

Astronomy 150: Killer Skies

Lecture 33, April 16

Guest Lecturer today!

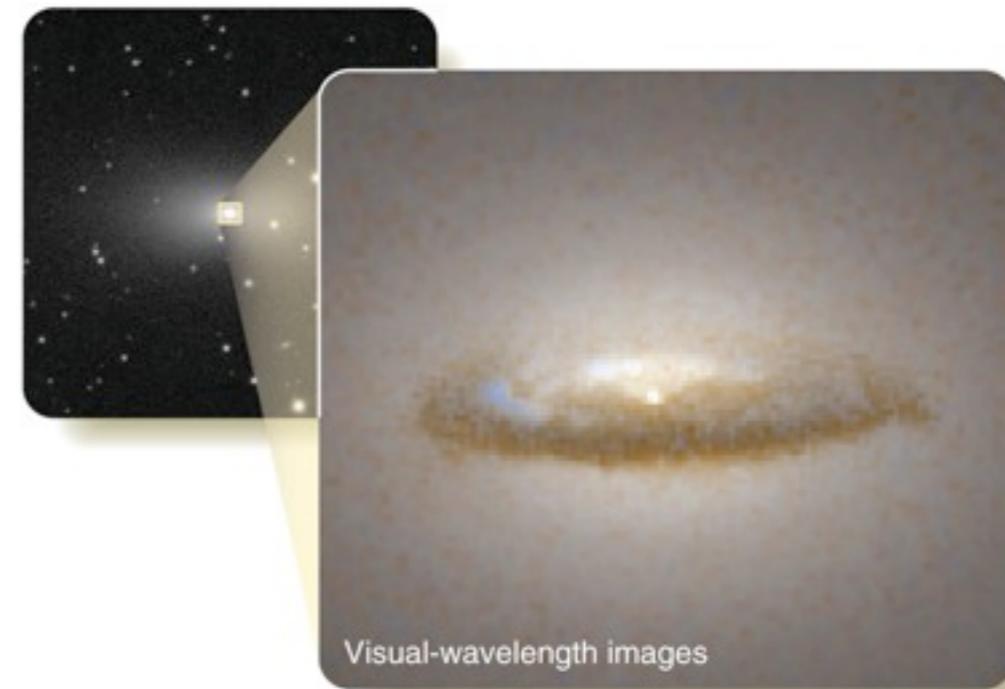
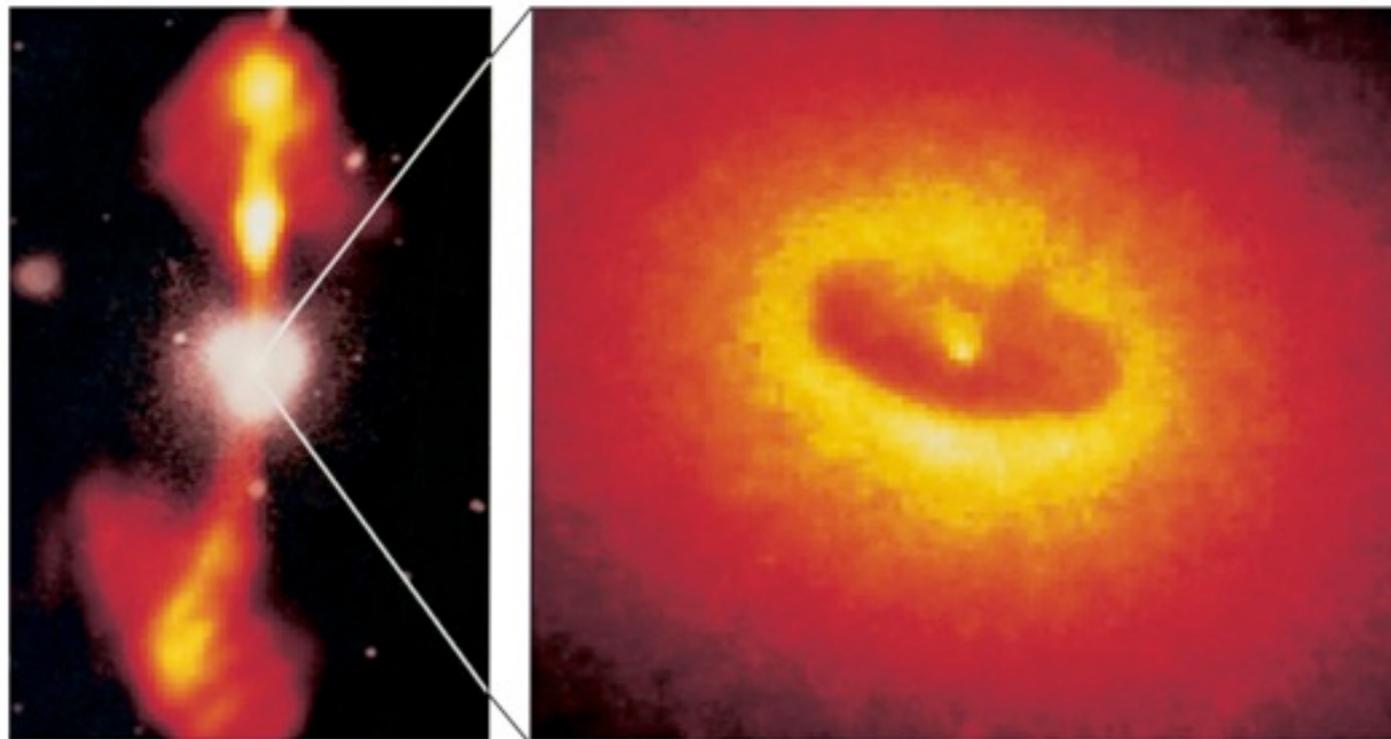
- ▶ Prof. Robert Brunner, expert on galaxies!

Assignments:

- ▶ HW10 due Friday
- ▶ Computer Lab 2 also due Friday
- ▶ so you can work on this: **no lecture on Wednesday!**
see you again on Friday

Last time: Cosmic Blowtorches

Today: **Our Milky Way Galaxy**



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Computer Lab 2

Another chance to analyze real research-grade data and draw conclusions about the cosmos

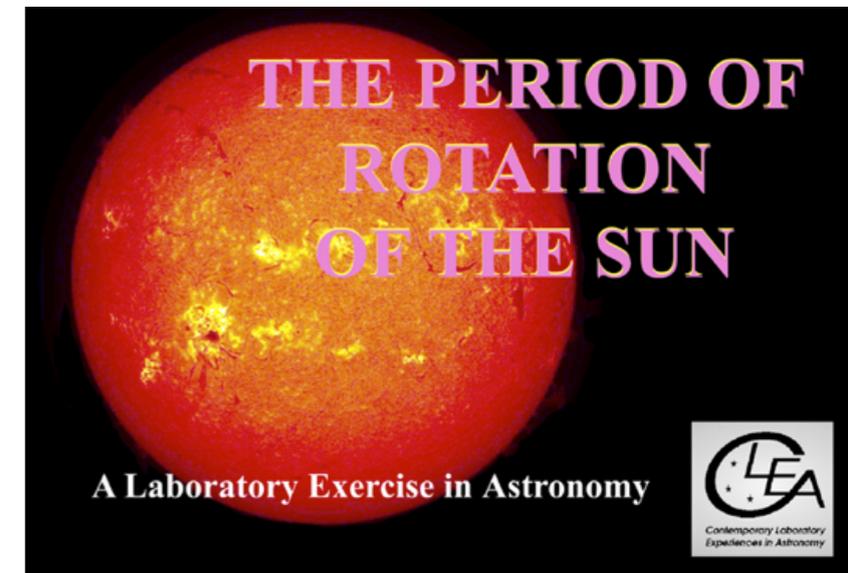
You can choose one of two:

- ▶ Solar Rotation
- ▶ Galaxy Zoo

details on course website

“Second Chance”: **If you do both labs** then

- ▶ higher score will count as your Lab 2 score
graded out of 100, counts for 50 points out of 1000 on final grade
- ▶ lower score will **replace up to 20 points** on final grade total for **one** low Homework, Observing, or Lab 1 score
not “extra” points on final point total, **only replaces one** existing low score
- ▶ in practice: **maybe worth doing if you have any** of these
 - a homework score < 20 out of 20
 - a Lab 1 score < 40 out of 100
 - an Observing or Planetarium score < 20 out of 25



just so we are totally clear...

No lecture Wednesday!
Work on Computer Lab!

See you this Friday
when the lab is due.

Recap: Our Milky Way Galaxy

Milky Way in sky (2D):

- ▶ irregular diffuse glow
- ▶ circles the celestial sphere
- ▶ most of MW light is from huge numbers of distant stars



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Milky Way in 3D space:

- ▶ a **disk** of stars
- ▶ **razor-thin**: like stack of 3 DVDs



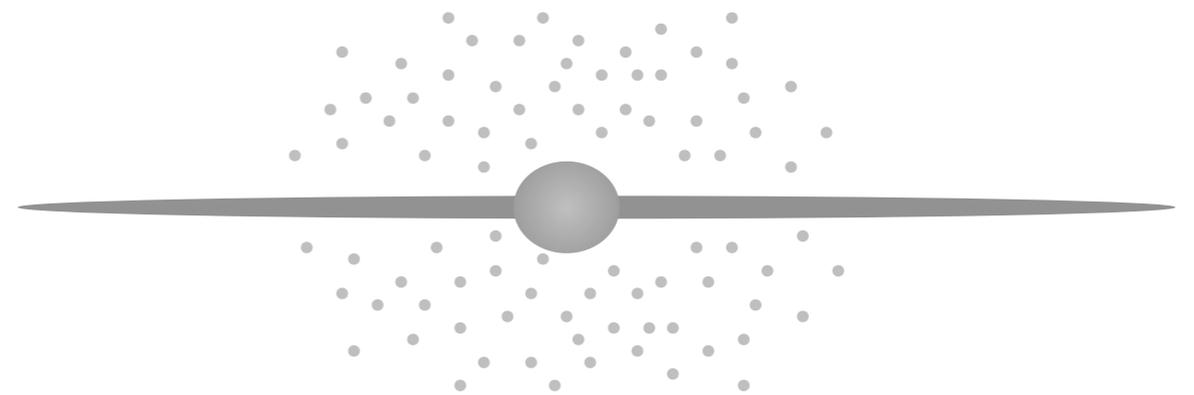
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Milky Way disk: edge-on view

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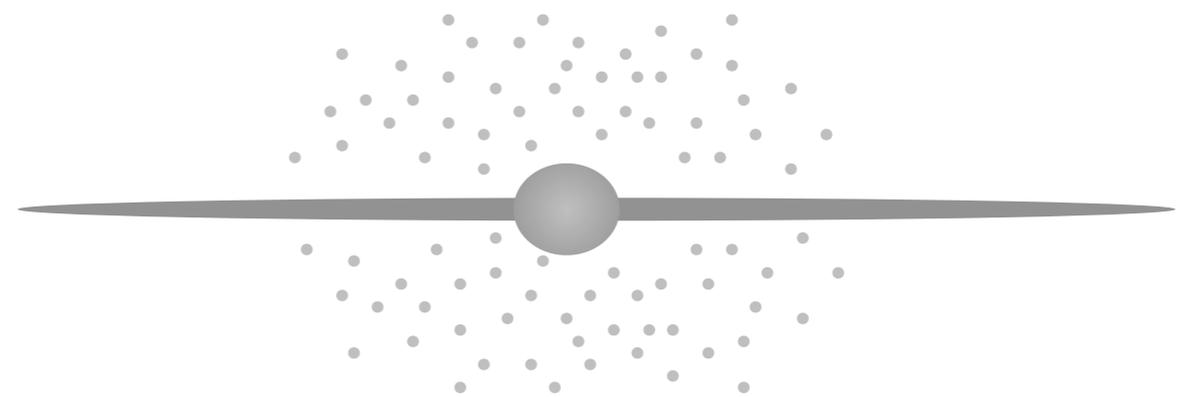
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- ▶ Herschels: looks like we're at center
- ▶ but: obscuring dust fills MW
- ▶ Shapley: look at pattern of globular clusters of stars



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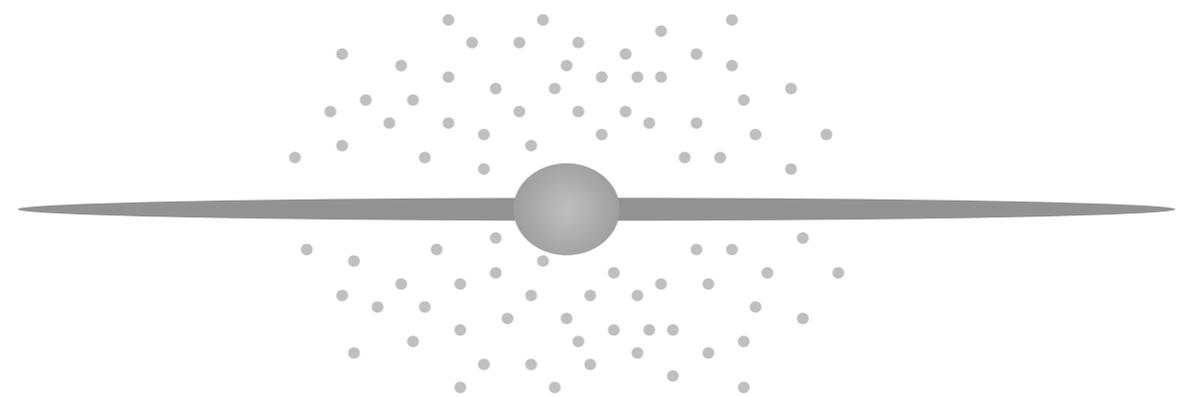
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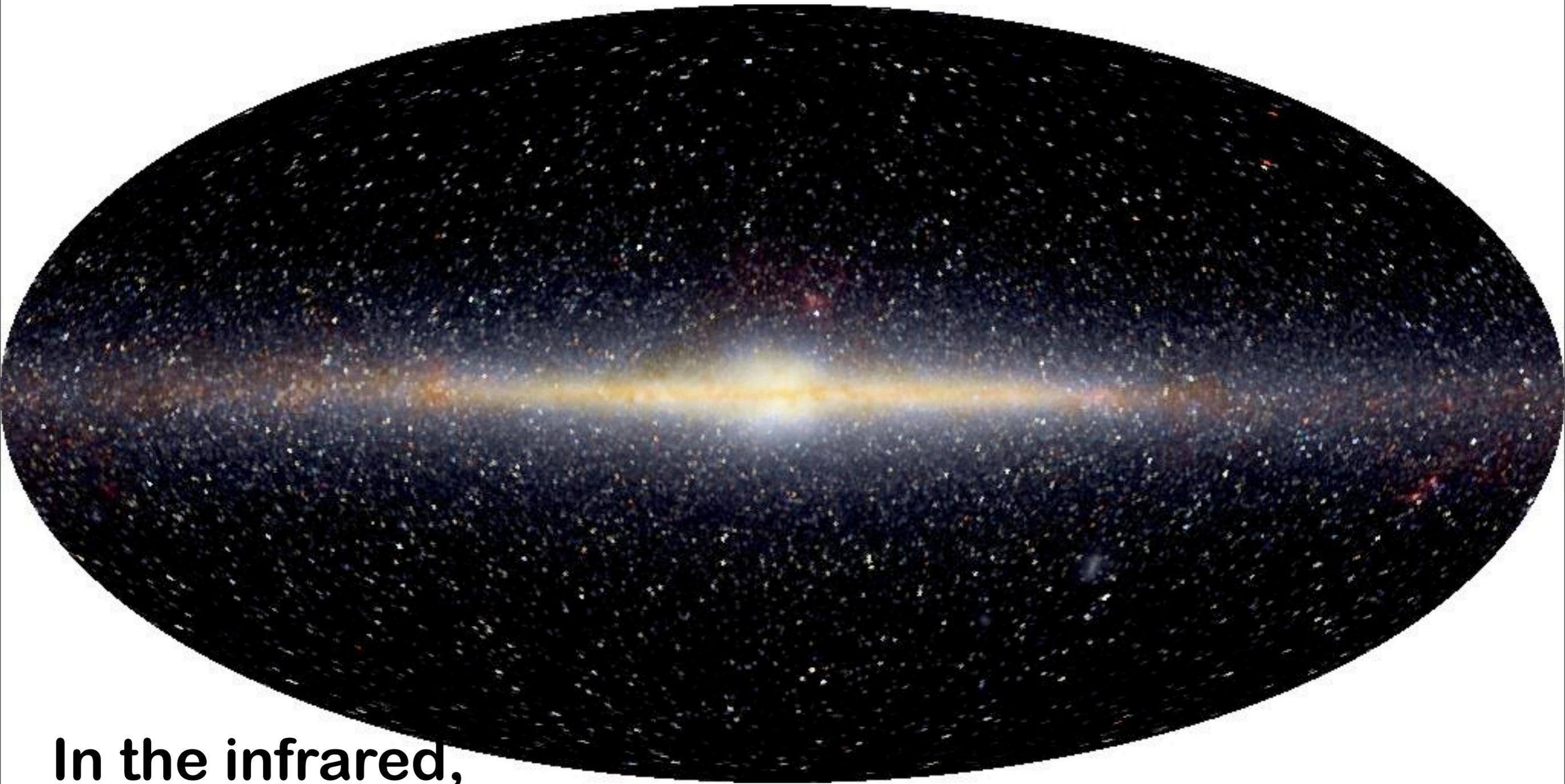
Result: we are off-center!

- ▶ we live in Milky Way suburbs!



