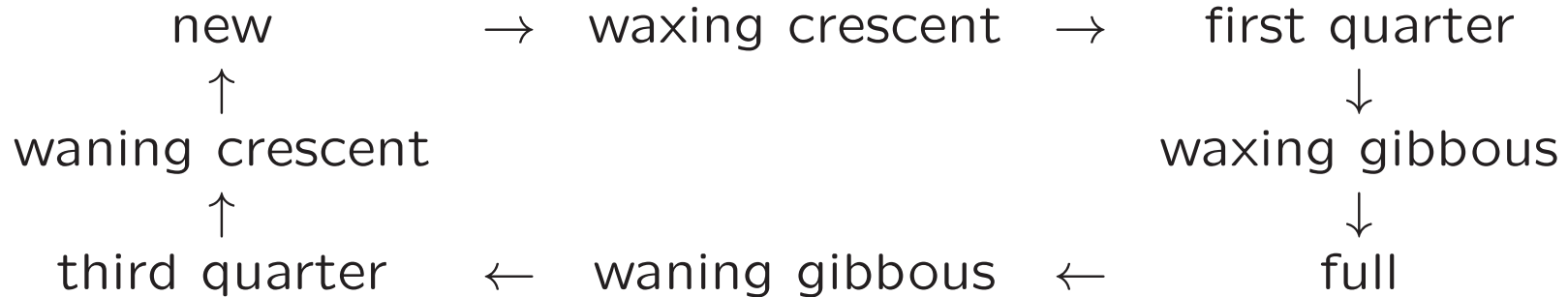


Astro 210  
Lecture 4  
Jan 26, 2011

Announcements

- HW1 posted; due at *start* of class Friday  
note juicy 10-point bonus (requires planning & digital camera)
- Office hours: Instructor 2-3pm Wed; TA 10:30-11:30am Thurs
- **register** your iClicker; link on course webpage
- first Planetarium show one week from today  
info online: **reservations**, schedules, directions, report form
- if this is your first class: see me afterward!

## Phases of the Moon



*diagram: phases as seen on sky*

*Q: what is basic physical origin of phases? Why do we sometimes see only part of the Moon illuminated?*

phases simple but beautiful

basic effect: **see illuminated moon from different angles**

phases **not** due to Earth blocking sunlight

i.e., phases are not eclipses!

*diagram: top view, sky views*

excellent exercise in translating situation in 3-D space

to 2-D projection on sky

for each Sun-Earth-Moon position, ask:

*Q: how much of Moon's surface is illuminated by the Sun?*

*Q: how much of the illuminated portion can we see from Earth?*

*Q: what does this look like in the sky?*

## iClicker Poll: Moon Phase and Rise Time

Note: each phase rises and sets at a specific time of day

When does the 1st Quarter moon rise?

- A noon
- B 3pm
- C dusk
- D midnight
- E 3am

## Phases and Latitude

Note: lunar phase pattern depends on observer latitude  
i.e., which from hemisphere you look

all of the above discussion was for *northern* observers

*Q: what aspects of the situation change for southern observers?*

*Q: what aspects do not change*

southern observers: “upside-down” relative to north  
(and vice versa)

→ moon image same (illuminated vs dark portion identical)  
but “upside-down”

unchanged:

- waxing or waning nature of phase  
(i.e., increasing or decreasing illumination)
- type of phase (crescent, gibbous, etc)

different:

- southern observers see reversed moon right ↔ left  
relative to northern observers
- e.g., waxing crescent illuminated on *left* side  
waning gibbous illuminated on *right* side

◦ For all homework, exam problems: assume *northern* observers  
on earth (and on the Moon!)

my suggestion: practice!

do the experiment in real life!

when moon up: point to moon, then sun, look at angle!

complete cycle of lunar phases in 29.5 days

⇒ 3rd fundamental measure of time: **month of phases**

*Any questions?*

# Eclipses

**Lunar Eclipse:** moon in earth's shadow

*diagram: Sun, Earth, Moon*

www: lunar eclipses

note: can still see Moon even when totally in Earth's shadow!  
appears much dimmer, and red

*Q: what's going on? why the red color?*



note that *direct* sunlight is totally blocked  
so light must be indirect, in fact:  
scattered light from earth's atm.  
red b/c blue is scattered more strongly, so only red is left  
in other words:

glow is from all the sunrises and sunsets on Earth!

