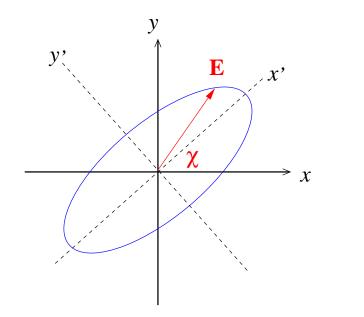
Astro 501: Radiative Processes Lecture 13 Sept. 26, 2018

Announcements:

• **Problem Set 4** due Friday note: blackbody fit to real data will not be perfect!

Last time: monochromatic plane waves *Q: what's that? properties? Q: polarization specail cases? general case?*



Ν

 $E'_x = \mathcal{E}_0 \cos\beta \cos(\omega t)$ $E'_y = -\mathcal{E}_0 \sin\beta \sin(\omega t)$ principle axes: $\mathcal{E}_0 \cos\beta$ and $\mathcal{E}_0 \sin\beta$

if $\beta \in [0, \pi/2]$: ellipse sweeps clockwise \rightarrow "righthanded" elliptical polarization, negative helicity if $\beta \in [-\pi/2, 0]$: "lefthanded", positive helicity

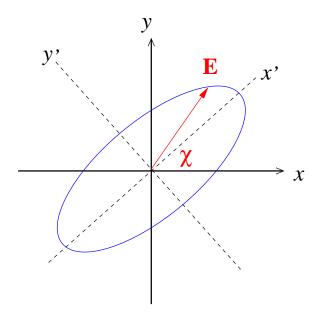
Q: what $\beta(s)$ give complete linear polarization? circular?

we want to relate x - y field parameters $\mathcal{E}_1, \mathcal{E}_2, \phi_1, \phi_2$ to x' - y' principle axes parameters $\mathcal{E}_0, \beta, \chi$

rotate x - y components by angle χ

ω

 $E_x = \mathcal{E}_0 \left(\cos \beta \cos \chi \cos \omega t + \sin \beta \sin \chi \sin \omega t \right)$ $E_y = \mathcal{E}_0 \left(\cos \beta \sin \chi \cos \omega t - \sin \beta \cos \chi \sin \omega t \right)$ matching to, e.g., $E_x = \mathcal{E}_1 \cos(\omega t - \phi_1)$:



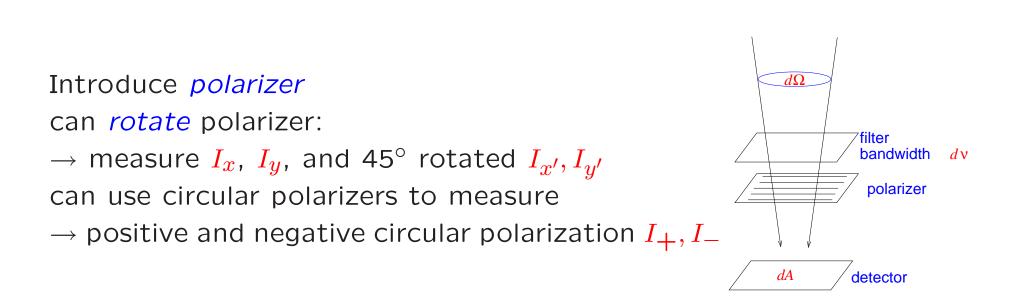
$$\mathcal{E}_1 \cos \phi_1 = \mathcal{E}_0 \cos \beta \cos \chi \tag{1}$$

$$\mathcal{E}_1 \sin \phi_1 = \mathcal{E}_0 \sin \beta \sin \chi \tag{2}$$

$$\mathcal{E}_2 \cos \phi_2 = \mathcal{E}_0 \cos \beta \sin \chi \tag{3}$$

$$\mathcal{E}_2 \sin \phi_2 = -\mathcal{E}_0 \sin \beta \cos \chi \tag{4}$$

Q: how can we determine polarization by intensity measurements with a polarimeters?



combine: Stokes parameters

$$I = I_x + I_y \tag{5}$$

$$Q = I_x - I_y \tag{6}$$

$$U = I_{x'} - I_{y'}$$
 (7)

$$V = I_{+} - I_{-}$$
 (8)

Q: what physically is each? can more than one of Q,U,V be nonzero? what does that correspond to? Q: range of values for Q? U? V? are they all independent?