Milky Way disk: edge-on view

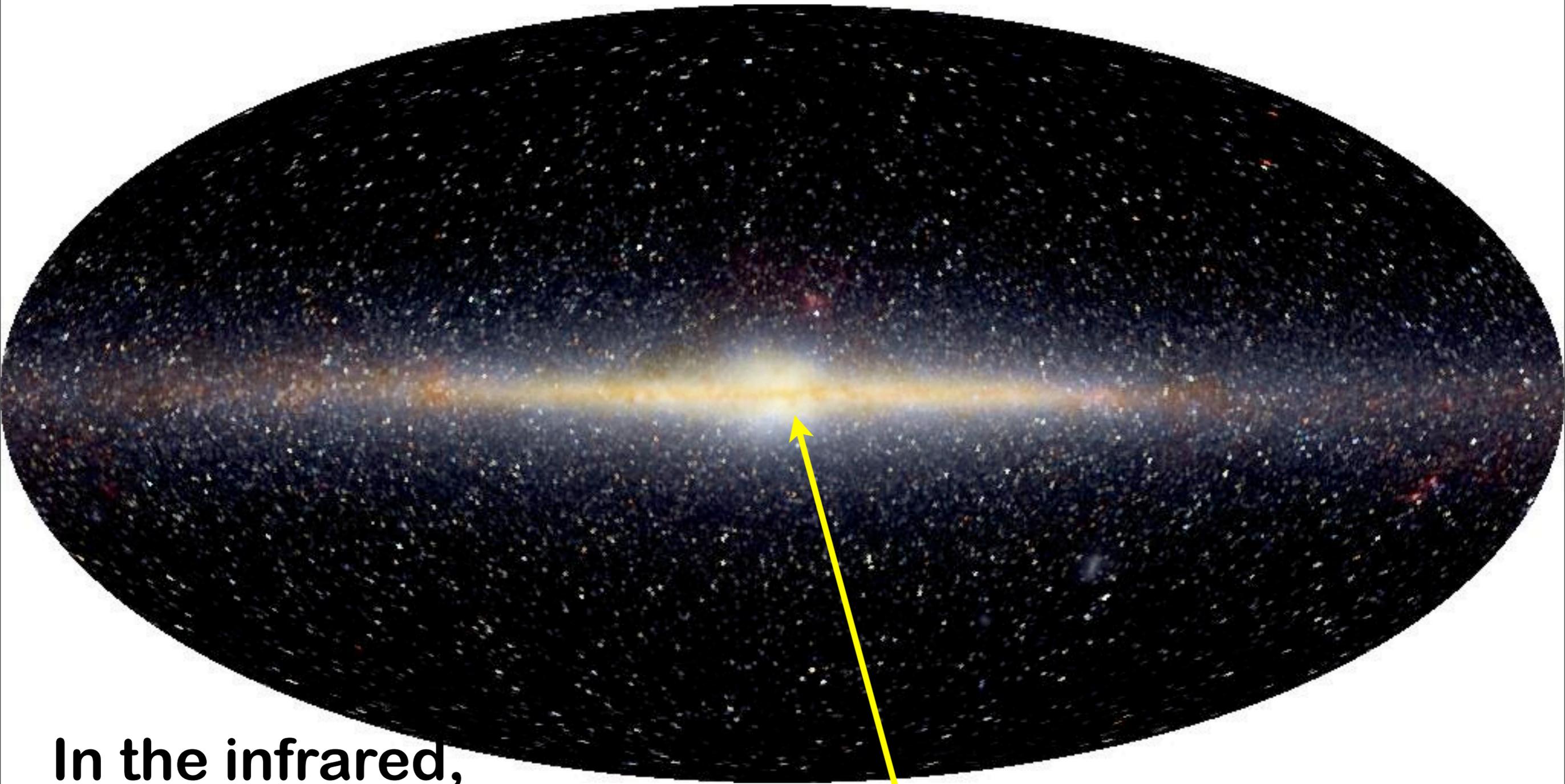
Shapley Vindicated



**In the infrared,
can see through dust**

http://antwrp.gsfc.nasa.gov/apod/image/0001/milkyway_cobe_big.jpg

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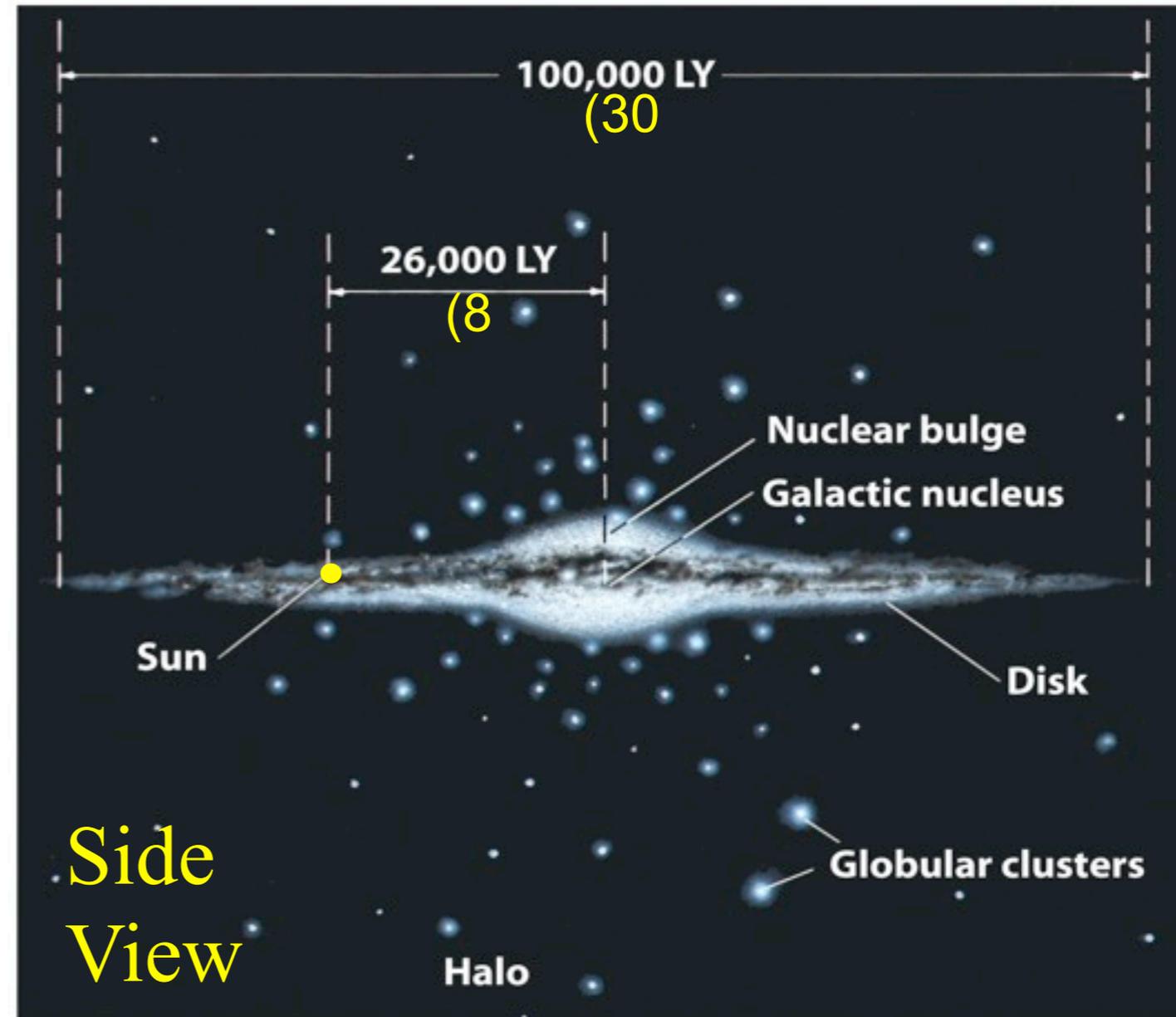
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Center of the Galaxy

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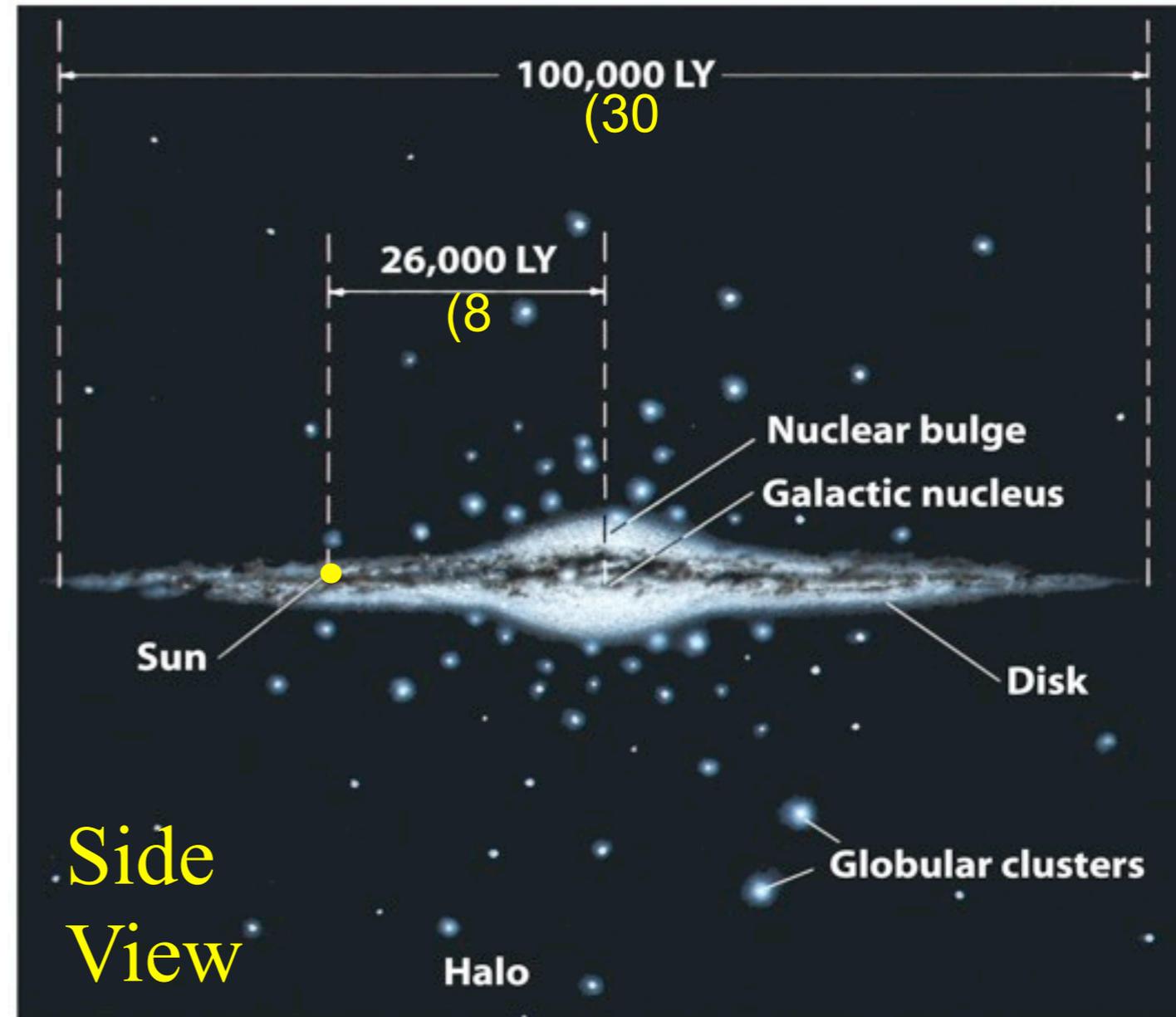
Our Galaxy

- **Globular clusters—
oldest stars**



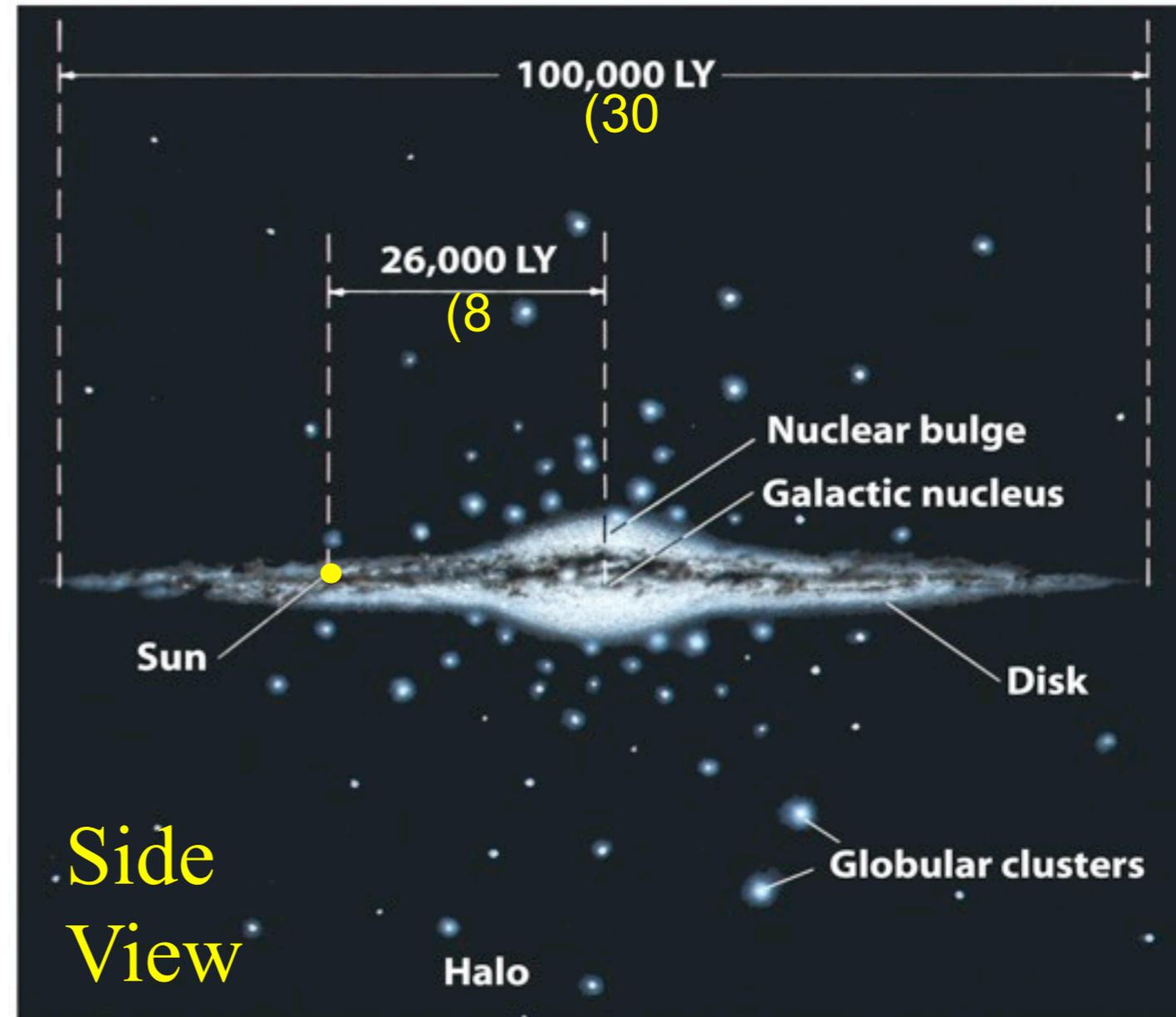
Our Galaxy

- Globular clusters— oldest stars
- Galactic nucleus— dense collection of stars (center of Galaxy)



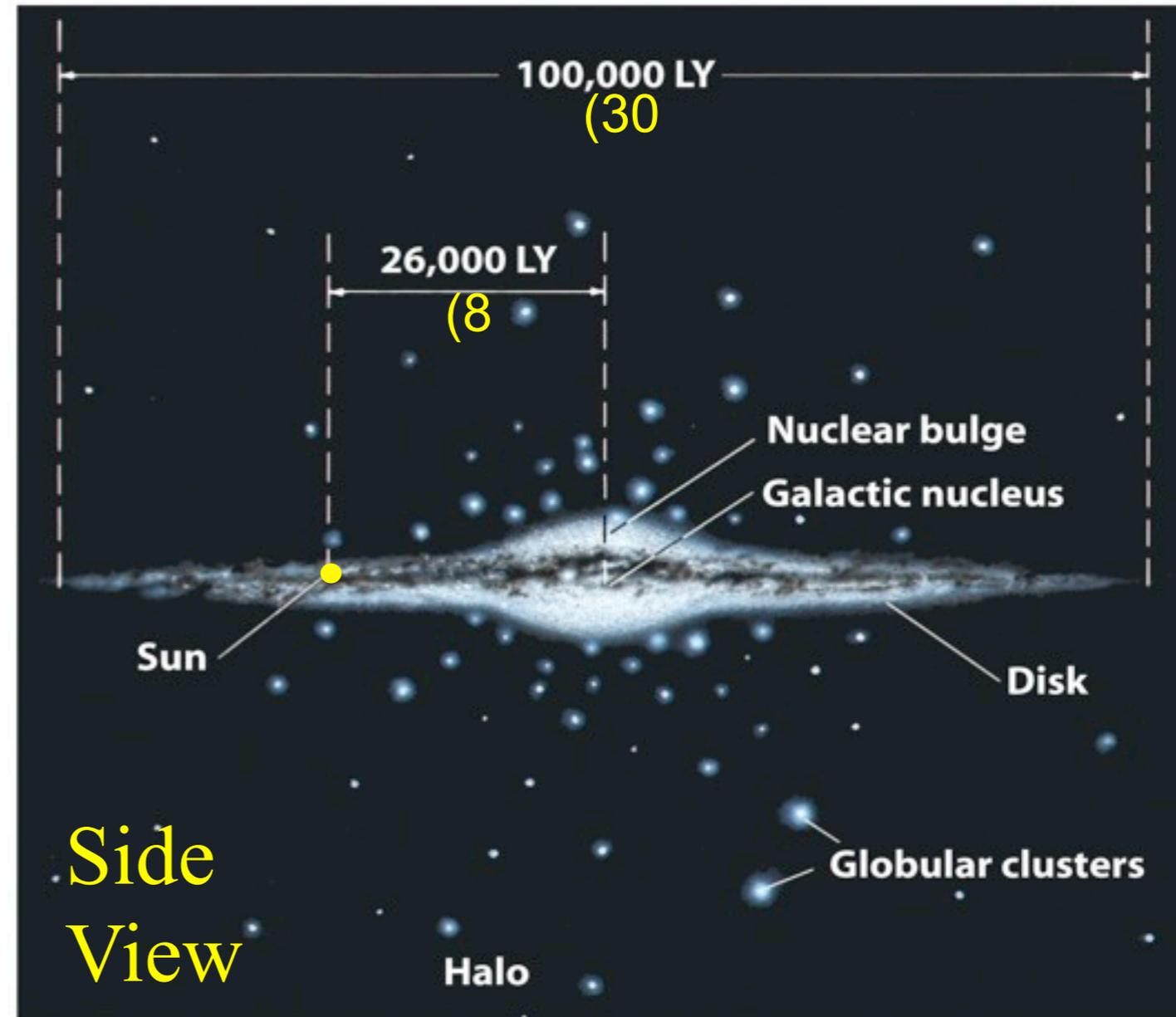
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- **Globular clusters**– oldest stars
- **Galactic nucleus**– dense collection of stars (center of Galaxy)
- **Nuclear bulge**– mostly old stars, but very densely packed



