

Astro 210
Lecture 18
March 2, 2011

Announcements

- HW5 due at start of class Friday
 - **Oops!** – typos found in Problem 2(d) which should refer to the situation in 2(c) erratum and corrected question posted
- Night Observing this week – *Dress warmly!* report forms, info online

Last time:

- atmospheres & properties of gasses
- ↳ *Q: how is a gas microscopically different from a solid or liquid?*
Q: what is a gas temperature, at a microphysical level?

Last time: gasses under a microscope

gasses: collections of particles, e.g., atoms, molecules

www: gas law simulation

- in space: “elbow room” – empty space between particles
- in time: constant random motion
collisions exchange energy, momentum
- *individual* particle velocities random, changing
but *distribution* (=histogram=bar graph) of particle speeds
is set by **temperature**

→ average speed $v_{\text{rms}} = \sqrt{3kT/m}$

Microscopic View of a Piston

Now consider a large number of gas particles

- in a sealed volume
- with a *piston* of area A

from microscopic viewpoint:

piston constantly bombarded by gas particles

if let free—would be pushed away

to resist bombardment, must *push* on piston = exert *force* F

define **pressure**

$$P = \frac{\text{force on piston}}{\text{area of piston}} = \frac{F}{A} \quad (1)$$

^ω Q: *how to intensify bombardment = pressure on piston?*

Hint—more than one way to do this

Pressure

collisions with walls → momentum transfer → force → pressure

www: piston simulation

ideal gas

pressure P , volume V , total number N of particles and **absolute** (Kelvin) temperature T all related by *ideal gas equation of state*:

$$PV = NkT \quad (2)$$

- N counts individual particles, typically very large!

alternatively: can count in units of moles of particles

i.e., in units of $N_{\text{AvO}} = 6 \times 10^{23}$

then # moles is $n_{\text{moles}} = N/N_{\text{AvO}}$

and $PV = n_{\text{moles}}RT$, where $R = N_{\text{AvO}}k = 8.3 \text{ Joules mole}^{-1} \text{ K}^{-1}$

- ↳ • since density $\rho = M/V = \mu N/V$, with μ gas particle mass can rewrite ideal gas law: $P = \rho kT/\mu$

Planetary Atmospheres

Terrestrial Atmospheres:

- atmospheres are tiny (or zero!) fraction of planet mass
- no light gasses (H_2 , He),
only heavier N_2 , O_2 , CO_2 (if anything!)

Jovian Atmospheres:

- a significant fraction of planet mass
- mostly H_2 , He; some heavier species

Q: why the difference? what factors important?

competition: gravity versus thermal motion

gravity → keep particles

thermal motion → run away

- gravity → escape speed

$$v_{\text{esc}} = \sqrt{2GM/R} = 2.4 \times 10^3 \text{ m/s for Moon}$$

- thermal motion → avg thermal speed

$$v_{\text{rms}} = 2.6 \times 10^3 \text{ m/s for H on Moon}$$

$v_{\text{rms}} > v_{\text{esc}}$: many atoms can escape

H lost from Moon (check!)

really: not all particles have same speed

so **always** some escape

real question: timescale

to **keep** atm for age of SS

need $v_{esc} \geq 6v_{rms}$ (rule of thumb)

species	v_{esc}/v_{rms}	lifetime
H on moon	0.9	hours
H on earth	2.1	days
sodium on moon	4	$\sim 10^3$ yrs
O ₂ on earth	12	$\rightarrow \infty$: no escape!

www: Sodium lunar ‘‘atmosphere’’

solar UV photon on Moon ‘‘soil’’ \rightarrow thin vapor of sodium

↘ ejected but replenished

LET'S GO: SOLAR SYSTEM

The Earth

astro-trivia: Earth logo/icon is \oplus

recall: $R_{\oplus} = 6.4 \times 10^6 \text{ m} = 6400 \text{ km}$

get mass from $g = GM_{\oplus}/R_{\oplus}^2$ (need G !)

