Astro 210 Lecture 39 April 29, 2011

Announcements

- HW 11 due now last one! no HW next week!
- Final Exam: Monday May 9, 1:30–4:30 pm info online
- ICES course evaluation available online *please* fill it out–I *do* read & use results

Last time: our home galaxy-the Milky Way

- Q: basic shape? where are we??
- Q: rotation curve—what is it? what does it look like? What does it mean?

Dark Matter: Results to Date

Dark matter identity is still *unknown* \Rightarrow one of the top questions in all of science!

But we do know what dark matter is not

- hot gas X ruled out
- black holes/white dwarfs/neutron stars X ruled out
- neutrinos X ruled out

N

*** exotic particles left over from big bang** not ruled out

The most exotic possibility is the only one left standing! more on this as we discuss cosmology...

Galaxies: Beyond the Milky Way

Edwin Hubble (1920's): galaxies fill universe typical separation $\sim 10^6$ pc = 1 Mpc (megaparsec) most distant 1000's of Mpc \Rightarrow galaxies are huge masses of stars

Galaxies sizes range large (like MW) to small "dwarf"

different structure ("morphology") & star, gas contents

spiral galaxies

- morphology: disk+bulge
- stars in both disk and bulge
- \bullet gas, dust evident in disk \rightarrow ongoing star formation

www: barred spiral zoom: star formation

elliptical galaxies

- morphology: elongated sphere, no disk!
- stars in spherical/spheroidal distribution
- no/very little gas, dust!
 - \rightarrow no ongoing star formation—no ingredients!

irregular galaxies

- morphology: no clear, simple shape/symmetry
- stars, gas, and dust all present
- www: HST merging galaxies

4

Milky Way one of many galaxies-a typical spiral

Revolution Re-Revisited

Copernican Revolution I (17th Century):

Earth is one typical planet among many not center of solar system

Copernican Revolution II (earth 20th Century):

Sun is one typical star among many not center of Milky Way Galaxy

Copernican Revolution III (1920's):

Milky Way is one typical galaxy among many Universe much larger than previously thought ... stay tuned for more...

Motions Within Galaxies

galaxies have (huge) masses

- \rightarrow each star feels gravity of all other stars
- \rightarrow all stars are in motion

in spiral galaxies, disk stars in circular orbits around center

in elliptical galaxies, stars in elliptical-like orbits around center *Q: but then how does the galaxy have a spherical shape?*

can measure rotation curves for galaxies:

- \rightarrow in fact, it's easier than measuring our
- ο own Galactic rot curve

iClicker Poll: Dark Matter in Other Galaxies

Our Galaxy has a "flat" rotation curve \rightarrow dark matter But what about other galaxies?

Vote your conscience!

How common are flat rotation curves and thus dark matter?

- A dark matter found in < 1% of galaxies
- B
 - dark matter found in < 10% of galaxies
- С
- dark matter found in < 50% of galaxies
- dark matter found in > 90% of galaxies

If it isn't dark, it doesn't matter!

⇒All galaxies have dark matter

Dark matter "halo" much larger than visible galaxy diagram

and **most** of every galaxy's mass is in the form of dark matter!

Dark Matter fills Universe!



Structure of the Universe

www: 2dF survey and map

Galaxy distribution in space:

- small scales: clumpy
- large scales: smooth

on large scales, Universe is

1. homogeneous:

galaxies fill space with (nearly) uniform density

2. isotropic:

universe looks same in all directions

5 Q: What's a U that is isotropic but not homogeneous? Q: What's a U that is homogeneous but not isotropic?

Motions of Galaxies

measure velocity respect to us i.e., in galaxy spectra, look for shifts in lines

Results:

(almost) all galaxies show redshift

all galaxies move away from us!

Hubble (1929): galaxy dist, speed related $v \propto d$, or

$$v = Hd$$

(1)

 $\stackrel{\text{\tiny L}}{\mapsto}$ \rightarrow Hubble's law

Hubble Law:

$$v = Hd \tag{2}$$

H: "Hubble constant" (sometimes written H_0) value: $H_0 \approx 70 \text{ km/(s \cdot Mpc)}$

Hubble sez: farther \rightarrow faster

draw random (homog) dots Q: add velocity vectors—what's the pattern?

What does it mean?

OBAFGKM(LT) Contest Winners