Stokes Parameters

for *monochromatic waves*, Stokes parameters related to $\mathcal{E}_1, \mathcal{E}_2, \phi_1, \phi_2$ and $\mathcal{E}_0, \beta, \chi$ bases:

$$I = \mathcal{E}_{1}^{2} + \mathcal{E}_{2}^{2} = \mathcal{E}_{0}^{2}$$
(9)

$$Q = \mathcal{E}_1^2 - \mathcal{E}_2^2 = \mathcal{E}_0^2 \cos 2\beta \cos 2\chi \tag{10}$$

$$U = 2\mathcal{E}_1 \mathcal{E}_2 \cos(\phi_1 - \phi_2) = \mathcal{E}_0^2 \cos 2\beta \sin 2\chi \qquad (11)$$

$$V = 2\mathcal{E}_1 \mathcal{E}_2 \sin(\phi_1 - \phi_2) = \mathcal{E}_0^2 \sin 2\beta$$
 (12)

and thus

сл

$$\mathcal{E}_0 = \sqrt{I} \tag{13}$$

$$\sin 2\beta = V/I \tag{14}$$

$$\tan 2\chi = U/Q \tag{15}$$

since wave has 3 independent parameters, Stokes parameters must be *related*

$$I^2 = Q^2 + U^2 + V^2 \tag{16}$$

Quasi-Monochromatic Waves

natural light generally not a pure monochromatic wave with a single, definite, complete state of polarization

rather: a *superposition* of components with many polarizations

consider wave with *slowly varying* amplitudes and phases

 $E_1(t) = \mathcal{E}_1(t) \ e^{i\phi_1(t)}$; $E_2(t) = \mathcal{E}_2(t) \ e^{i\phi_2(t)}$ (17)

"slow": wave looks completely polarized on timescalse ω^{-1} but amplitudes and phases drift over intervals $\Delta t \gg \omega^{-1}$ \rightarrow polarization changes

but also wave is *no longer monochromatic* frequency spread: "bandwidth" $\Delta \omega \sim 1/\Delta t \ll \omega$ \rightarrow quasi-monochromatic wave

0

Q: effect on Stokes?

Stokes Parameters for Quasi-Monochromatic Light

real measurements represent averages over timescales during which polarization can change

Stokes parameters become averages

$$I = \langle E_1 E_1^* \rangle + \langle E_2 E_2^* \rangle = \left\langle \mathcal{E}_1^2 + \mathcal{E}_2^2 \right\rangle$$
(18)

$$Q = \langle E_1 E_1^* \rangle - \langle E_2 E_2^* \rangle = \left\langle \mathcal{E}_1^2 - \mathcal{E}_2^2 \right\rangle$$
(19)

$$U = \langle E_1 E_2^* \rangle + \langle E_2 E_1^* \rangle = 2 \langle \mathcal{E}_1 \mathcal{E}_2 \cos(\phi_1 - \phi_2) \rangle$$
 (20)

$$V = -i\left(\langle E_1 E_2^* \rangle - \langle E_2 E_1^* \rangle\right) = 2\left\langle \mathcal{E}_1 \mathcal{E}_2 \sin(\phi_1 - \phi_2) \right\rangle \quad (21)$$

but for quasi-monochromatic waves

$$I^2 \ge Q^2 + U^2 + V^2 \tag{22}$$

- quasi-monochromatic polarization is still in general *elliptical*
- but drifts can reduce degree of polarization

$$I^2 \ge Q^2 + U^2 + V^2 \tag{23}$$

- maximum polarization when equality holds: *completely elliptically polarized*
- minimum when Q = U = V = 0: unpolarized
- arbitrary wave is *partially polarized*

useful to define *polarized* intensity

$$I_{\rm pol} = Q^2 + U^2 + V^2 \tag{24}$$

and since $I_{pol} \leq I$, define fractional degree of polarization

$$\Pi \equiv \frac{I_{\text{pol}}}{I} = \frac{\sqrt{Q^2 + U^2 + V^2}}{I}$$
(25)

 \odot

note: can always decompose Stokes parameters

$$\begin{pmatrix} I \\ Q \\ U \\ V \end{pmatrix} = \begin{pmatrix} I - I_{\text{pol}} \\ 0 \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} I_{\text{pol}} \\ Q \\ U \\ V \end{pmatrix}$$
(26)

sum of unpolarized and polarized components

Superposition and Stokes

consider composite wave that is superposition of many independent waves

electric field components are given by superposition

$$E_1 = \sum_k E_1^{(k)}$$
; $E_2 = \sum_k E_1^{(k)}$ (27)

each term k of which has different phase

PS4: phases specified, can calculate sum explicitly

but generally, phases are random

so field products average out phases from different waves

$$\left\langle E_i E_j^* \right\rangle = \sum_k \sum_\ell \left\langle E_i^{(k)} E_j^{(\ell)*} \right\rangle = \sum_k \left\langle E_i^{(k)} E_i^{(k)*} \right\rangle \tag{28}$$

but due to this averaging, *Stokes parameters are additive*

$$I = \sum_{k} I^{(k)} \tag{29}$$

$$Q = \sum_{k} Q^{(k)} \tag{30}$$

$$U = \sum_{k}^{n} U^{(k)} \tag{31}$$

$$V = \sum_{k} V^{(k)} \tag{32}$$

How Do Charges Generate Radiation

Thus far: vacuum Maxwell solutions support EM waves

- \bullet speed c
- transverse
- $\vec{B} = \vec{n} \times \vec{E}$

Maxwell **sources** are charges and currents

But how do sources *generate* radiation? Strategy: study point charge, then superpose

Consider a point charge *at rest* Q: what are ρ , \vec{j} everywhere? \vec{E} , \vec{B} everywhere?

