Astro 350 Lecture 13 Sept. 26, 2012

Announcements:

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- Hour Exam 1 Friday; info online discussion at end of class today
   Office hours: me today at 1pm; TA Thurs 9:30-10:30 am
- *Discussion* 4 due today
- *Discussion 5* posted today, due next Wednesday
- Bonus Participation: class portrait on Compass identify yourself to help me learn your name

Last time: began stars & cosmology Q: the Sun is made of a huge mass of gas, which can flow and be compressed, yet the Sun maintains a spherical shape and constant size—why?

# The Lives of Stars

The life of a star is a struggle against its own gravity

- if gravity force balanced by pressure, star is stable and to keep pressurized, must stay hot!
- if pressure weaker then gravity, star unstable collapses under its own weight

#### **Birth**

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stars formed when cold gas clouds collapse due to gravity compression  $\rightarrow$  heating, until T at center  $\rightarrow 10^7$  K "birth" when first nuke reactions begin

#### Youth and Midlife (Main Sequence) – All Stars

in core of star, nuclear reactions convert  ${\rm H}$   $\rightarrow$   ${\rm He}$ 

- energy release  $\rightarrow$  heat  $\rightarrow$  maintains outward pressure
  - $\rightarrow$  balances inward gravity  $\rightarrow$  stability! ("hydrostatic equilibrium")

## Hydrogen Burning in Stars

interstellar gas is mostly (about 75%) hydrogen stars formed from this gas  $\rightarrow$  stars begin as mostly H

nuclear "burning" of hydrogen to helium:

- key reactions occur in "chains"
- first step involves pre-existing solar ingredients
- input for each new step is output from previous step

Dominant reactions: "'pp" Chain  

$$p + p \rightarrow \frac{2H}{e^{+}} + \frac{e^{+}}{\nu} + \frac{\nu}{e^{-}} + e^{+} \rightarrow \gamma + \gamma$$
  
 $^{2}H + p \rightarrow ^{3}He + \gamma$   
 $^{3}He + ^{3}He \rightarrow ^{4}He + 2p$ 

<sup>ω</sup> Net effect:  

$$4p + 2e^- \rightarrow 2n2p = {}^{4}\text{He} + \text{energy} + ...$$

each "p-p reaction" creates:

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• <sup>2</sup>H=<u>np</u> "deuterium"
"heavy hydrogen" nucleus
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## • *e*<sup>+</sup> "positron"

antimatter: positively charged anti-electron! more later about antimatter then  $e^- + e^+ \rightarrow \gamma + \gamma$  energy! annihilation

#### • *ν* "neutrino"

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very low-mass  $(m_{\nu} \ll m_e)$  particle only created in nuclear reactions ("weak" decays) very weakly interacting particle once born, go thru Sun, Earth, your body but almost never interact  $\nu$  escape diagram

## The Nuclear Powered Sky

Before 1930's:

- a mystery how the Sun could burn for billions of years
- no known energy source would work

In the 1930's: nuclei, nuclear reactions, nuclear energy discovered it was realized that this can power the Sun and all stars www: Nobel Prize: Hans Bethe

The Sun is a mass of incandescent gas a gigantic nuclear furnace Where hydrogen is burned into helium, at temperatures of millions of degrees

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<sup>-</sup> Lou Singer and Hy Zaret, 1959; cover: They Might Be Giants 1993

## **Inner Space and Outer Space**

Lesson: a deeper understanding of "inner space" i.e., the microscopic world led to a deeper understanding of "outer space" i.e., the astronomical/cosmological world

Q: how could we be so sure?

Can we get even more direct confirmation? *Q: is another way to confirms the Sun is a nuclear reactor?* A "smoking gun" signature?

# **The Evidence: Solar Neutrinos**

If the Sun takes  $4p \rightarrow {}^{4}\text{He} = 2p2n$ then it *must* convert  $2p \rightarrow 2n$  $\rightarrow$  *must* produce neutrinos! in fact: most made via  $pp \rightarrow de^{+}\nu$ 

The Sun radiates neutrinos as well as photons!

Moreover:

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- since ν are weakly interacting they come directly from the solar core
   → messengers from the center of the Sun!
- but luckily, weakly interacting  $\neq$  non-interacting  $\Rightarrow$  solar neutrinos are potentially observable!
- clever experiments can try to "catch" them

# **In Search of Solar Neutrinos**

experiments have been built to "see" solar neutrinos by observing rare cases of  $\nu$  interactions with atoms all use huge underground detectors *Q*: why huge? why underground?

Two types:

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1. "radiochemical" – vats of fluid
see element change due to \nu
ex: chlorine fluid \nu + {}^{37}\text{Cl} \rightarrow {}^{37}\text{Ar} + e^-
collect Ar atoms (radioactive!)
www: Davis chlorine experiment
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2. "scattering" - vats of ultrapure water
see light pulses from
high-energy e^- scattered by \nus
www: SNO ball
www: Super-K Sun image
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Upshot:

- **\star** All experiments detect solar  $\nu$ s!
- $\star$  Amount (flux) is just as predicted
- *Q*: what fundamental fact(s) is/are confirmed?

# **Solar Neutrino Results**

I. proof that Sun powered by nuke fusion II.  $\nu$ s give view into solar core III. these are  $\nu$  telescopes!

A new window on the Universe: **Nobel Prize 2002!** 

Poetry reading: John Updike, "Cosmic Gall"

Imagine: 100 years ago, you try to explain that the Sun and all stars create tiny invisible particles that pass through us all the time in huge numbers and are essential byproducts of the working of stars

Q: a lesson for cosmology?

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## Hour Exam 1

www: Exam Info on Course Website

www: Front Page: Instructions and Equations

Any Questions?

# **Sample Questions**

#### **Multiple Choice**

An object orbiting the Sun with a semimajor axis of 9 AU has what orbital period?

- (a) 3 years
- (b) 9 years
- (c) 27 years
- (d) depends on orbit eccentricity
- (e) none of the above

Star A and star B are observed to have the same flux, but the parallax of star A is **larger** than the parallax of star B. Thus the **luminosity** of star A is \_\_\_\_\_ the luminosity of star B.

- (a) larger than
- (b) smaller than
- (c) the same as
- (d) not enough information to answer
- (e) none of the above

#### Short Answer

Sketch the rotation curve observed for our Galaxy. Be sure to:

- label both axes
- indicate the center of the Galaxy
- indicate the visible edge of the Galaxy

Sketch the rotation curve our Galaxy **would** have if all of the matter were in the form of stars. Be sure to label both axes and identify the Galactic center and visible edge.

Explain why the observed rotation curve of our Galaxy gives evidence for dark matter.

Give an example of possible form of dark matter and explain why it meets the requirements to be dark matter.