**solar eclipse:** observer in moon's shadow

Note: Earth larger than Moon

*Q: what does this immediately imply for solar eclipses?*

since Moon smaller than earth, whole earth cannot be in shadow  
in fact, only a small region,  $\sim 100$  mi, at a given time  
“eclipse path”

interesting coincidence: Moon and Sun have  
almost identical angular size

*and what's more:* Moon's distance changes (not circular orbit)

www: moon perigee/apogee comparison

together this means: two kinds of solar eclipses

www: annular eclipse

www: total eclipse

www: looking back on Earth's shadow

10 Note: there is *not* a solar or lunar eclipse every month!

if Moon's orbit plane around Earth  
were same as Earth's around the Sun (i.e., the ecliptic)  
then *would* have eclipses monthly

the fact that we *don't* means that  
the two orbits are *not* coplanar!

★ moon's orbit plane slightly tilted w.r.t. ecliptic  
so moon is typically below or above ecliptic

www: eclipse diagram

only eclipse when Moon orbit crosses ecliptic plane  
happens *twice* a year → “season of eclipses”

geometrically: intersection of planes is a *line*  
and intersection of line with closed orbits is *2 points*

Note: eclipse “season” last about  $\sim 1$  month  
i.e., time window of alignment about equal to Moon orbit period  
→ can sometimes have two or three eclipses in same season

*Next eclipses: partial solar June 1 and July 1, lunar June 15  
watch over summer, impress someone with your ASTR210 skills!*

*Any questions?*

# Planets

Known since ancient times:

Mercury, Venus, Mars, Jupiter, Saturn

Discovered later, with telescopes:

Uranus, Neptune, Pluto

**daily motion:** westward w.r.t. horizon  
i.e., .rise in east, set in west

w.r.t. fixed stars:

always stay close to ecliptic

usually move eastward

www: SOHO LASCO planet movie

but sometimes westward: "retrograde motion"

www: retrograde motion of Mars

so motion not uniform in angular speed

in fact, not uniform even when in "direct" (non-retro) motion

retrograde motion not random in occurrence!

key patterns observed:

- for each planet, retrograde onset is *periodic* with different periods for different planets
- retrograde occurrence *correlated* with position relative to Sun:

Mercury, Venus	Mars, Jupiter, Saturn (& others)
always stay close to sun on sky <i>never</i> seen opposite Sun	can move freely along ecliptic can be opposite sun
retro when in <b>conjunction</b> (i.e., when closest to Sun on sky)	retro when in <b>opposition</b> (i.e., opposite Sun on sky)

## These Patterns Cry Out For Explanation

you may have noticed—I've heaped a lot of facts on you.

Do you have to memorize them? Do *I* have them memorized?

No! There's a simpler way of remembering.

→ build a **model** of the solar system's geometry and dynamics  
organize, explain all of this data!

Science is built up with facts, as a house is with stones.

But a collection of facts is no more a science  
than a heap of stones is a house.

—Henri Poincaré

Crucial point:

when making model for motions of planets

have to explain *all* observed features;

turns out the retrograde motion, in all its detail,

gave people the hardest time...

## Building a Scientific Model

Scientific Models must:

- explain observations
- predict future observations

The principle of science, the definition, almost, is the following: *The test of all knowledge is observation.* Experiment is the *sole judge* of scientific “truth.”

The first principle is that you must not fool yourself—and you are the easiest person to fool.

—Richard Feynman