

Astro 406
Lecture 12
Sept. 23, 2013

Announcements:

- **PS 4 due Friday**
- iClicker scores posted on Compass; check for accuracy!
- ASTR 401: outline due today

Last time: internal Galactic motion—orbital in the disk

we want to construct a *rotation curve* $v(R)$ vs R

Q: challenges in measuring circular speed $v(R)$ pattern?

Q: what Galaxy component(s) provides useful velocity tracers?

└ *Q: what observables can we directly measure?*

Galactic Rotation Observed: Gas Components

Doppler requires a line to be shifted

but nature has been kind: strong lines available:

- 21 cm from atomic H
- rotational lines from CO molecules

bonus: nature has been *really* kind: both are in radio

→ no dust extinction → can see whole Galaxy!

observables:

- Doppler shift → line-of-sight (radial) velocity v_{los}
- measure in plane as a function of Galactic longitude ℓ

www: 21 cm (ℓ, v_{los}) plot

www: CO (ℓ, v_{los}) plot

2

Q: $v \neq 0$ for all ℓ : implications?

Q: what regions populated? unpopulated?

Galactic Rotation Overview

$v \neq 0$ for all ℓ :

- *our Galaxy does not rotate as a solid body*
- different Galactocentric radii have different v, Ω
 \Rightarrow *differential rotation*

Signal only populates some regions:

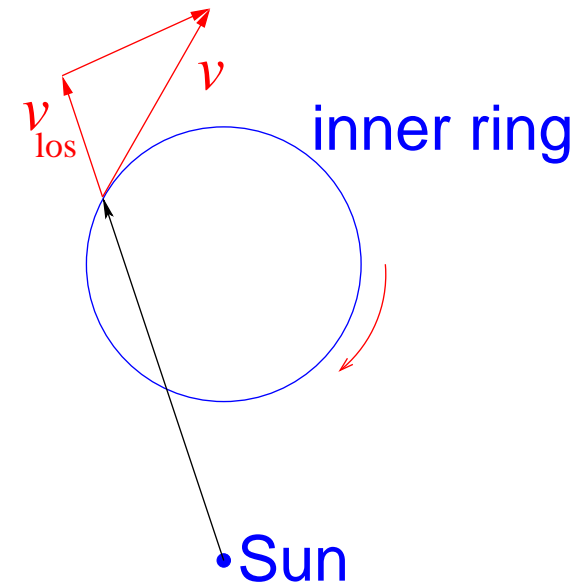
- for $\ell \in (90^\circ, 180^\circ)$, $v_{\text{los}} < 0$ only
and $v_{\text{los}} > 0$ only for $(270^\circ, 360^\circ)$
- in other regions, both v_{los} signs found

Q: expectations for one inner ring?

One Inner Ring

features:

- v_{los} changes sign across Galactic center at thus is zero towards center at $\ell = 0$
- v_{los} maximum on sightline tangent to ring i.e., when ℓ_{max} satisfies $R = R_0 \sin \ell_{\text{max}}$ and here $v_{\text{los}} = v(R)$
- for $|\ell| < \ell_{\text{max}}$: signal from 2 points
Q: are v_{los} signs the same?
- no gas found at $|\ell| > \ell_{\text{max}}$
→ no signal at these longitudes



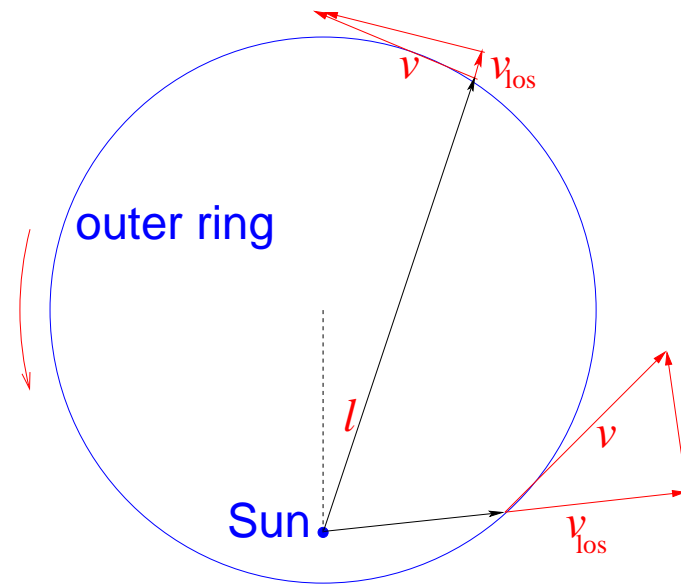
observations: for $0 < \ell < 90^\circ$, strong $v_{\text{los}} > 0$

- *inner rings have larger Ω than us*
- *Galactic rotation carries us in this direction*

One Outer Ring

Now imagine a single *outer* ring
i.e., with $R > R_0$
and assume $\Omega(R) < \Omega(R_0)$

Q: v_{los} sign for $\ell \in (0, 90^\circ)$?
($90^\circ, 180^\circ$)?
($180^\circ, 270^\circ$)?
($270^\circ, 360^\circ$)?