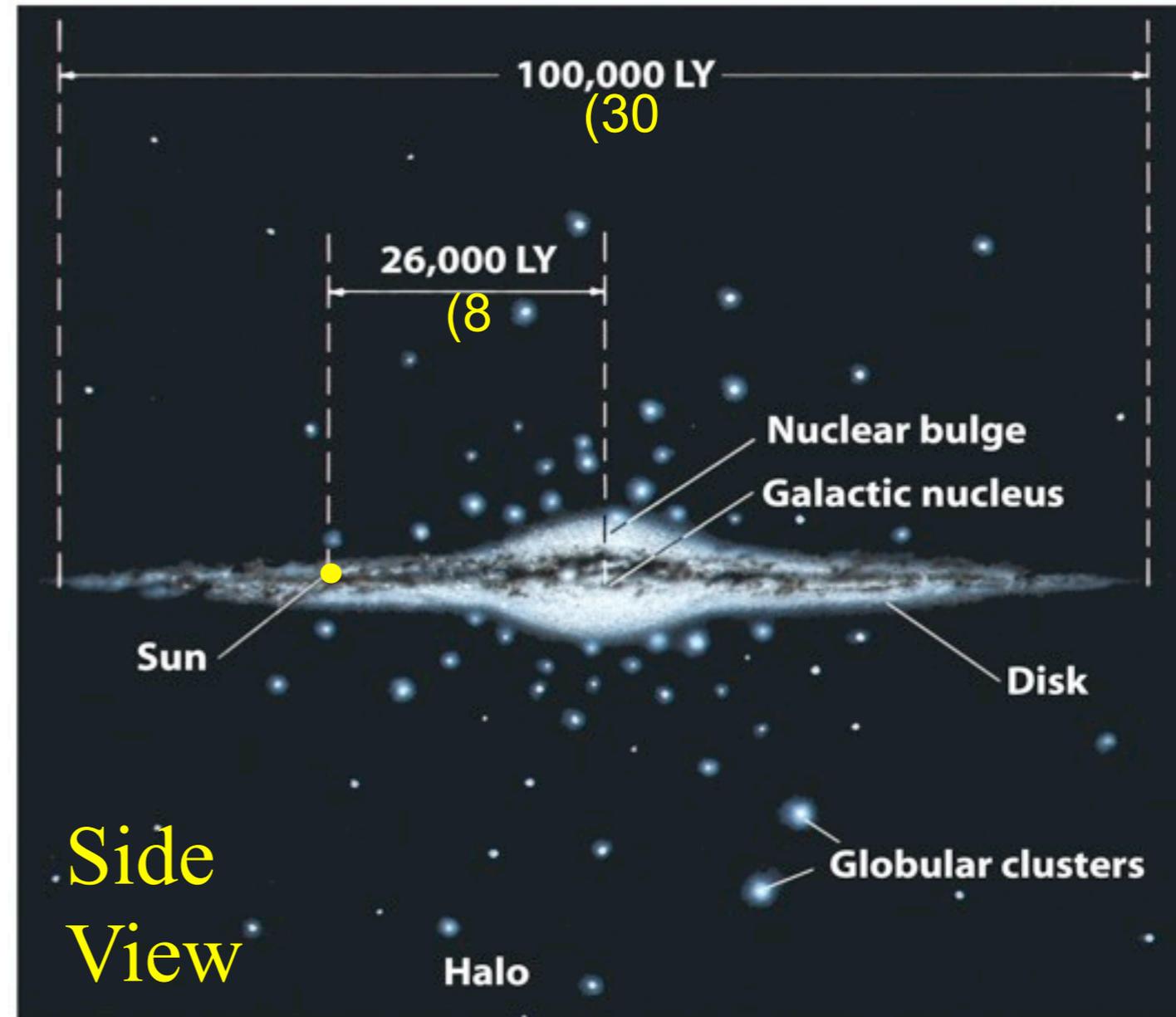
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Our Galaxy

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- Spiral arms and the disk– mostly young stars and lots of dust
- Note position of the Sun, just over half way out.



The Disk

The disk of our Galaxy contains most of its visible mass

- ▶ **90% of the Galaxy's stars**

It's where “the action” occurs

- ▶ **Star formation, nebulae, etc..**

Relatively thin

- ▶ **1,000 lyrs thick vs. 100,000 lyrs across**



The disk in infrared light

Spiral Arms?

Other disk galaxies show spiral arms

- ▶ **Made massive stars, diffuse nebulae, and most of the giant molecular clouds**

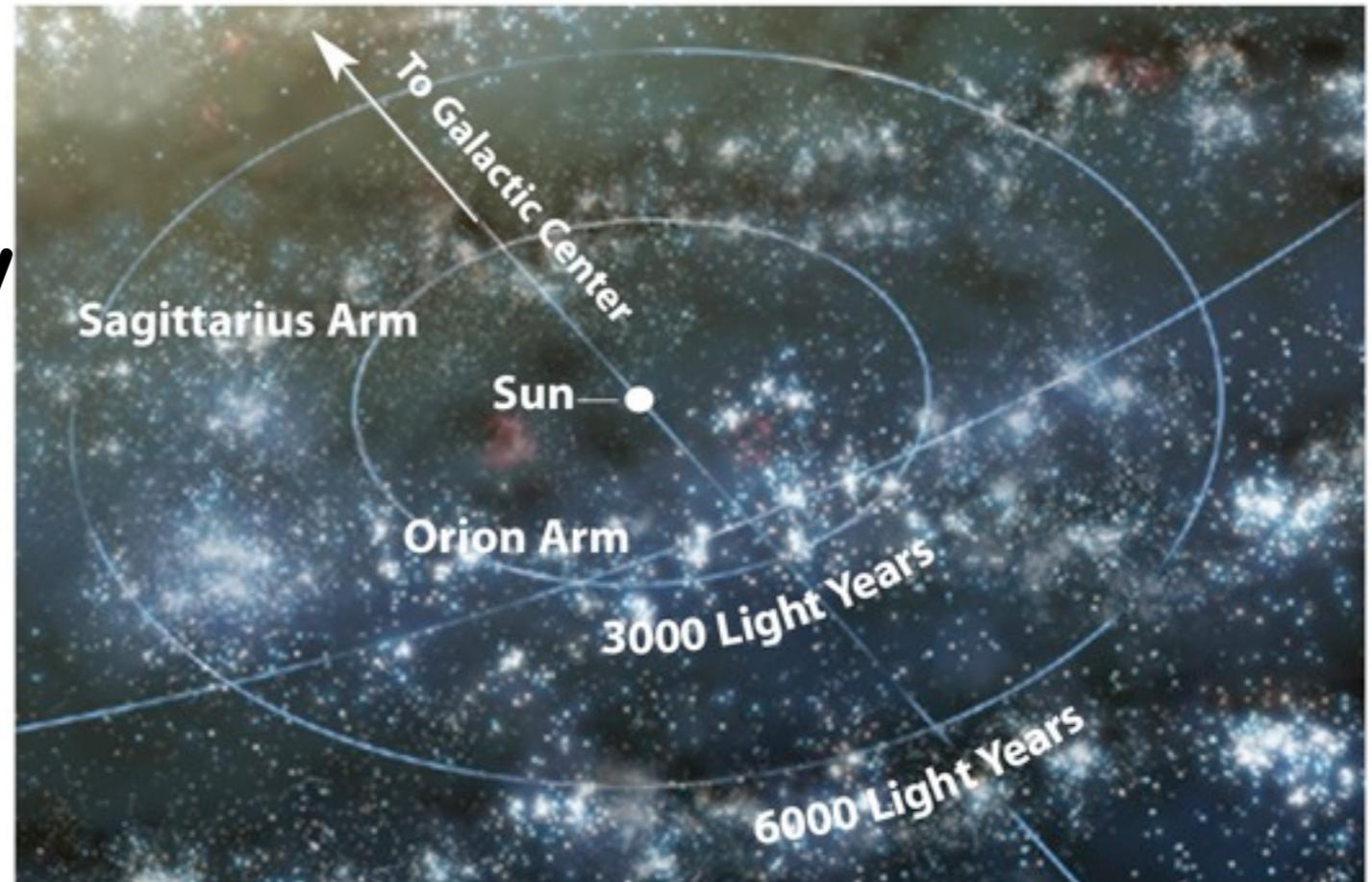
How do we know our Galaxy has them?

It's the problem of not seeing the forest for the trees

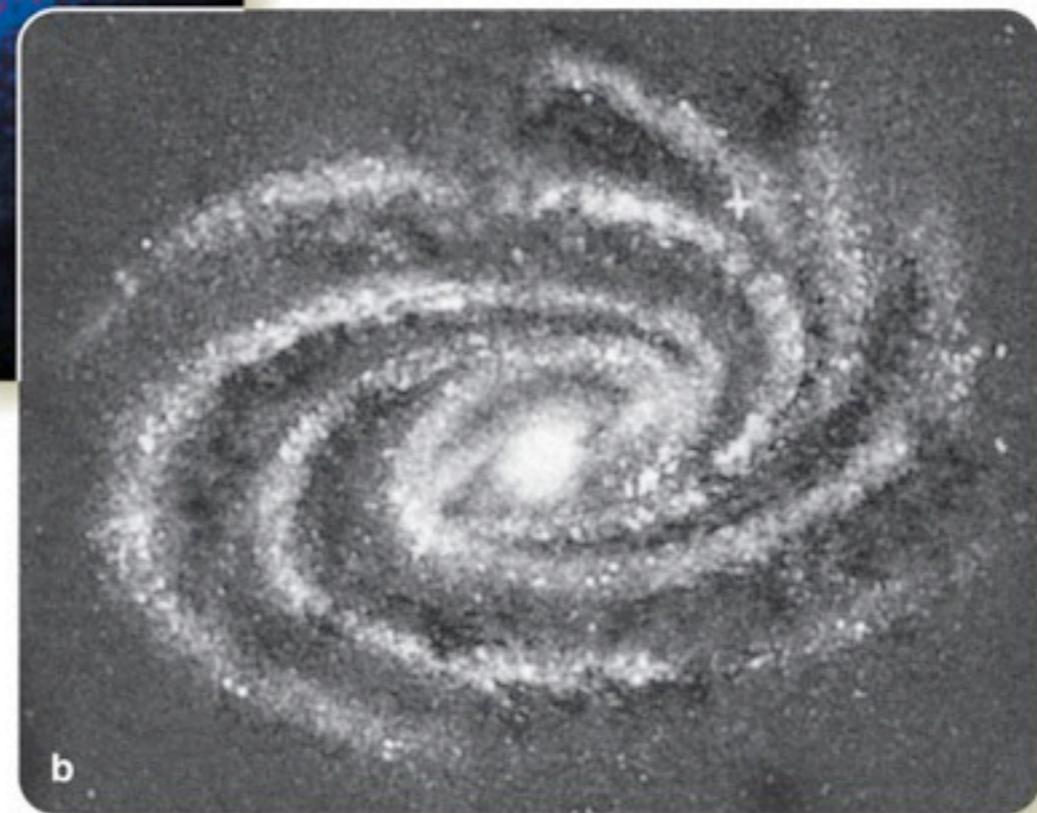
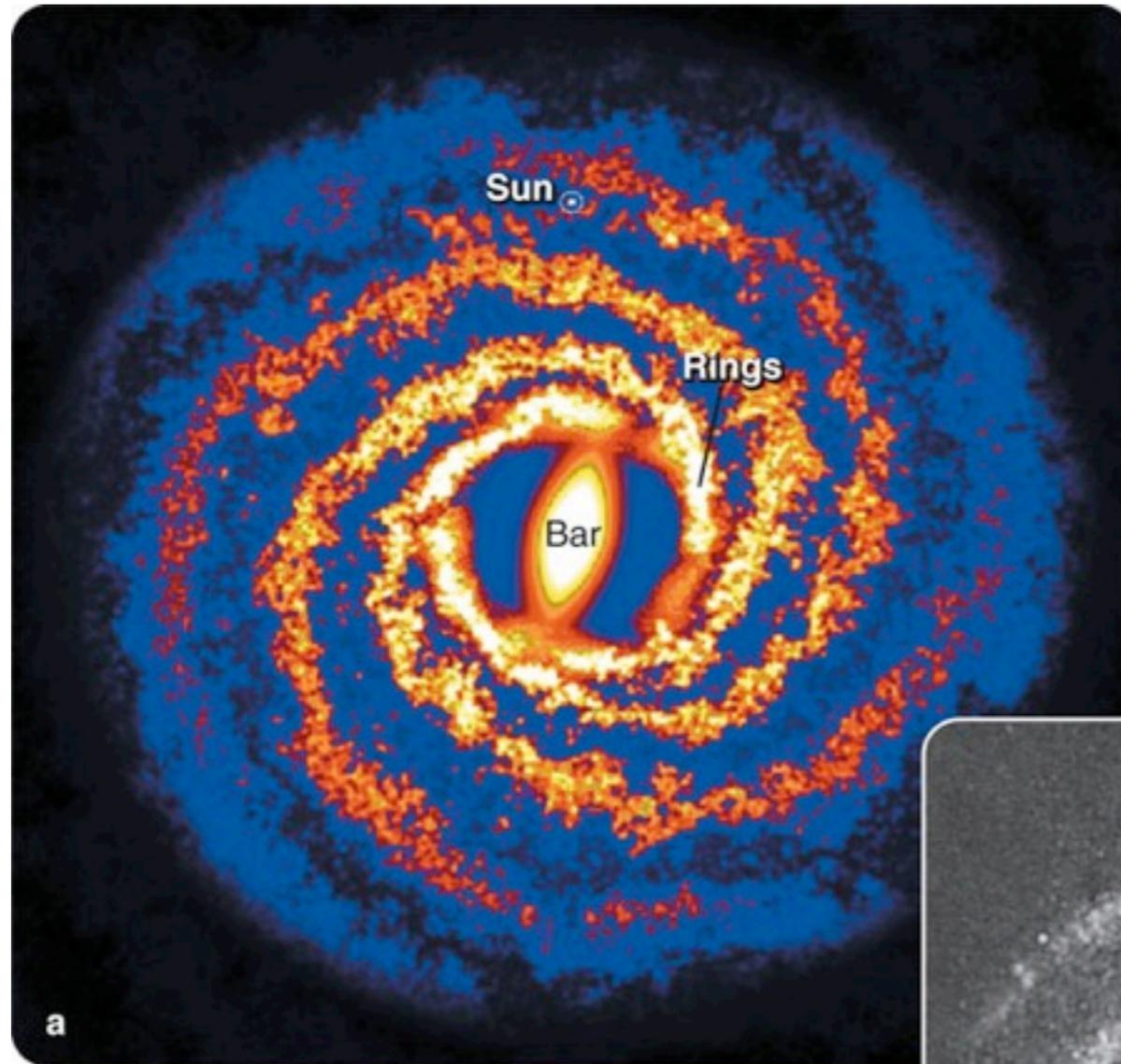


Hints of Spiral Arms

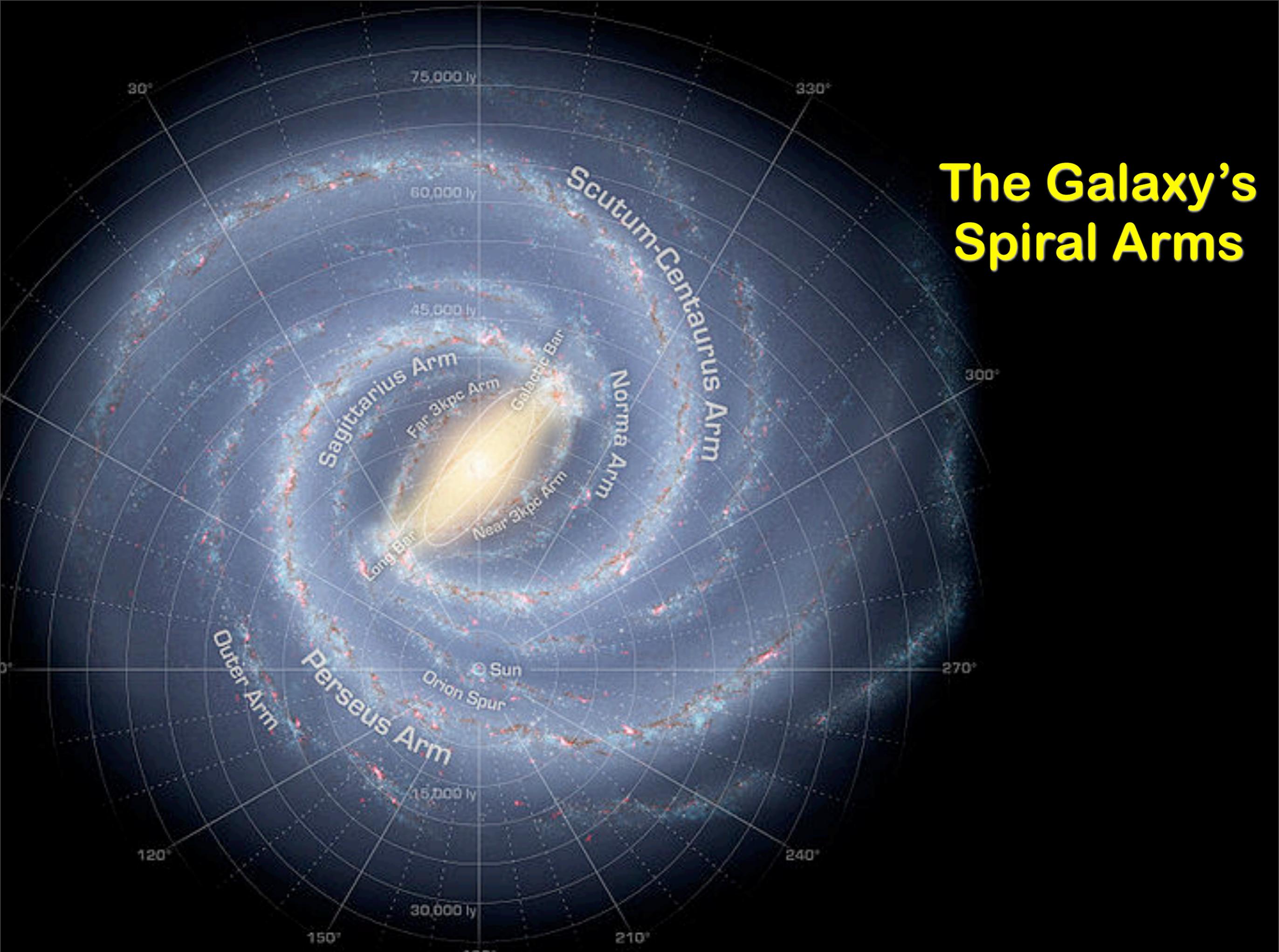
We plot the locations of nearby massive, bright stars in our Galaxy
Find the stars are arranged in arms
Our Sun is in-between spiral arms



The Galaxy's Spiral Arms



The Galaxy's Spiral Arms



Milky Way: Motions

- Milky Way contains about 100 billion stars**
- ▶ **and thus has mass of at least $10^{11} M_{\text{sun}}$**

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But what is the pattern of motion?

