carried by particles or waves

Examples:

ordinary visible light! e.g., flashlight, sunlight, starlight, ...
completely benign and indeed necessary for life!

but also invisible EM waves: radio, UV, X-ray...

and even non-EM particles: neutrinos...

⚠ Beware Confusion: "radiation" so defined ≠ radioactivity!
more on radioactivity later...

Thermal Radiation

hot objects **glow** (think stove burner)
temperature – radiation connection

microscopic picture:
temperature → atom motion

but atoms made of charged particles
motion → changing EM forces → light

thermal body (at some temp T)
emits EM radiation: which λ emitted?

iClicker Poll: Is Red-Hot Really Red?

Lightbulb on low power: wire appears to glow red
but look at glow through prism (diffraction grating)

What colors will you see?

- A red only
- B red and orange
- C red, orange, and yellow
- D red, orange, yellow, green
- E red, orange, yellow, green, blue

spectrum of light; depends on T

perfect absorber of light: “blackbody”

absorbed energy \rightarrow heats up

re-emits according to T

“blackbody radiation” = thermal radiation

Thermal Spectrum: Light as Thermometer!

peak λ is color seen: $\lambda_{\text{peak}} \propto 1/T$

where T is *absolute* temperature in **Kelvin**

recall: $T_{\text{Kelvin}} = 273 + T_{\text{Celsius}}$; room temp ≈ 290 K

hotter \rightarrow more blue \rightarrow shorter λ

“Wien’s Law”

Turn the equation around: $T \propto 1/\lambda_{\text{peak}}$, namely

$$T = 3000 \text{ K} \left(\frac{10^{-6} \text{ m}}{\lambda_{\text{peak}}} \right) \quad (1)$$

so: can find T just from light!

\Rightarrow spectrum as **thermometer**

color measures temperature