An *exterior ring* at $R > R_0$ with $\Omega(R) < \Omega(R_0)$:

- $v_{\text{los}} < 0$ for $\ell \in (0, 180^\circ)$
- $v_{\text{los}} > 0$ for $\ell \in (180^\circ, 360^\circ)$

Recall *interior ring* result for $\Omega(R) > \Omega(R_0)$:

- $v_{\text{los}} > 0$ for $\ell \in (0, \ell_{\text{tan}})$
- $v_{\text{los}} < 0$ for $\ell \in (360^\circ - \ell_{\text{tan}}, 360^\circ)$

Note: if $\Omega(R)$ *decreases* with R

- only some regions populated
- matching the observed pattern!

observational confirmation that $\Omega(R)$ decreases with R

Q: *implications for orbit periods?*

◦ Q: *what does Galactic motion look like?*

Q: *what about orbit velocities?*

orbit periods: $P(R) = 2\pi/\Omega(R)$

→ $P(R)$ increases

⇒ interior orbits “laps” us, but
we “lap” exterior orbits!

orbit velocity: $v(R) = \Omega(R) R$

- product of decreasing and increasing functions of R
- need to examine in detail to see what “wins”

Milky Way Rotation Curve: Summary

Measure $v(R)$ vs R , and plot: **rotation curve**

www: 21 cm velocity profile: SG fig 2.18

Trends:

- difficult to measure accurately for $R \lesssim 2$ kpc
some indication of roughly linear increase $v(R) \propto R$
- beyond this, $v(R) \rightarrow v_0 \approx 220$ km/s = const
→ **“flat rotation curve”**

But recall: for circular orbits

$$m(r) = \frac{v^2 r}{G} \quad (1)$$

∞ since $v(r) = v_0$, $m(R) \propto R!$

Q: where have you seen this before (hit: PS2)?

Q: what does this mean for Galactic density?

the linear increase of $m(r) \propto r$
is identical to result from isothermal sphere!
...about which you are experts (PS2!)
e.g., density, for $r \gtrsim R_0$

$$\rho(r) \propto \frac{1}{r^2}$$

Q: other implications of $m(R) \propto R$?

Milky Way Rotation Curve: Implications

Flat rotation curve gives

$$m_{\text{enclosed}}(R) \propto R \quad (2)$$

trend continues for largest measured R
in far outer regions with little/no stars and gas

Mass grows even when *no luminous matter present!*

$$m(50 \text{ kpc}) \gg M_{\star} + M_{\text{gas}}$$

we have a big discrepancy!

Q: what are logically possible ways out?

Flat Rotation Curves: New Gravity Physics?

Flat rotation curve + gravitational dynamics

$$\rightarrow M_{\text{tot}} \gg M_{\text{lum}}$$

What explains this mismatch?

Approach 1: reject assumptions, avoid conclusions

Note that result rests on Newtonian gravity

applied to Galactic scales

but we have never tested gravity on these scales

What if there is a fundamental problem with our present understanding of gravity?

- then mass mismatch really probes new gravity behavior
- this *is* an active area of research!

11

Q: arguments against this approach?

Q: alternatives if this approach fails, i.e., gravity theory is right?

Flat Rotation Curves: Dark Matter

New gravity theory, like any new theory must explain *all* existing data – not just the new anomaly

- Newton/Einstein explain terrestrial, solar system, stellar data with great precision
- new theory must do all of this too
must “look like” standard gravity on small scales
but change only on large scales $\gtrsim 1$ kpc
- very difficult to do – but not ruled out!

Approach 2: accept assumptions, live with conclusions

Alternative possibility: we *do* understand gravity

and thus must live with $M_{\text{tot}} \gg M_{\text{lum}}$

→ most ($\sim 80 - 90\%$) of MW is **dark matter**

Our Dark Halo

How far does the dark matter extend?

flatness of rotation curve well-measured out to $R = 20$ kpc

Magellanic clouds at ~ 50 kpc also fall on trend

→ dark matter extends at least this far

new component Milky Way structure: **dark halo**

dominates mass of Galaxy

visible stars and gas embedded at center of halo

Q: what can we say about the nature of dark matter?