Do Galaxies Spin?



Spiral galaxies really suggest it. Our Galaxy probably looks more like the right galaxy.

Rotation of the Galaxy



Rotation of the Galaxy

Similar to the planets orbiting the Sun, the stars and gas of the Galaxy orbit the nucleus



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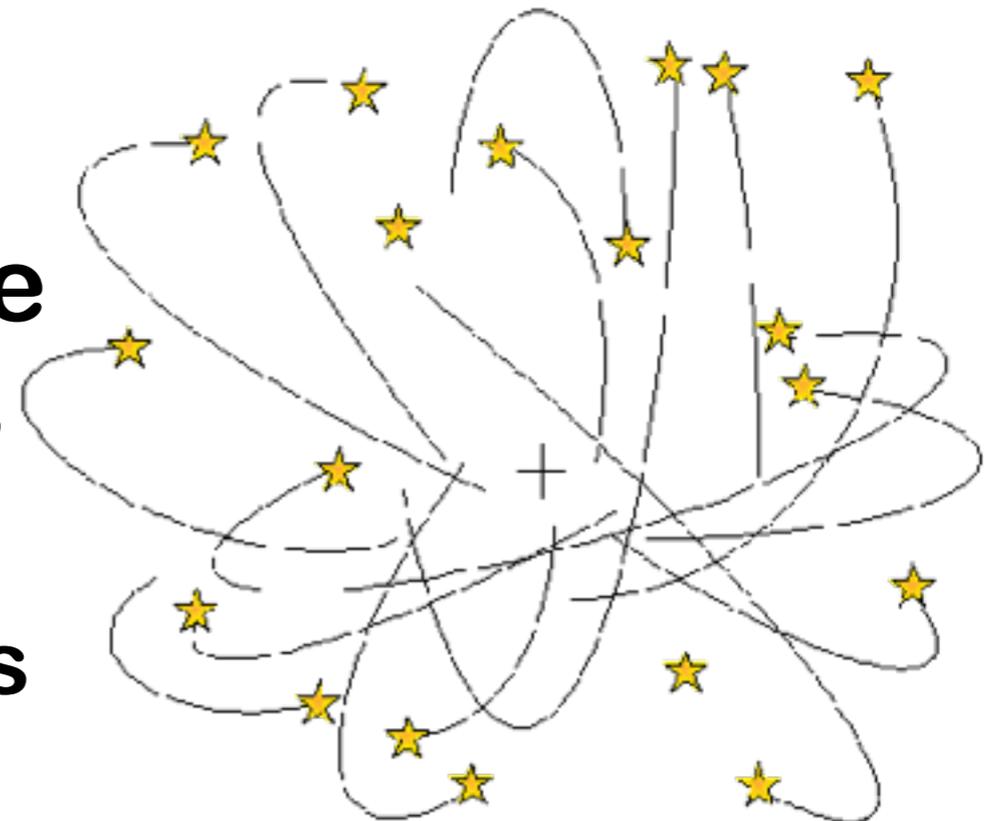
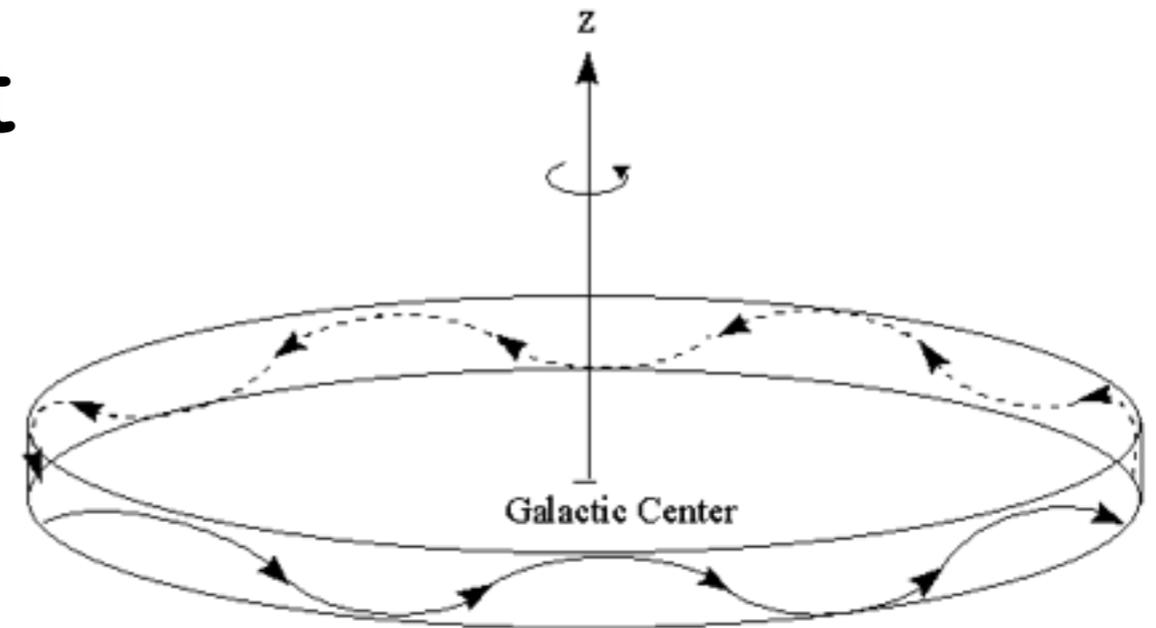
Rotation of the Galaxy

Stars in the disk all orbit the Galaxy in the same direction

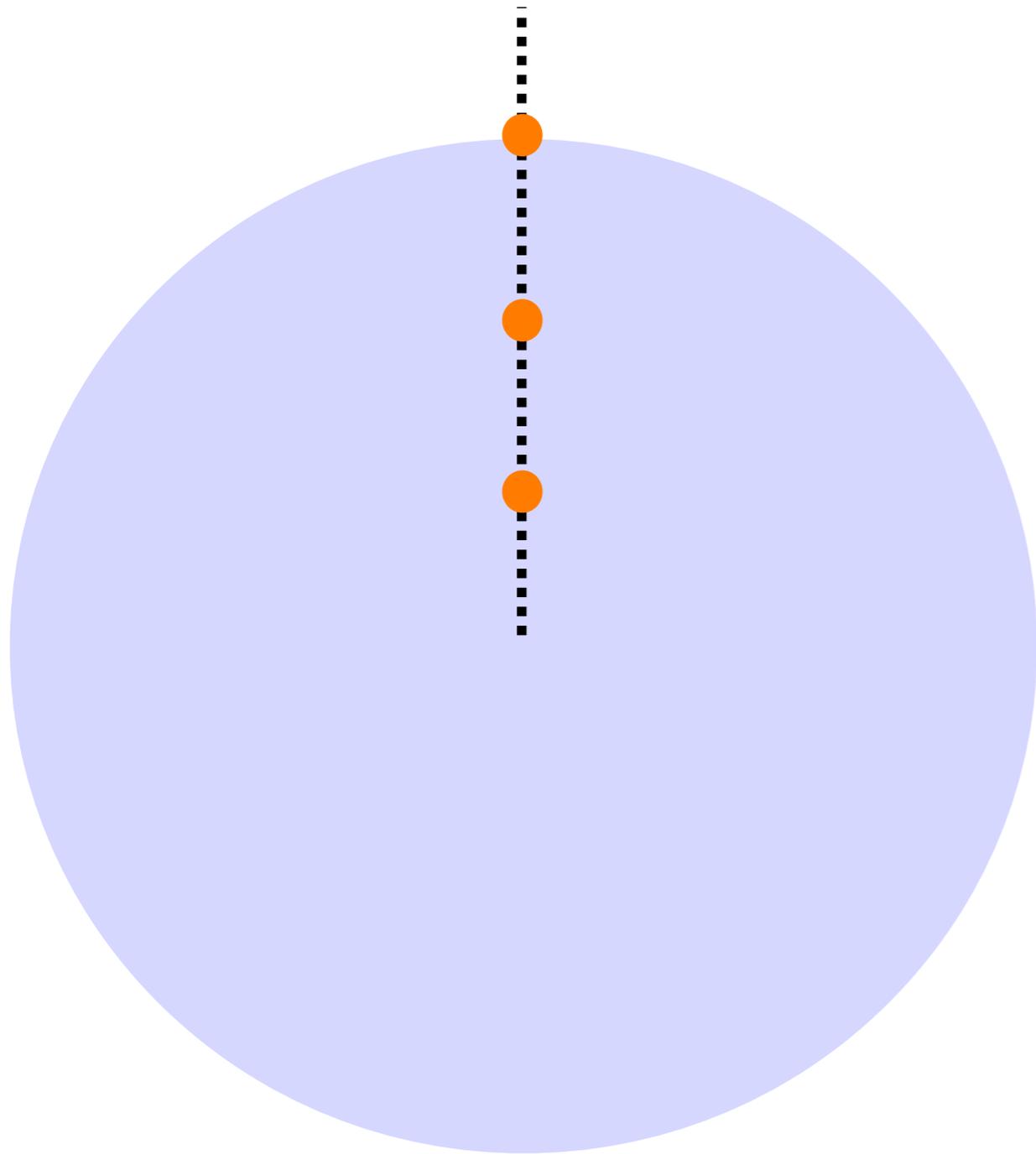
- ▶ organized pattern
- ▶ Stay in the disk (they may bob up and down)
- ▶ Orbits roughly circular

Stars in the halo and bulge orbit the Galactic nucleus randomly

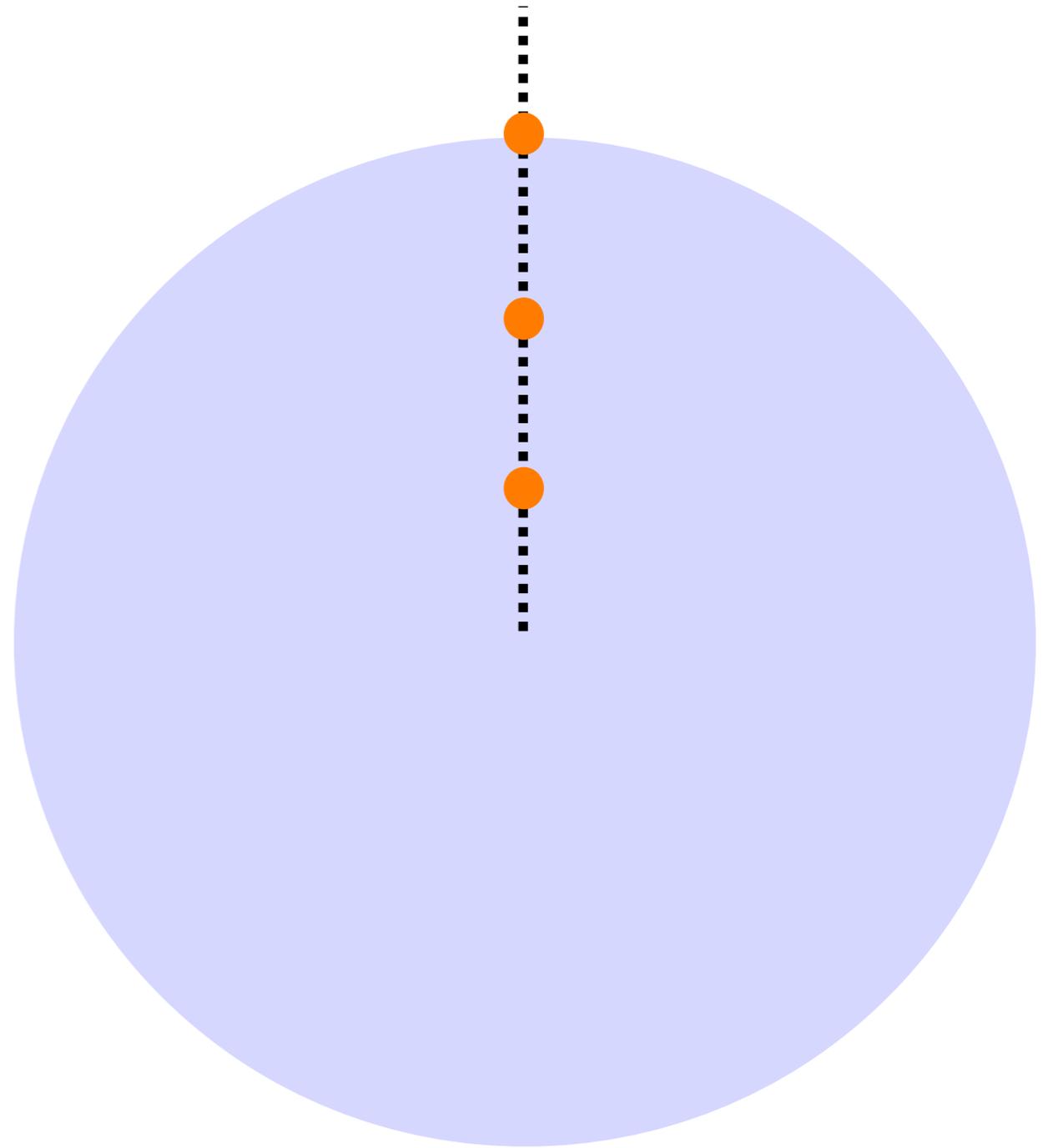
- ▶ No organization to the orbits
- ▶ Many very elliptical orbits
- ▶ collection of random orbit appears spherical



Solid vs. Differential Rotation

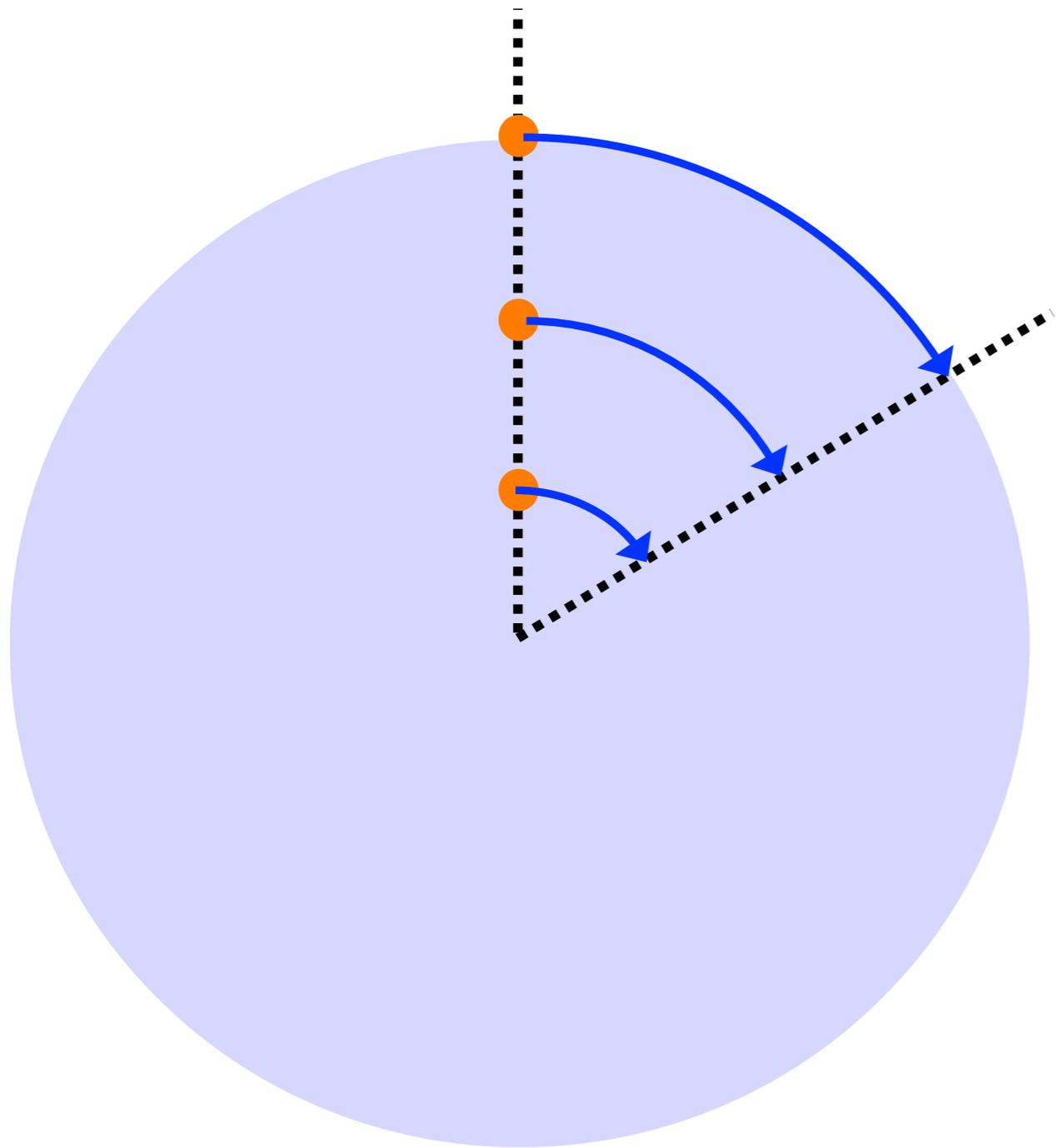


Same angular speed
(degrees per year)

