

Astro 210
Lecture 10
Feb 9, 2011

Announcements

- HW3 due Friday
HW1 Q8 bonus still available
- Office hours: Instructor 2-3pm Wed
TA 10:30-11:30am Thurs
- Planetarium: new shows added
registration, report form, info online

Last time: Kepler from Newton

solve $\vec{F} = m\vec{a} = m\ddot{\vec{r}}$ with $\vec{F} = -GMm/r^2 \hat{r}$

gives back Kepler's laws, and so

- agrees precisely with observed planet orbits
- also explains how orbits arise from gravity
- and gives, e.g., circular speed: $v_c = \sqrt{\frac{GM}{r}}$
- and updates Kepler III: $a^3 = \left(\frac{GM}{4\pi^2}\right) P^2$

Newtonian gravity: possible orbits

line, circle, ellipse, parabola, hyperbola

Gravity and energy

- bound vs unbound orbits

Q: which is which? what is condition for bound vs unbound?

- escape speed $v_{\text{esc}} = \sqrt{2GM/r}$

Q: why is this important?

Escape Speed

At radius r , define

escape speed

$$v_{\text{esc}} = \sqrt{2GM/r}$$

- if launch from r with $v_{\text{launch}} < v_{\text{esc}}$
then $TE < 0$: *fall back!* (elliptical orbit)
- if launch from r with $v_{\text{launch}} > v_{\text{esc}}$
then $TE > 0$: *escape "easily"*: $v > 0$ at $r = \infty$
- if launch from r with $v_{\text{launch}} = v_{\text{esc}}$ exactly
then $TE = 0$ exactly, *"just barely" escape*

So: escape speed is *minimum speed* needed to leave a gravitating source

Example: escape speed from earth

$$v_{\text{esc}} = 11 \text{ km/s} = 25,000 \text{ mph!}$$

ω

predict the future: if toss object with $v < 25,000$ mph, falls back
but if $v > 25,000$ mph *Q: example?* never returns!

finally, the more “generic” unbound orbit:

hyperbola

$$r(\theta) = \frac{(e^2 - 1)a}{1 + e \cos \theta} \quad (1)$$

$$e > 1, TE > 0$$

$v > 0$ at $r = \infty$: nonzero speed far from M

Recall: at large r , hyperbola \rightarrow *straight line*

But Newton says: $d\vec{v}/dt = -\frac{GM}{r^2} \hat{r}$

so as $r \rightarrow \infty$, then $d\vec{v}/dt \rightarrow 0$

\Rightarrow gravity negligible, $\vec{v} \rightarrow$ const: *free body*=straight line!

orbit of unbound “flyby”:

- ∇ distant nearly free body \rightarrow passing: pulled toward M
 - \rightarrow distant deflected nearly free body

Two-Body Problem

Thus far: cheated! (i.e., simplified)

- (1) neither Sun nor planets “nailed down”, and
- (2) Newton III → planets exert net force on Sun
- ⇒ (3) Sun moves too! (but larger M , so less accel.)

How to analyze 2-body system?

imagine a box, with mass M , with no net forces on it (floating in space).

Q: how would it move?

Q: what if the box has pieces in it—still same answer?

box of mass M : without forces, moves inertially
i.e., as free body \rightarrow constant \vec{v}

now open box: contains two pieces, mass m_1 and m_2
no matter what pieces do, box still has constant \vec{v}

now imagine moving with same velocity as box:
so to you, box is at rest

if at one time, know where the two particles are (draw)

then later if particle 1 has moved (draw new position) then:

Q: can you say anything about where particle 2 has to be?

diagram: draw m_1 , m_2 , then dotted box around, and $M = m_1 + m_2$

define center of mass (COM):

$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2} \quad (2)$$

draw \vec{R} on diagram

a mass-weighted average

Can show: $\ddot{\vec{R}} = 0 \rightarrow \dot{\vec{R}} = \text{const.}$

\rightarrow can pick inertial frame where $\dot{\vec{R}} = 0$, choose $\vec{R} = 0$: origin of coordinates

Planet & Sun as a Two-Body System

in Sun-planet system: center of mass is “fixed” (free body)
but Newton III says that since Sun pulls on planet
then planet pulls back on Sun

→ *both* accelerate

→ *both* orbit around center of mass

Q: but what's the difference in the motions?

Q: who's more correct: Copernicus or Ptolemy?

Testing Newton's Gravity

Moons of Jupiter: obey Kepler's laws

→ Jupiter's gravity works like Sun's, Earth's

1830's: Uranus observed orbit did *not* follow predictions of Newtonian solar system model

⇒ the death Newton's gravity?

recall: theory must explain *all* data, not just some!
so despite Newton's great job with planets, moons even *one clear failure is enough*

Q: *so do we have to throw out Newtonian gravity?*

◦ Q: *why hesitant to throw out?*

Q: *if not abandon, what's another solution to the problem?*

iClicker Poll: Uranus Discrepancy

1830's Problem: *measured* Uranus orbit *doesn't match* predictions of Newtonian Gravity *theory*

Vote your conscience!

Which seems more likely to you?

- A** Newton's gravity theory *correct*, but not all gravity sources had been included
- B** Newton's gravity theory *incorrect* (or at least incomplete)

Q: *what experiment/observation would tell which is right?*

maybe haven't included all sources of gravity?
maybe unknown/unseen object causes U's deviations?
⇒ a new planet?

if unknown object, could predict where should be
did this, looked. saw:

www: Neptune

1846: Neptune found at right position

▷ *predicted* by Newton's gravity ("dark matter")

very impressive! victory snatched from jaws of defeat!
triumph of Newtonian dynamics and gravity

many other confirming observations

www: binary star orbits

Heliocentric vs Geocentric Finale

What is the main lesson, for the practice of science, of the geocentric vs heliocentric shift?

Note:

not asked *content* of science (don't say lesson=heliocentric) but rather the *practice*—what does it tell us here and now about how to do science?

Geocentric vs Helocentric: Lessons

For me, a big lesson is **Humility!**

naive to think: “Greeks dumb, we’re smart”

rather a sobering reminder: sometimes, same observations can be explained in radically different ways

also: can have bias not even aware of
shapes how view world, seems reasonable to everyone
humbling! examples in QM, relativity

what’s more...probably going on still today!

remember: all astronomy, all science ultimately tentative

In this course: my guess: $\sim 80\%$ stand test of time

but don’t know which 20% is wrong...so have to learn it all!

that said, not everything up for grabs or matter of taste...

confidence/uncertainty varies tremendously

My Wagers