$\rightarrow M_{\oplus} = gR_{\oplus}^2/G = 6.0 \times 10^{24} \text{ kg}$

Average density:

$$\rho_{\text{av},\oplus} = 3/4\pi M_{\oplus}/R_{\oplus}^3 = 5,500\text{kg}/\text{m}^3$$

between rocks and iron \rightarrow some of both

Earth's Interior

crust: 16–40 km → *thin!*

mantle: ~ 3000 km

outer core: ~ 2,200 km

inner core: ~ 1,200 km

crust: solid. ocean basins—basalt: O, Si, Al, Mg

continental plates—granite: O, Si, Al, Na, K

mantle: “plastic”. Fe, Mg, Si, O

outer core: liquid. Fe, Ni

inner core: solid. Fe, Ni

heaviest elements lowest → settling (“differentiation”)

Q: how do we know?

How do we know?

“Refraction” of Earthquakes

Demo: slinky

Earthquakes “emit” waves → use seismographic info to learn earth structure

www: diagram

Plate Tectonics

www: plot of earthquake sites

www: plot of volcano sites

www: plate locations

crust not a single rigid solid

but collection of “plates”

motions in mantle (convection) cause plate motion

www: convection currents

www: satellite laser ranging

www: VLBI: radio telescopes used to detect motion

www: drift animation

plates move, and interact: collide, slide, buckle

⇒ “plate tectonics”

leads to observed geological features

www: Mountain

www: volcano (Kilauea, HI)

www: San Andreas

ex: San Andreas fault in CA: sliding plates ⇔

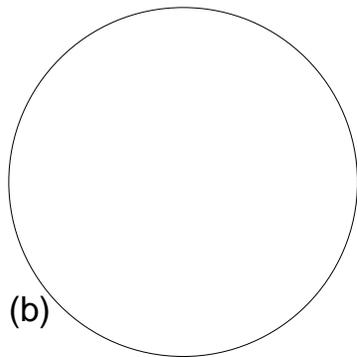
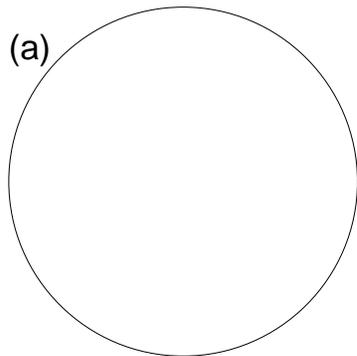
Earth is evolving!

iClicker Poll: Earth's Orbit

Earth orbit around Sun: ellipse, eccentricity $e_{\oplus} = 0.017$

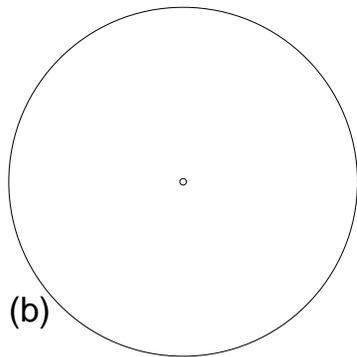
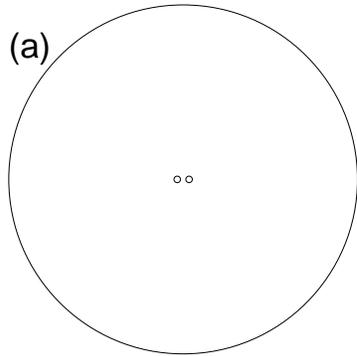
Below: one **ellipse** with $e = e_{\oplus}$, one **circle** $e = 0$

Which of these has the eccentricity of Earth's ellipse?



Survey Says

Foci labeled



Seasons and Eccentricity

Due to elliptical orbit, Earth-Sun distance changes

$$r_{\text{ap}} = (1 + e)a_{\oplus} = 1.017 \text{ AU}$$

$$r_{\text{peri}} = (1 - e)a_{\oplus} = 0.983 \text{ AU}$$

...but not by much! 3.2% swing!

if this were the whole story:

temperature swing by $\approx 2\% \times 300 \text{ K} = 6^{\circ} \text{ C} = 11^{\circ} \text{ F!}$

also: if yearly temperature variations were due to eccentricity...

Q: how should the seasons compare in the N and S hemispheres?

Q: what is actually observed?

Q: what does this imply about season origin?

Origin of the Seasons

seasons are **exactly opposite** in N and S hemispheres
i.e., right now is spring in S. America, will soon be summer

so: season origin must distinguish hemispheres

recall: Earth's spin axis tilted w.r.t. orbit by 23.5°

- when one hemisphere tilted closest, the other farthest:
this is summer/winter
- when tilt \perp Earth-Sun radius: fall/spring

Note: Earth closest to Sun in northern winter!

→ eccentricity has tiny effect on temperature