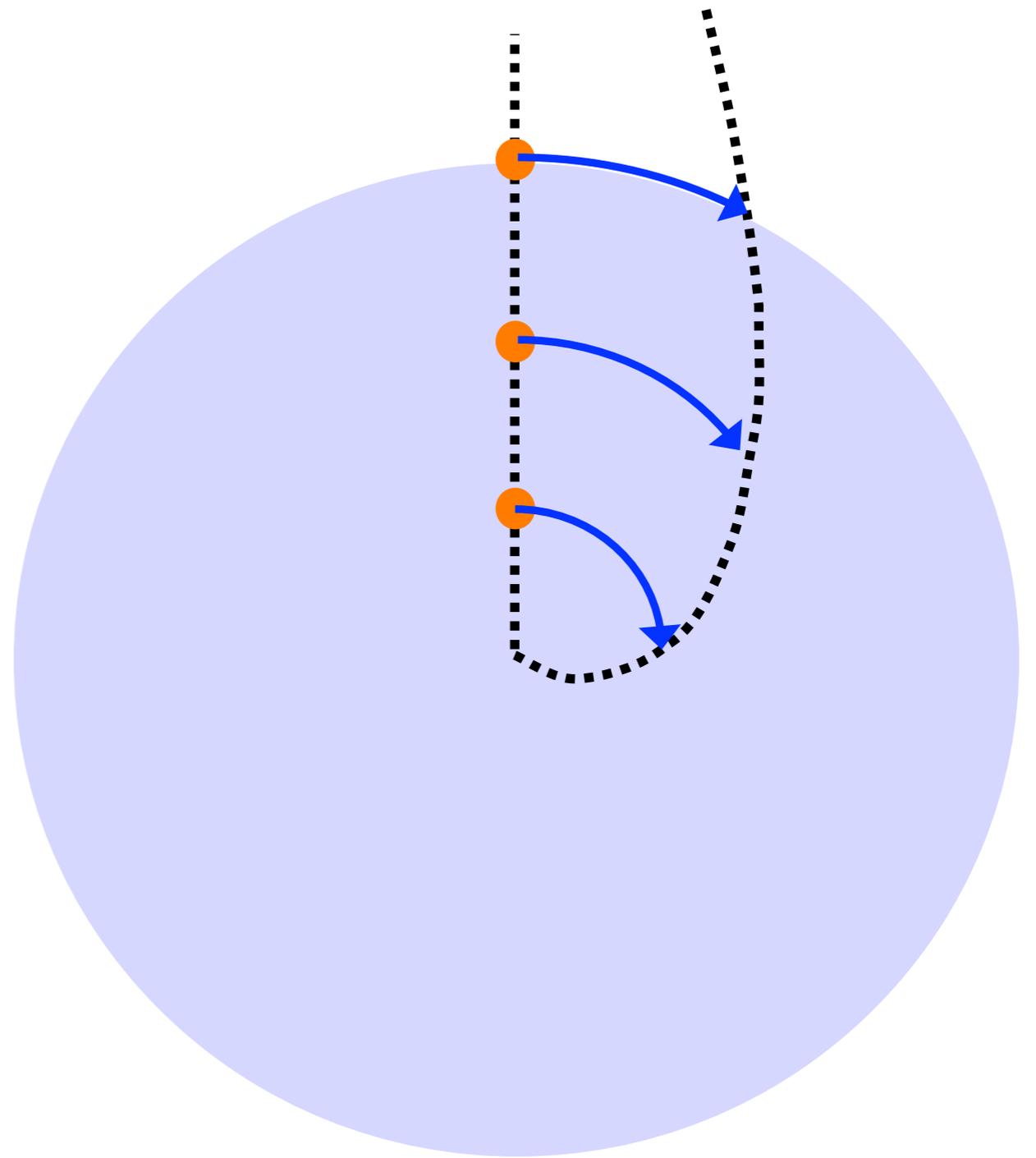


Same linear speed
(parsecs per year)

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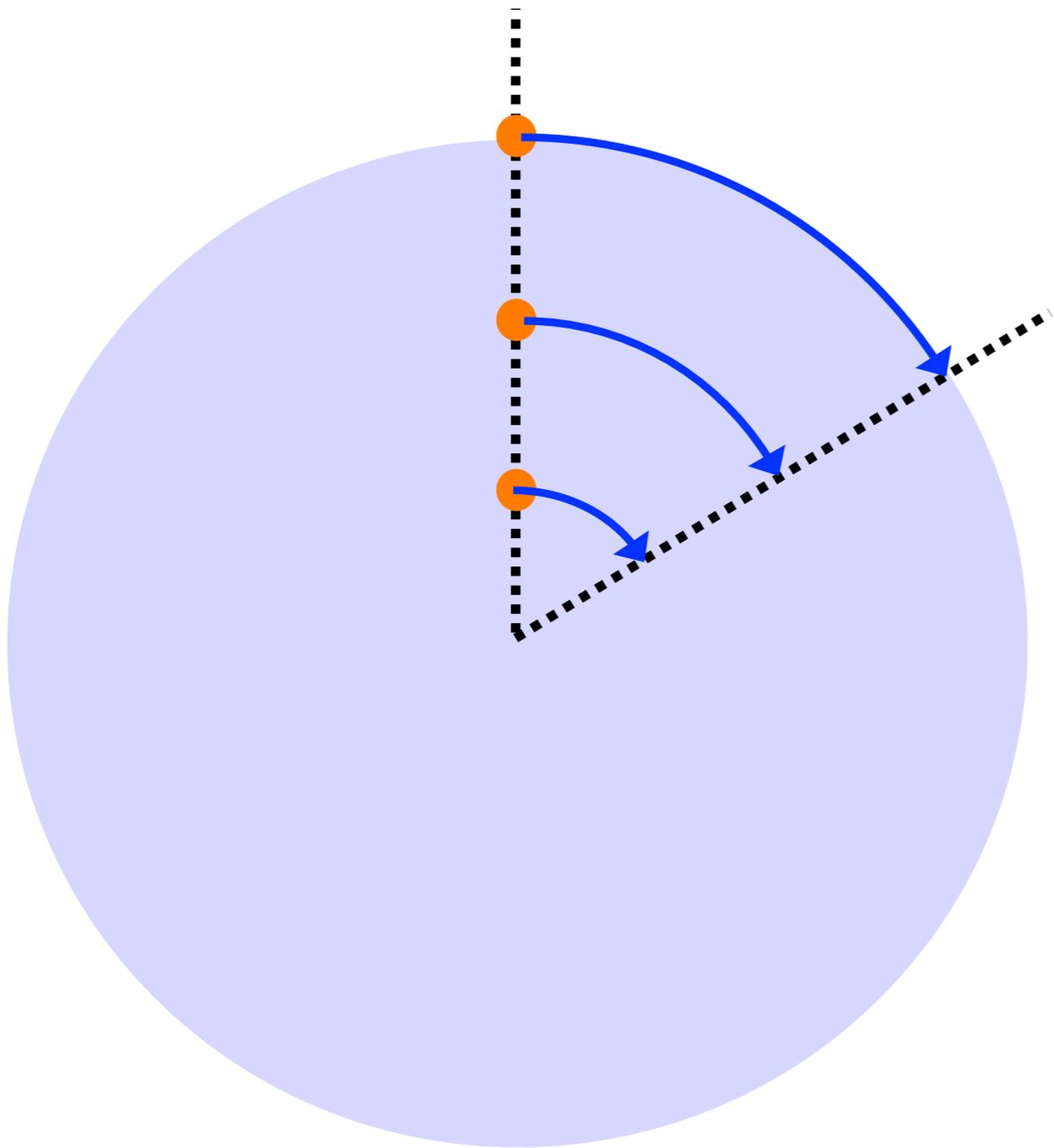


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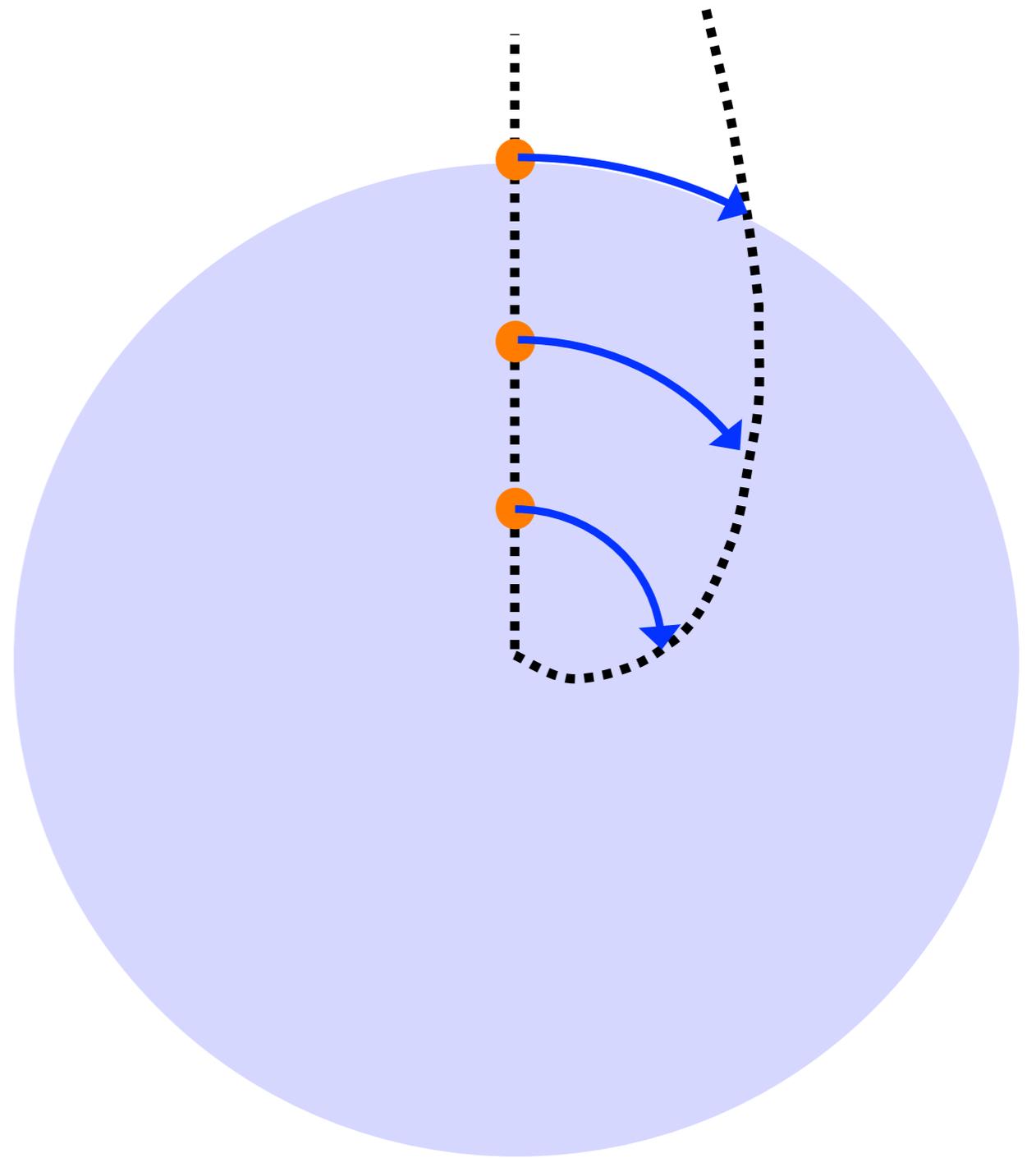


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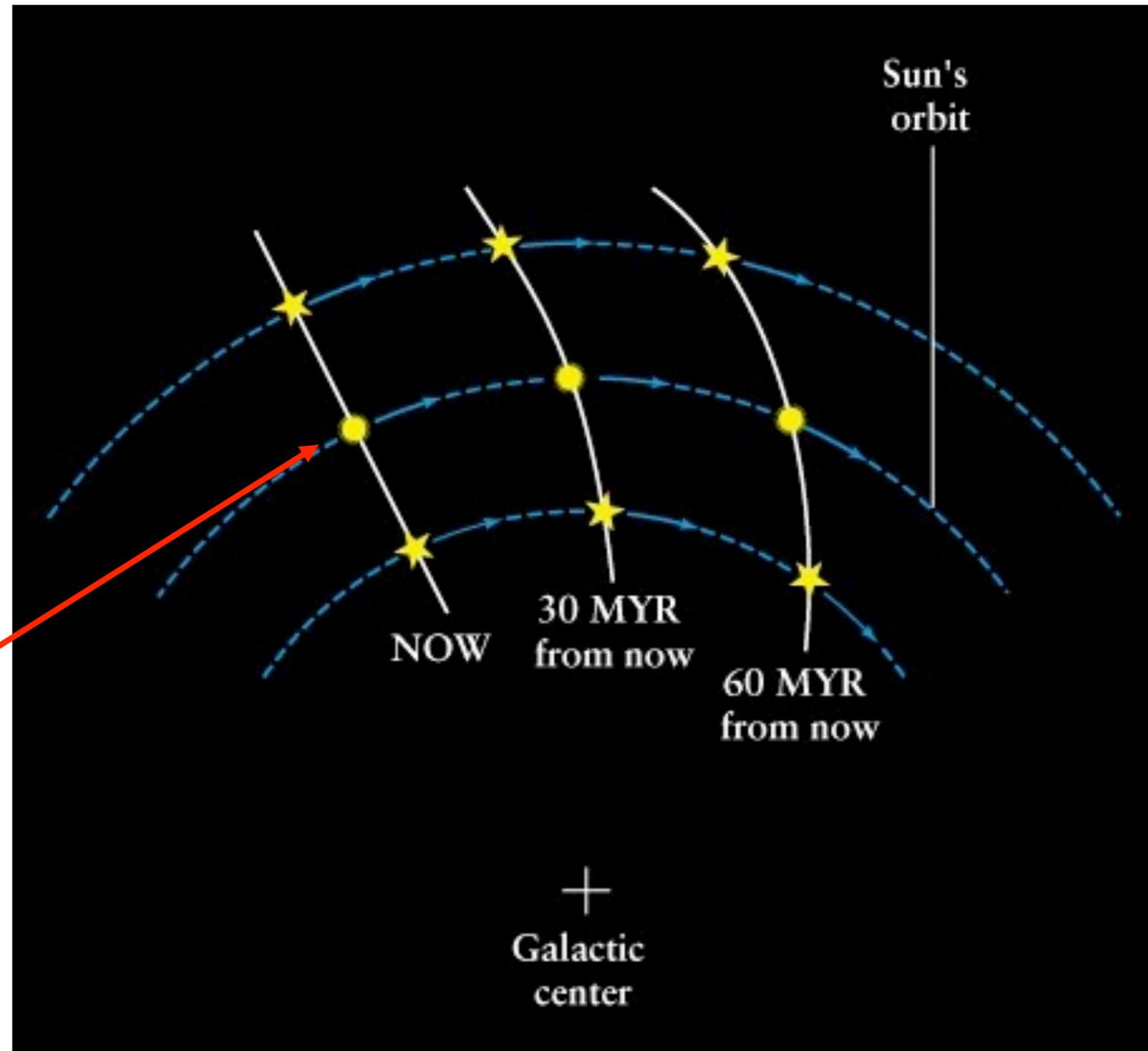


Same linear speed
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Is the Solar System Moving Too?

Yes... the whole Galaxy has differential rotation– us included



The Sun orbits at 220 km/s or about 500,000 mph– 230 million years per orbit!

Wow! That's fast!

Stop and think about it.

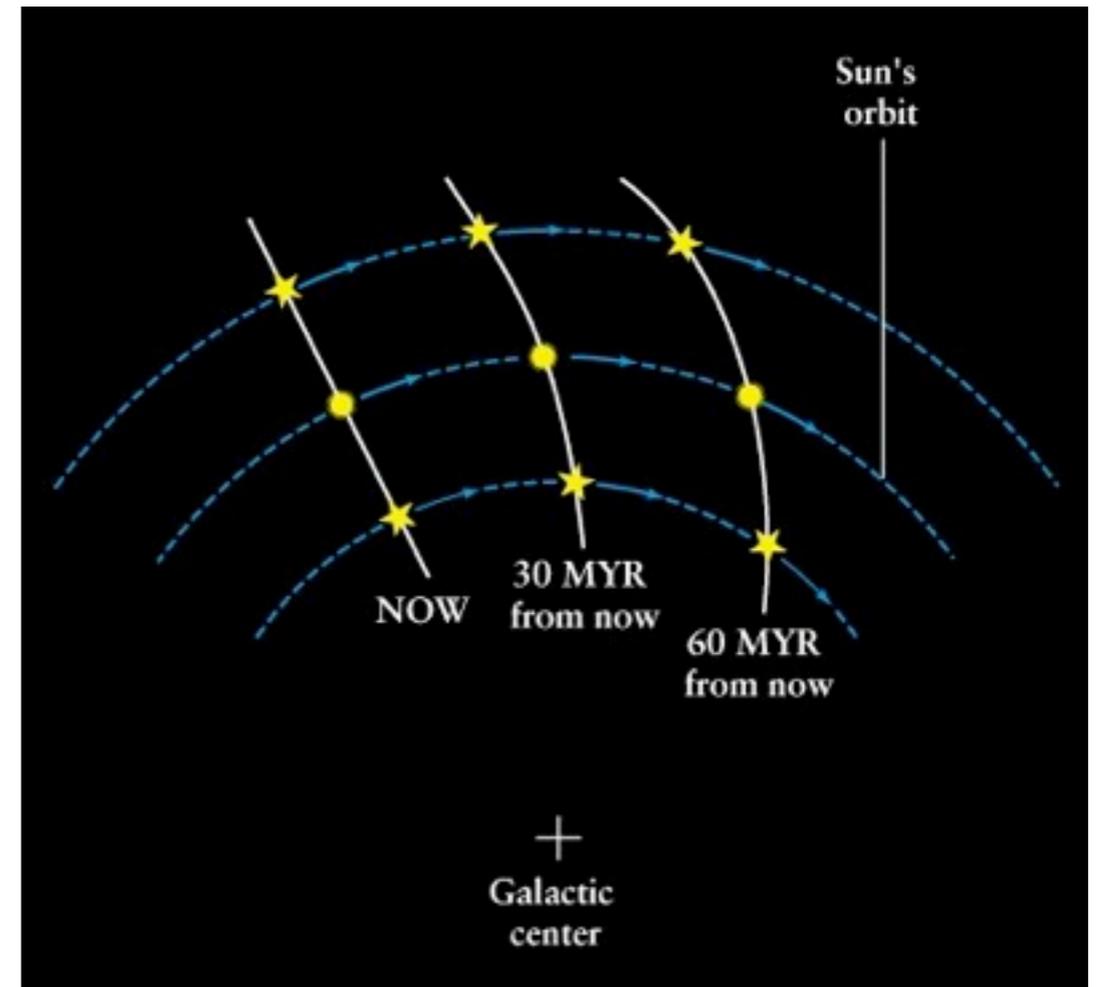
**That's traveling to Chicago
in 1 second!**

But Milky Way is big!