A Point Charge at Rest

Consider a point charge q at rest at origin $\vec{r} = 0$ charge density $\rho = q\delta(\vec{r})$ current density $\vec{j} = \rho \vec{v} = 0$

Gauss' Law: $\nabla \cdot \vec{E} = 4\pi \rho$ Spherical symmetry: $\vec{E} = E(r) \hat{r}$ Gauss' Theorem applied to sphere enclosing charge:

$$\int \nabla \cdot \vec{E} \, dV = \int \vec{E} \cdot d\vec{A} \int E \, dA = 4\pi r^2 E \tag{33}$$

$$= 4\pi \int \rho \ dV = 4\pi \ q \tag{34}$$

$$E(r) = \frac{q}{r^2} \tag{35}$$

Coulomb's Law!

and $\vec{j} = 0$ means $\vec{B} = 0$: no magnetic field

Q: how can things change if the charge moves?

An Accelerated Point Charge

consider a particle rapidly *decelerated* from speed v to rest over time δt



consider a later time $t \gg \delta t$

- Q: field configuration near particle ($r \ll ct$) ?
- Q: field configuration near particle ($r \gg ct$)?
- Q: consequences?

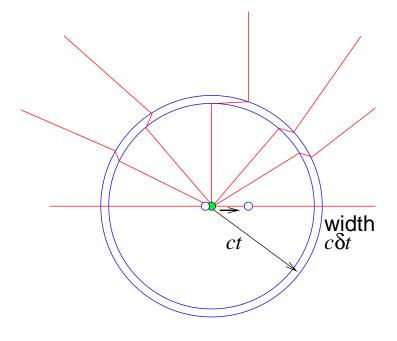
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for fields track particle location expected for constant velocity

- nearby: $r \ll ct$, fields radial around particle at rest
- far away: $r \gg ct$: fields don't "know" particle has stopped \rightarrow "anticipate" location displaced by ct from original particle radially oriented around this expected point

between the two regimes: $r = ct \pm c\delta t$ field lines must have "kinks" which

- have tangential field component
- tangential component is *anisotropic* and largest $\perp \vec{v}$



consider vertical fieldline $\perp \vec{v}$: kink radial width $c\delta t$ kink tangential width vt = (v/c)r

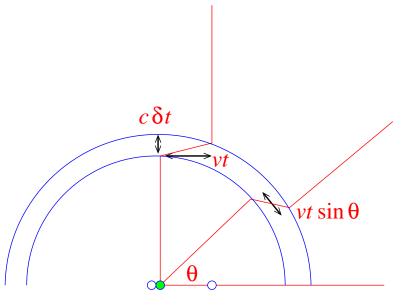
tangential/radial ratio is $(v/\delta t)r/c^2$ but $v/\delta t = a$, average acceleration: $\rightarrow E_{\perp}/E_r = ar/c^2$

more generally, tangential width is $vt \sin \Theta = (v/c)r \sin \Theta$ and so using Coulomb for E_r :

$$E_{\perp} = \frac{ar\sin\Theta}{c^2} E_r = \frac{qa}{c^2r}\sin\Theta$$
(36)

this is huge! *Q: why?*

G: relation to radiated flux?



We find acceleration leads to a propagating field perturbation that is **tangential = transverse!** just what we expect for EM radiation

so we expect also a transverse \vec{B} component, with

$$B_{\perp} = E_{\perp} = \frac{ar\sin\Theta}{c^2} E_r = \frac{qa}{c^2r}\sin\Theta$$
(37)

and thus a radial Poynting vector with magnitude

$$S = \frac{c}{4\pi} E_{\perp}^2 = \frac{q^2 a^2}{4\pi c^3 r^2} \sin^2 \Theta$$
(38)

this is also huge! Q: why?

Q: total radiated power per solid angle?

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Larmor Formula

Poynting flux:

$$S = \frac{c}{4\pi} E_{\perp}^2 = \frac{q^2 a^2}{4\pi c^3 r^2} \sin^2 \Theta$$
 (39)

- scales as $S \propto 1/r^2!$ as it must!
- note importance of $E_{\perp} \propto 1/r$ scaling

Total power into solid angle $d\Omega$: $dP = r^2 S d\Omega$ so power per solid angle

$$\frac{dP}{d\Omega} = r^2 S = \frac{cr^2 E_{\perp}^2}{4\pi} = \frac{q^2 a^2}{4\pi c^3} \sin^2 \Theta$$
(40)

Larmor Formula for Radiated power $\overrightarrow{\omega}$ *Q: lessons from magnitude? direction?* Larmor:

$$\frac{dP}{d\Omega} = r^2 S = \frac{cr^2 E_{\perp}^2}{4\pi} = \frac{q^2 a^2}{4\pi c^3} \sin^2 \Theta$$
(41)

- magnitude $P \propto a^2$: accelerated charges radiate
- direction: $dP/d\Omega \propto \sin^2 \Theta$ not isotropic! maximum orthogonal to accleration zero along acceleration