Dark Matter

What we know is all in the name!

- gravitates → has mass (“matter”)
- non-luminous: does not emit (visible) light (“dark”)
Q: but can say more about how it interacts with light
- also: cannot strongly absorb light *Q: why?* (no/weak EM interactions)
- probably: no/weak strong (nuclear) interactions *Q: why?*
that’s all we know for sure!

What is the DM?

Q: you tell me...

Dark Matter Candidates

black holes
neutron stars
white dwarfs
Jupiters, brown dwarfs

} compact objects

hot $\sim 10^6\text{K}$ gas

cold gas

neutrinos

relic exotic particles from earliest moments of big bang

Q: how do we figure out which (if any) are right?

iClicker Poll: Dark Matter Gut Feeling Twofer

Vote your conscience!

1. Which dark matter candidates seem **most plausible** to you?

1. Which dark matter candidates seem **least plausible** to you?

A black holes

B other compact objects:
planets, brown dwarfs, white dwarfs, neutron stars

C very hot gas or very cold gas

D neutrinos

E relic exotic particles

The Search for Dark Matter

Does dark matter really exist in the Milky Way?
If so, what form does it take?

⇒ difficult problem! (“highly non-trivial”)
have to work hard and be clever

Next week: we will cross off 5 lines from our list!
The rest: still open possibilities

Director's Cut Extras

Galactic Rotation and Doppler Surveys: Derivation

Basic idea: we can only measure the

- line-of-sight component of the
- velocity relative to us

Us:

- Galactocentric radius \vec{R}_0
- angular velocity $\vec{\Omega}_0$
- linear velocity $\vec{v}_0 = \vec{v}(R_0) = \vec{\Omega}_0 \times \vec{R}_0$

Them (i.e., some arbitrary point in disk):

- Galactocentric radius \vec{R}
- angular velocity $\vec{\Omega}$
- linear velocity $\vec{v} = \vec{v}(R) = \vec{\Omega} \times \vec{R}$

¹⁹ relative distance (i.e., distance from them to us):

$$\vec{d} = \vec{R}_0 - \vec{R}$$

Here Goes...

diagram: $\vec{R}_0, \vec{R}, \vec{d}, \ell$

relative velocity (them vs us):

$$\Delta \vec{v} = \vec{v}(\vec{R}) - \vec{v}(\vec{R}_0) = \vec{\Omega} \times \vec{R} - \vec{\Omega}_0 \times \vec{R}_0 \quad (3)$$

but $\vec{R} = \vec{R}_0 - \vec{d}$

$$\Delta \vec{v} = (\vec{\Omega} - \vec{\Omega}_0) \times \vec{R}_0 - \vec{\Omega} \times \vec{d} \quad (4)$$

and along **line-of-sight** \vec{d} :

$$v_{\text{los}} = \hat{d} \cdot \Delta \vec{v} = \frac{\vec{d}}{d} \cdot \Delta \vec{v} = \frac{\vec{d}}{d} \cdot (\vec{\Omega} - \vec{\Omega}_0) \times \vec{R}_0 - \frac{\vec{d}}{d} \cdot \vec{\Omega} \times \vec{d} \quad (5)$$

vector ID: $\vec{a} \cdot (\vec{a} \times \vec{b}) = 0$

$\vec{a} \cdot (\vec{b} \times \vec{c}) = (\vec{c} \times \vec{a}) \cdot \vec{b}$

$$v_{\text{los}} = \frac{\vec{d} \times \vec{R}_0}{d} \cdot (\vec{\Omega} - \vec{\Omega}_0) \quad (6)$$

and since $\vec{d} \times \vec{R}_0 = -dR_0 \sin \ell \hat{z}$, at last we get

Ta da! The Magic Formula

$$v_{\text{los}} = [\Omega(R) - \Omega(R_0)] R_0 \sin \ell = \left(\frac{v}{R} - \frac{v_0}{R_0} \right) R_0 \sin \ell \quad (7)$$

Q: what if $R = R_0$? $R < R_0$? $R > R_0$?

Prediction: if $\Omega(R)$ decreases w/ R

Q: v_{los} signs?

diagram: top view, plot v_{los} vs ℓ