**Earth has only orbited 50
times!**

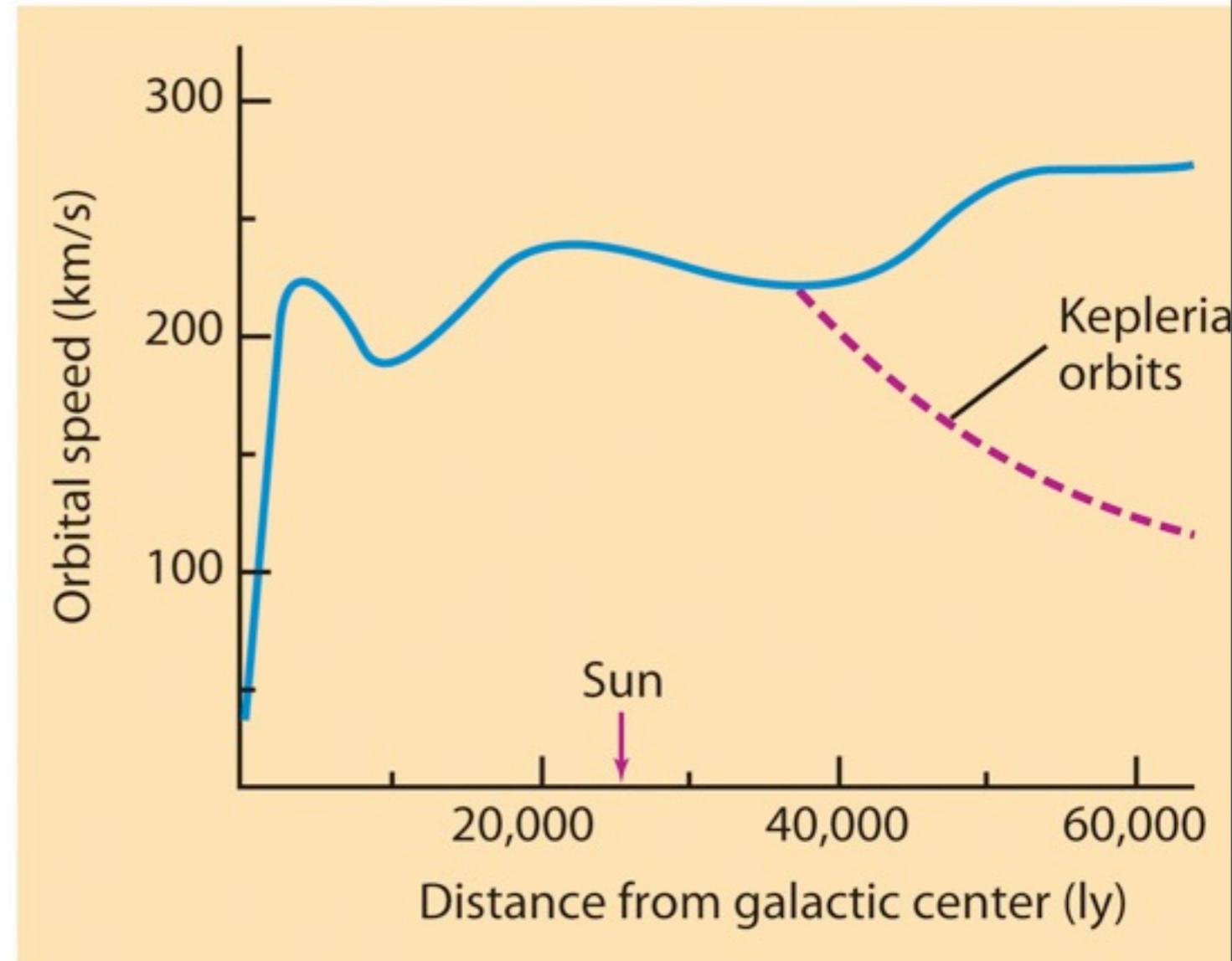
**Last time the Sun was
where at this spot in the
Galaxy, the dinosaurs were
just starting out.**

**$\frac{1}{4}$ way around, they were
extinct!**



The Rotation of the Galaxy

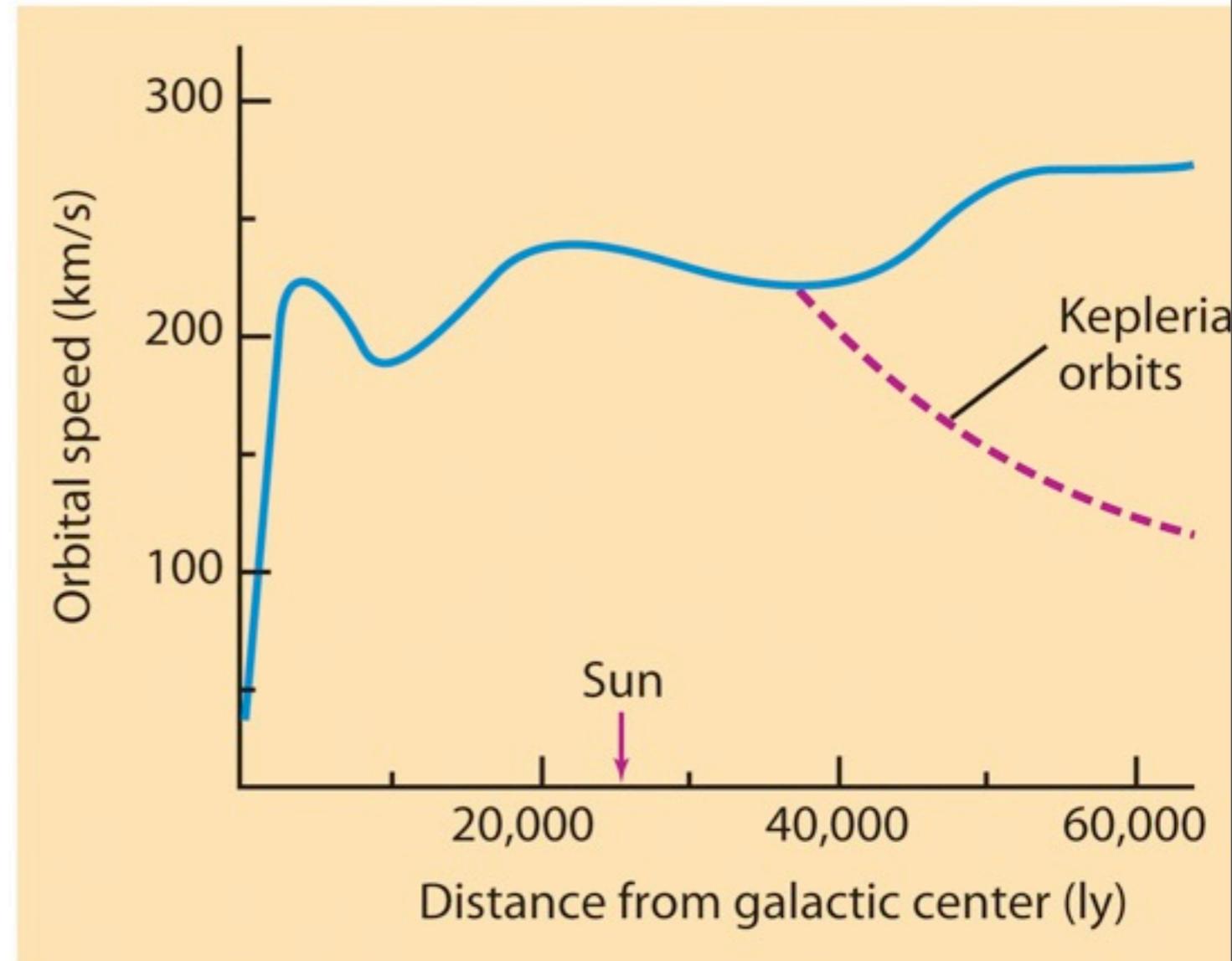
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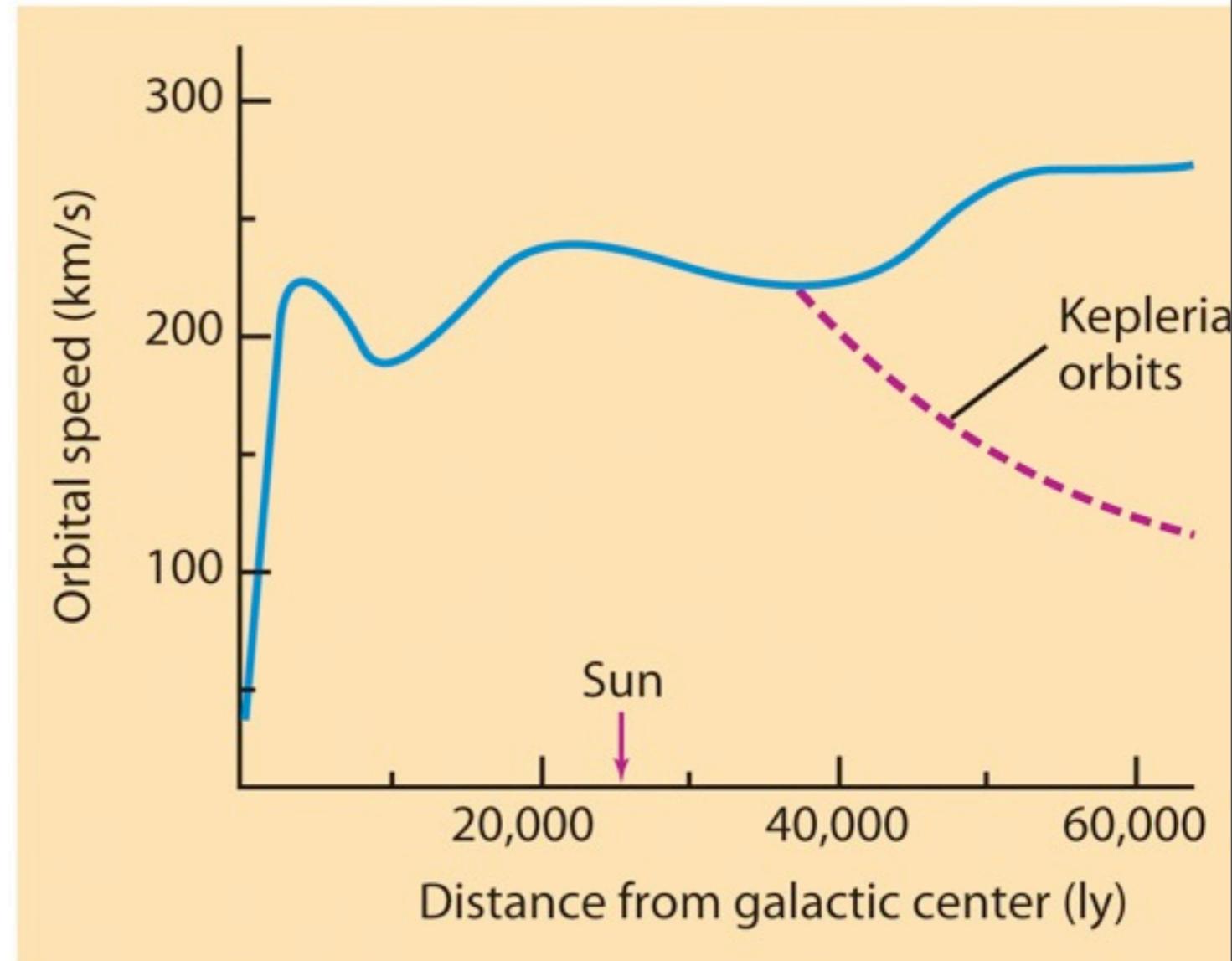
► plot: rotation speed versus distance



The Rotation of the Galaxy

Can measure rotation speeds at different distances from Milky Way center

- ▶ plot: rotation speed versus distance
- ▶ “rotation curve”

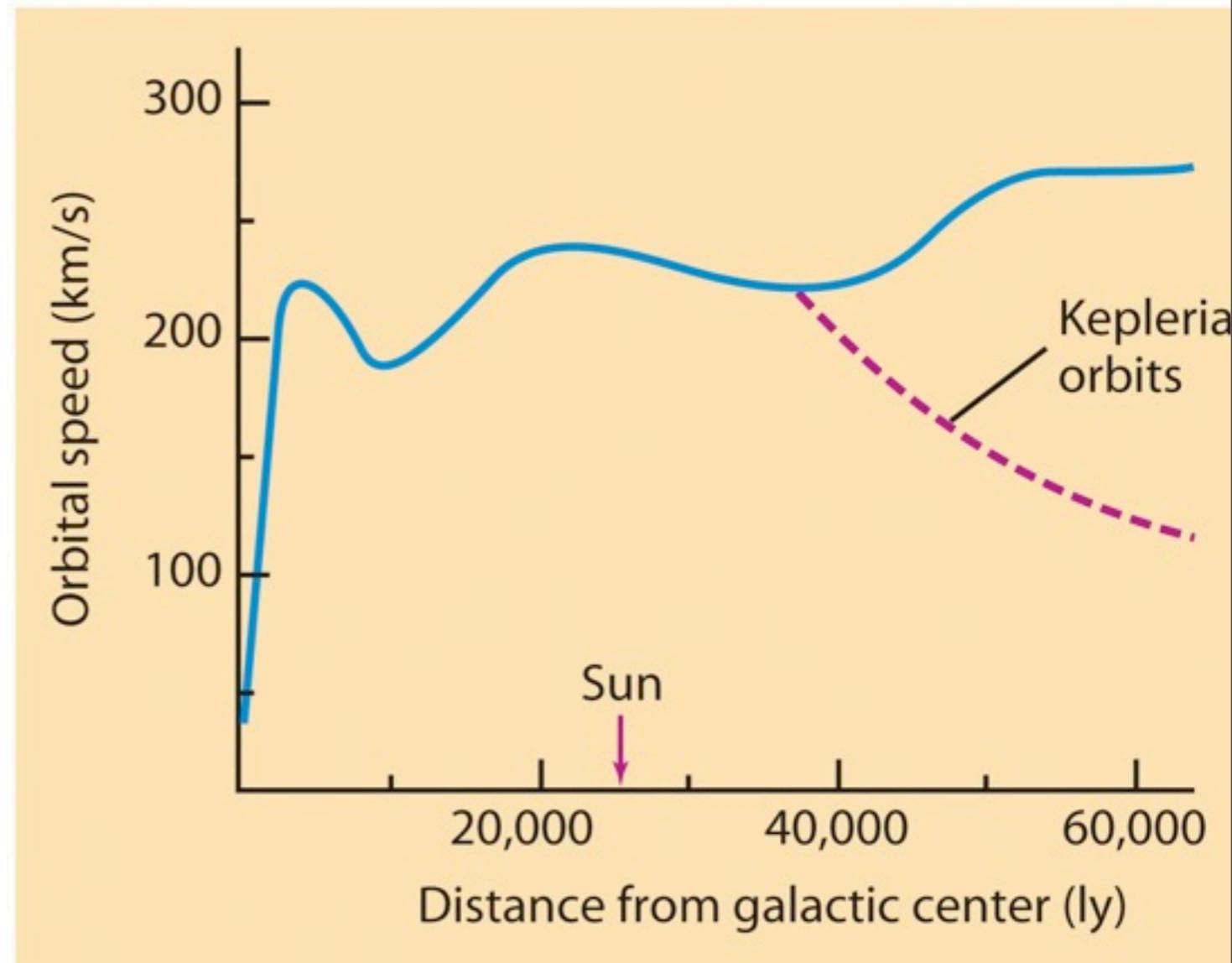


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Since we know the Sun’s speed, we can estimate how much mass **inside** our orbit.



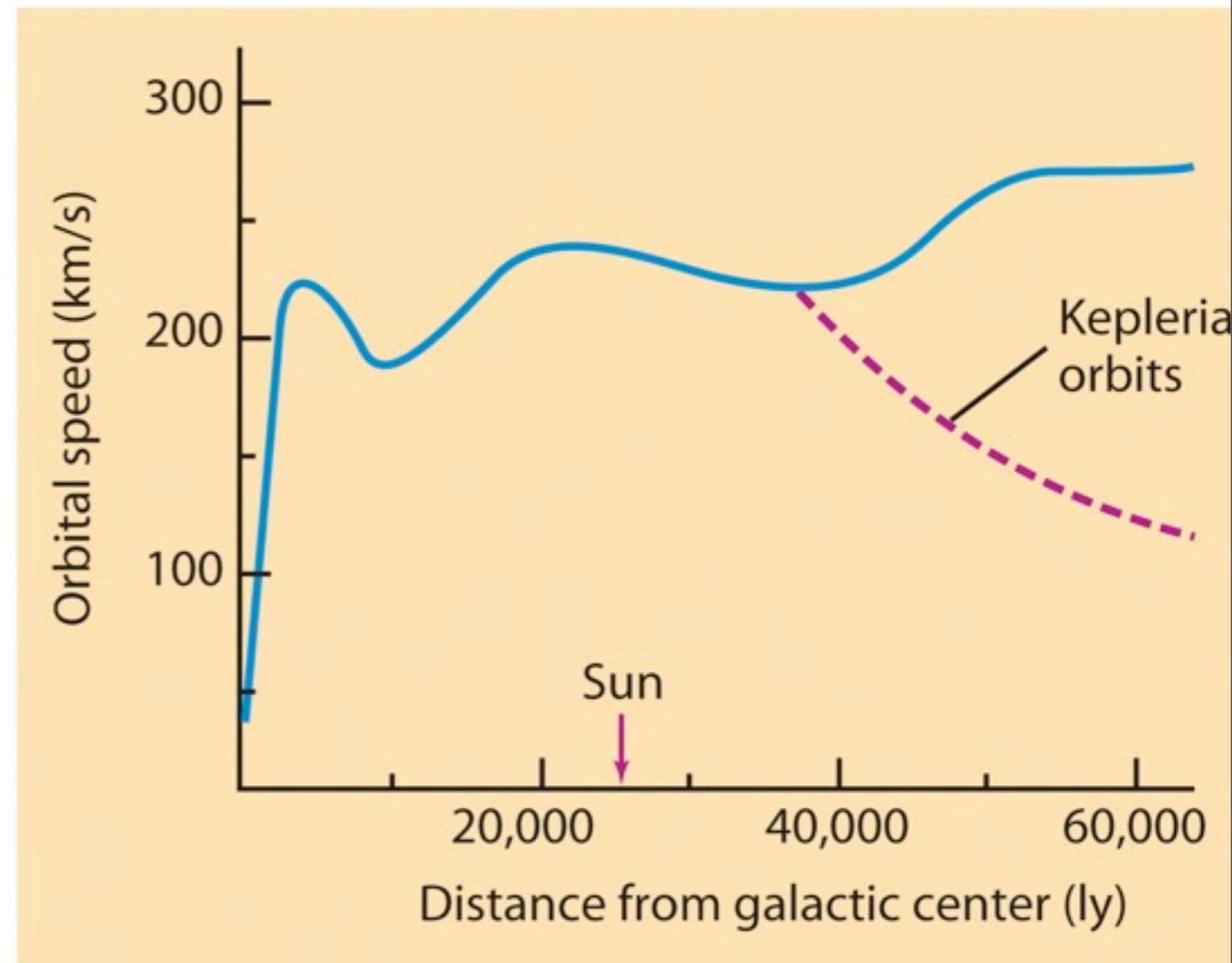
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- ▶ since gravity controls orbits



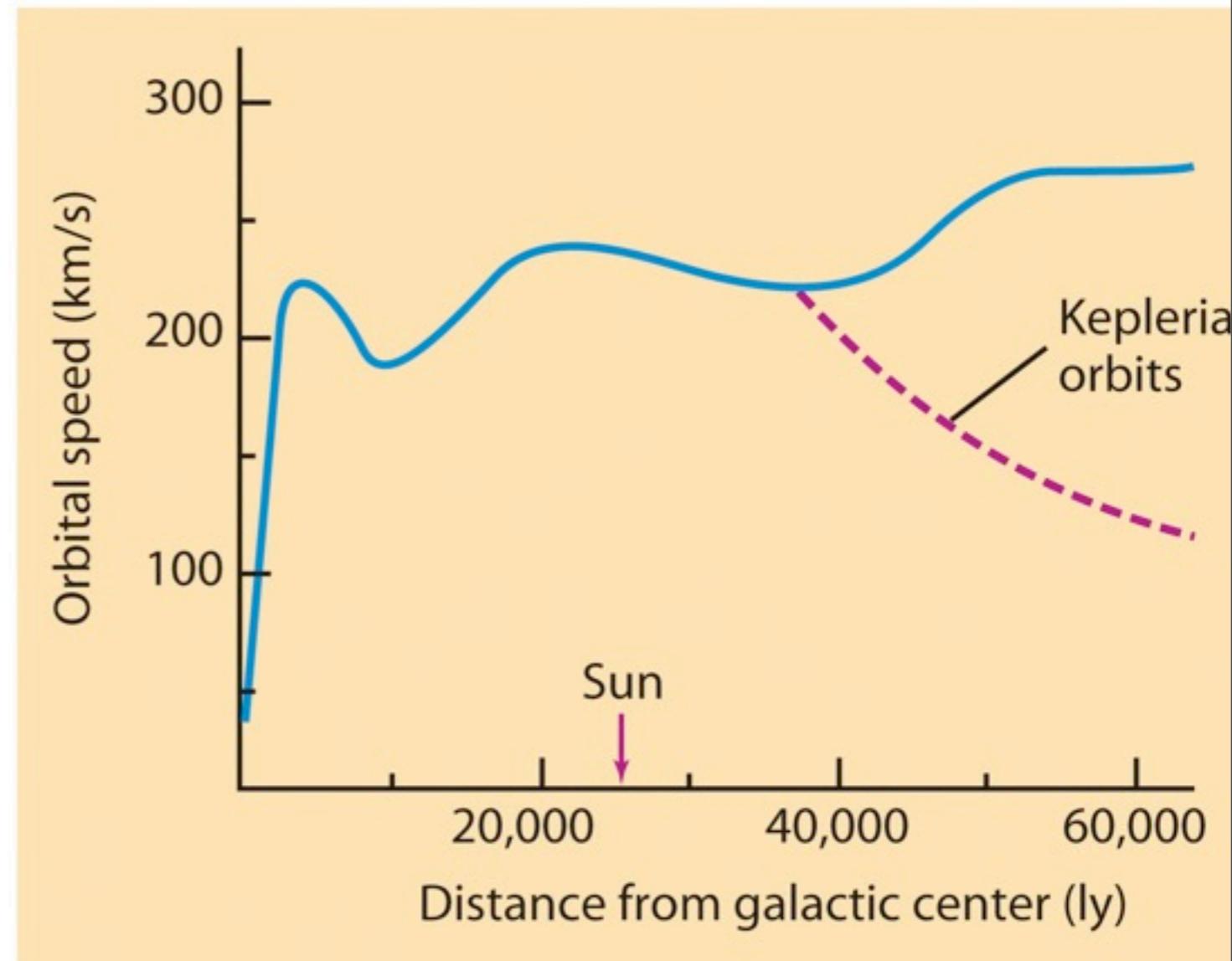
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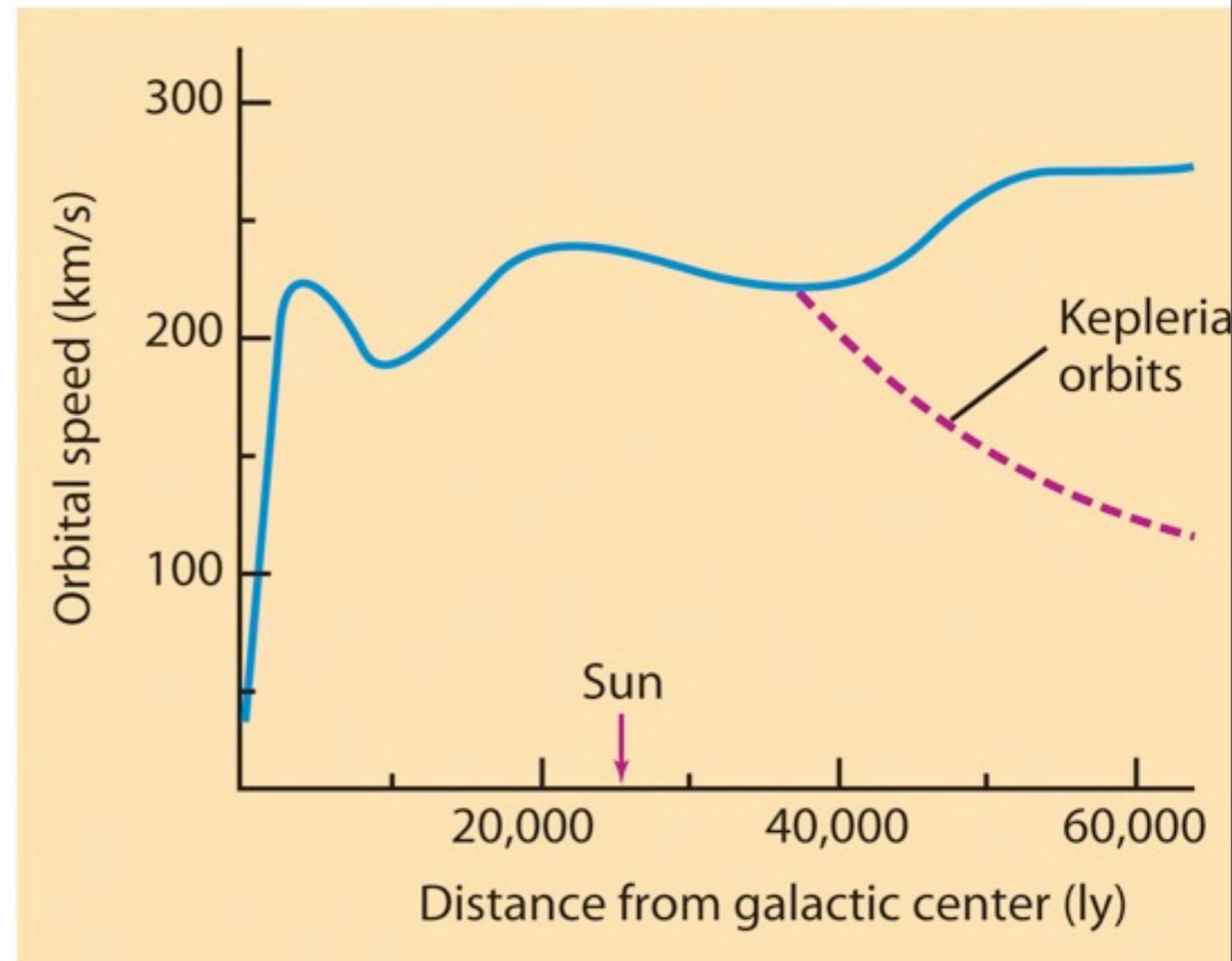
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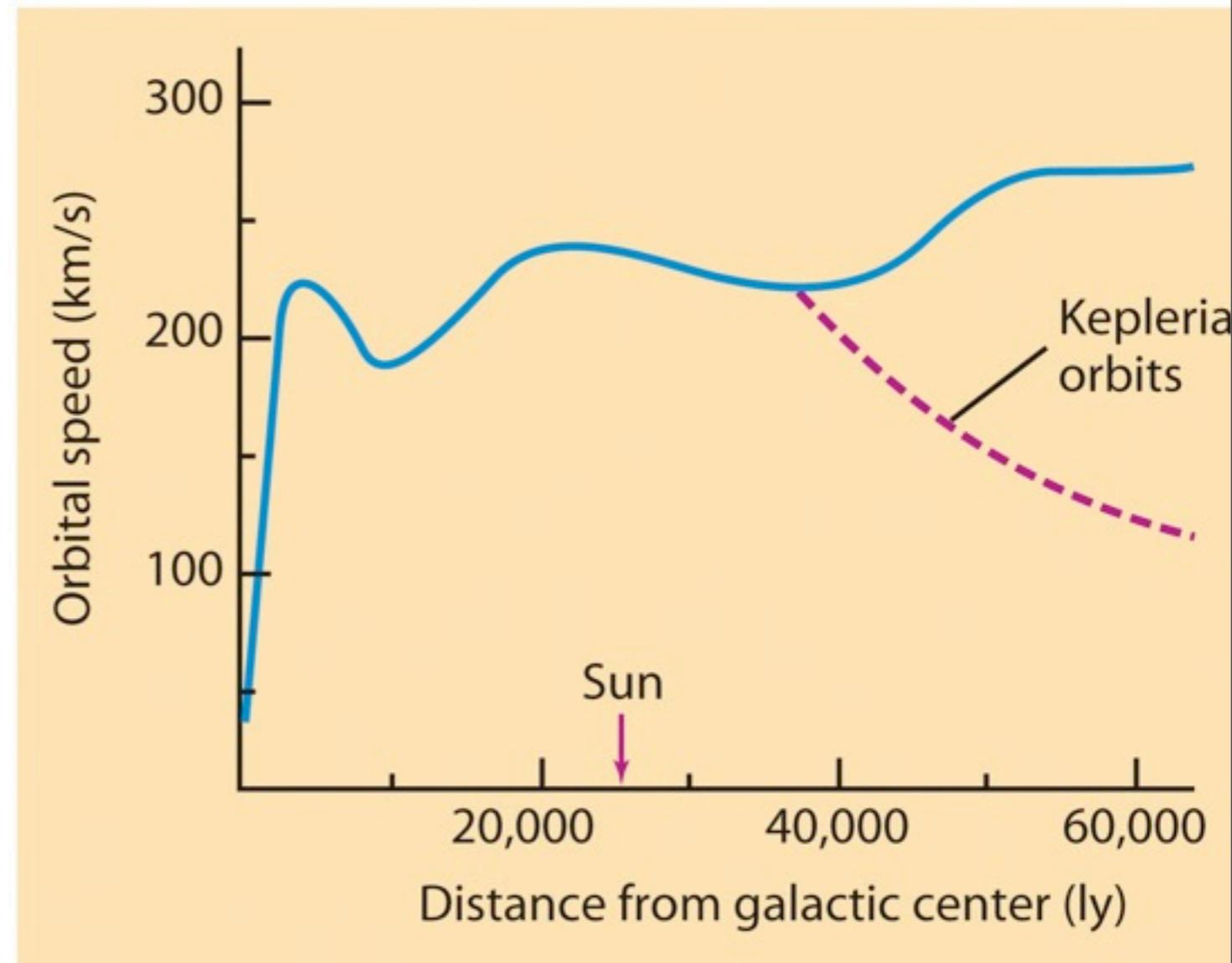
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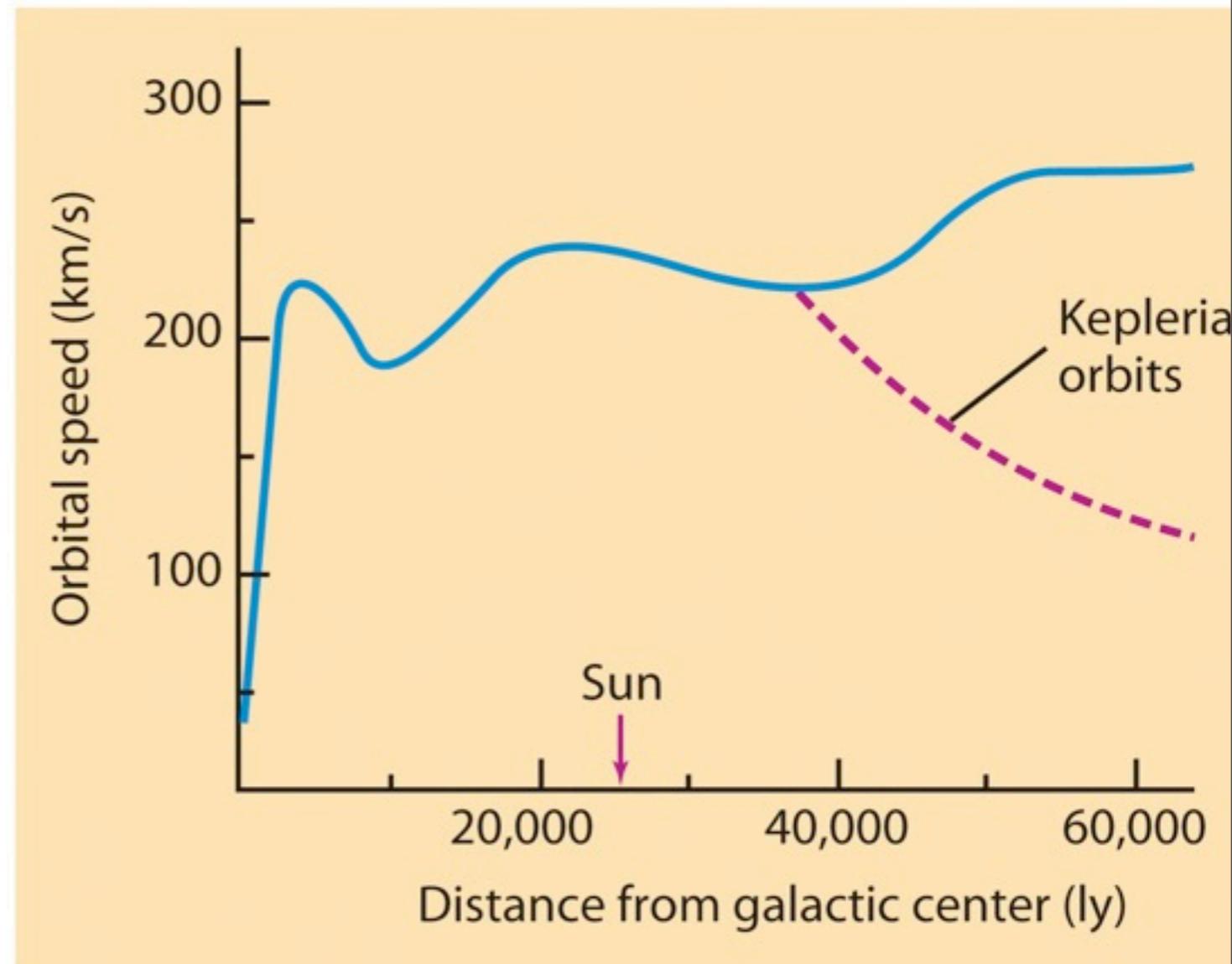
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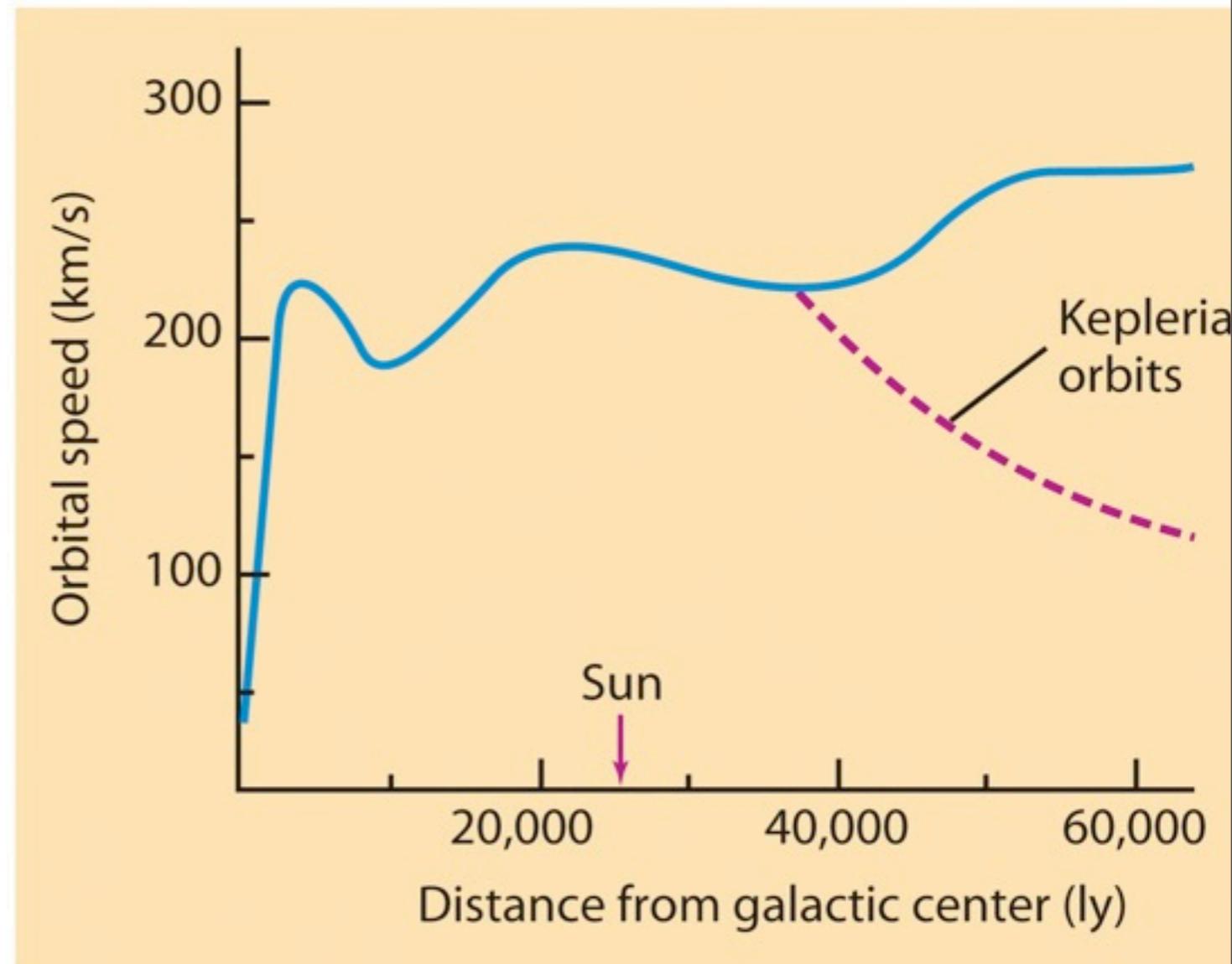
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speed is constant from 5000 lyrs out.



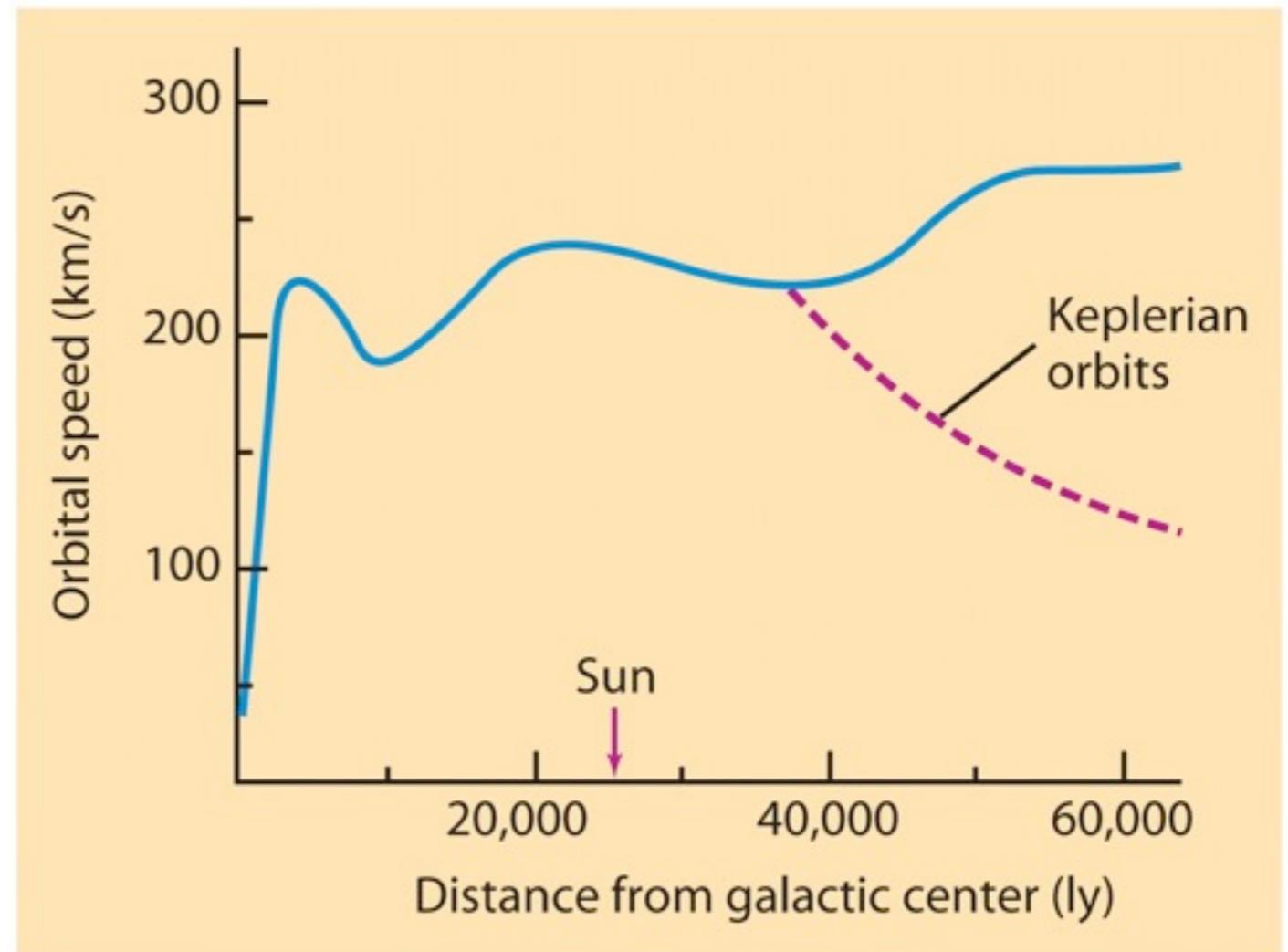
Whaa?

Something weird.

Rotation speed **does not drop off** in the outer Galaxy

- ▶ rotation curve is “flat”
- ▶ yet in the outer galaxy there is a **huge dropoff** in stars, gas, and dust

Instead: rotation speed remains constant, or even slightly increasing!



Rotation Curve Shows Hidden Mass

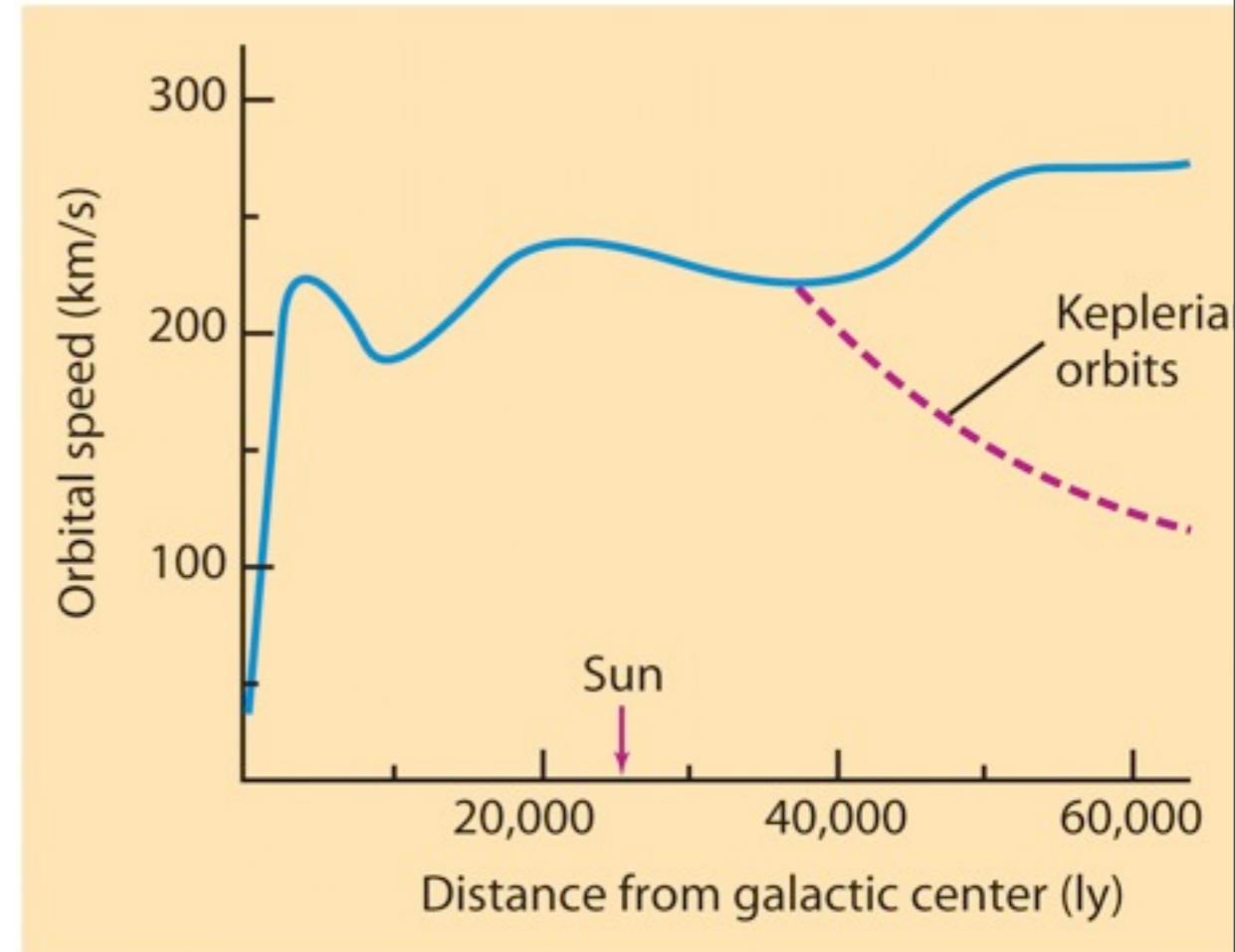
In the Solar System:

- ▶ >99% of mass is in Sun
- ▶ the farther away from the mass, the slower something orbits.

Compare Pluto and Mercury's orbital speed.

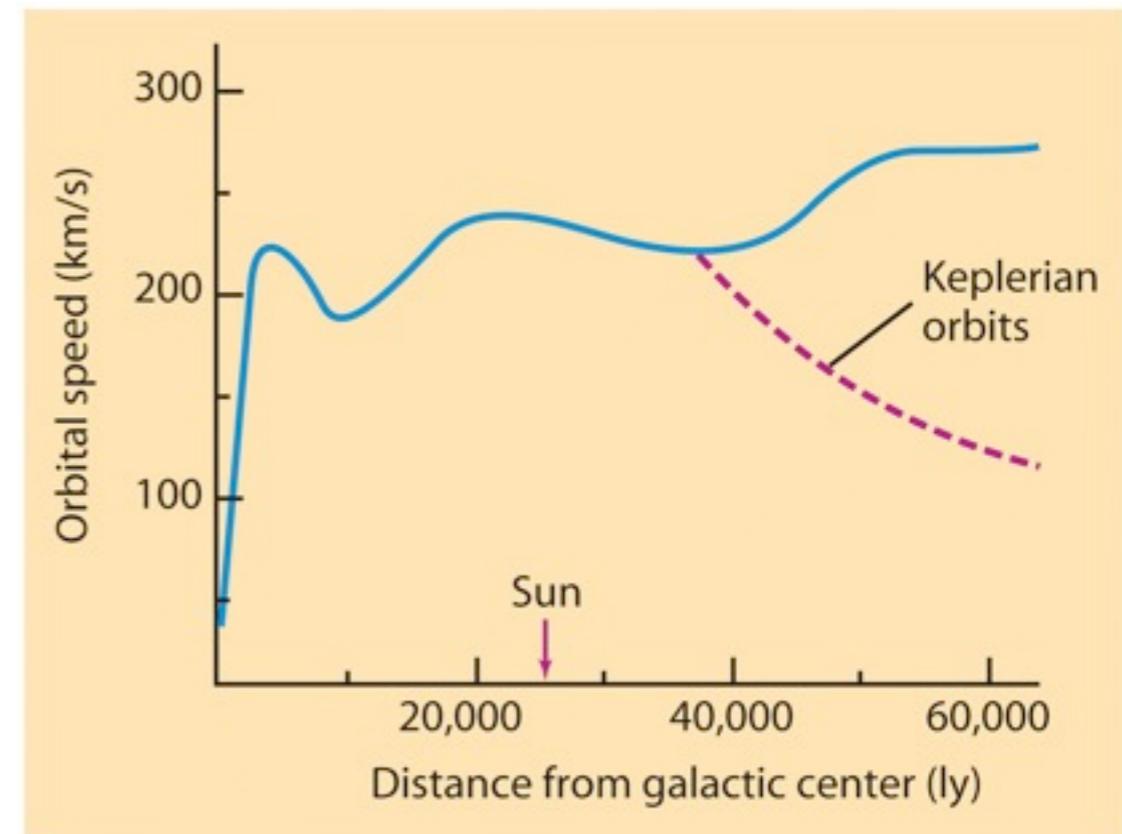
But in the Galaxy:

- ▶ orbit speed of stars in the outer Galaxy actually increases or is constant with distance from the center!



Rotation Curve Shows Hidden Mass

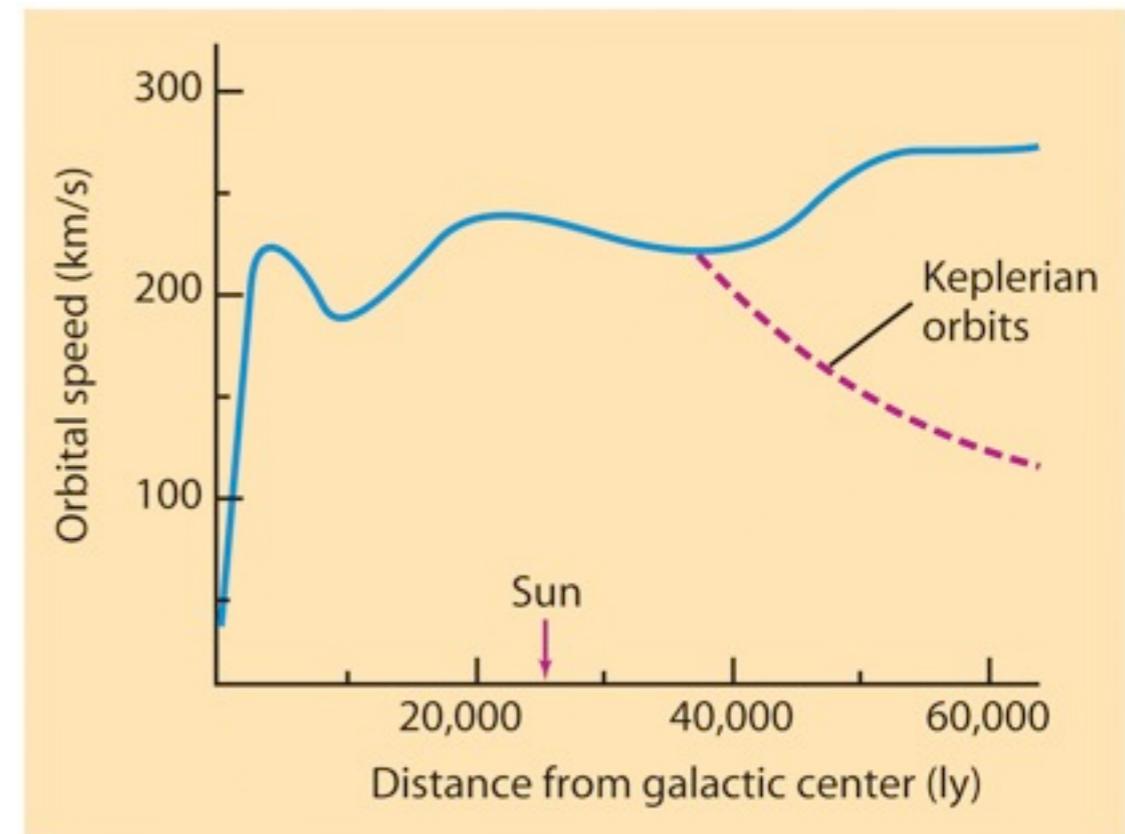
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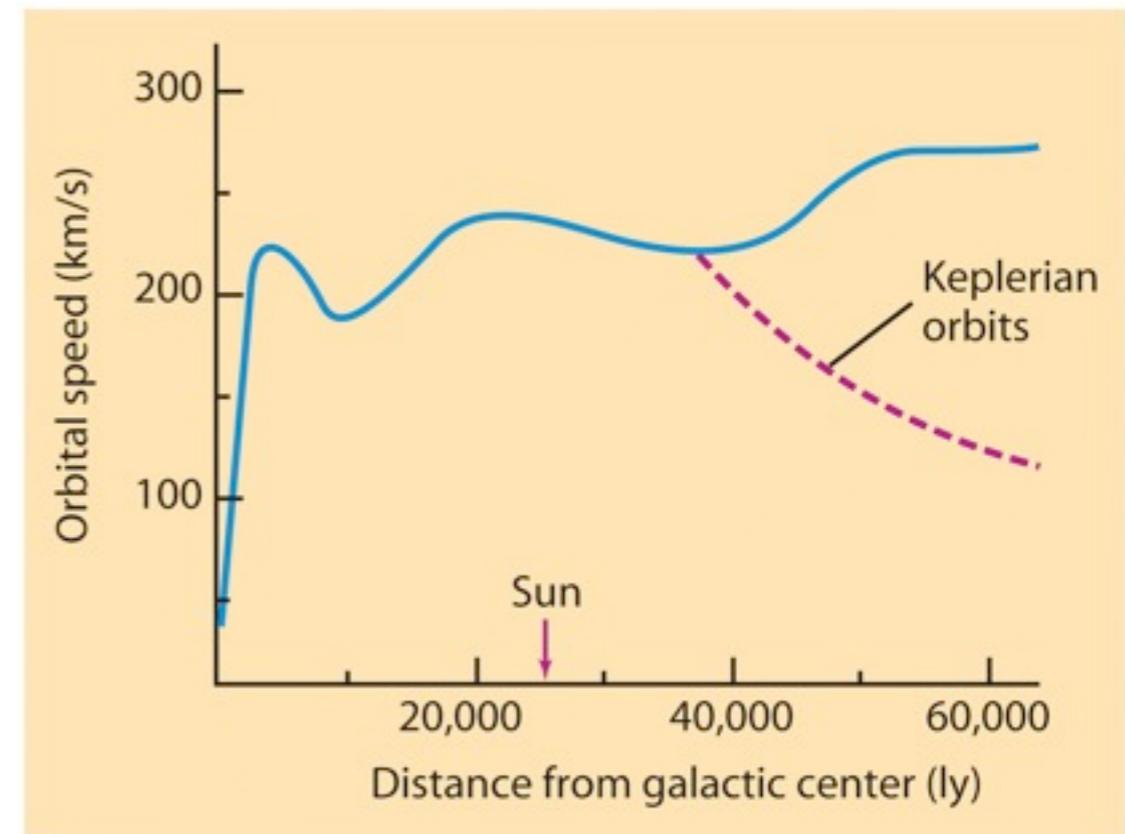
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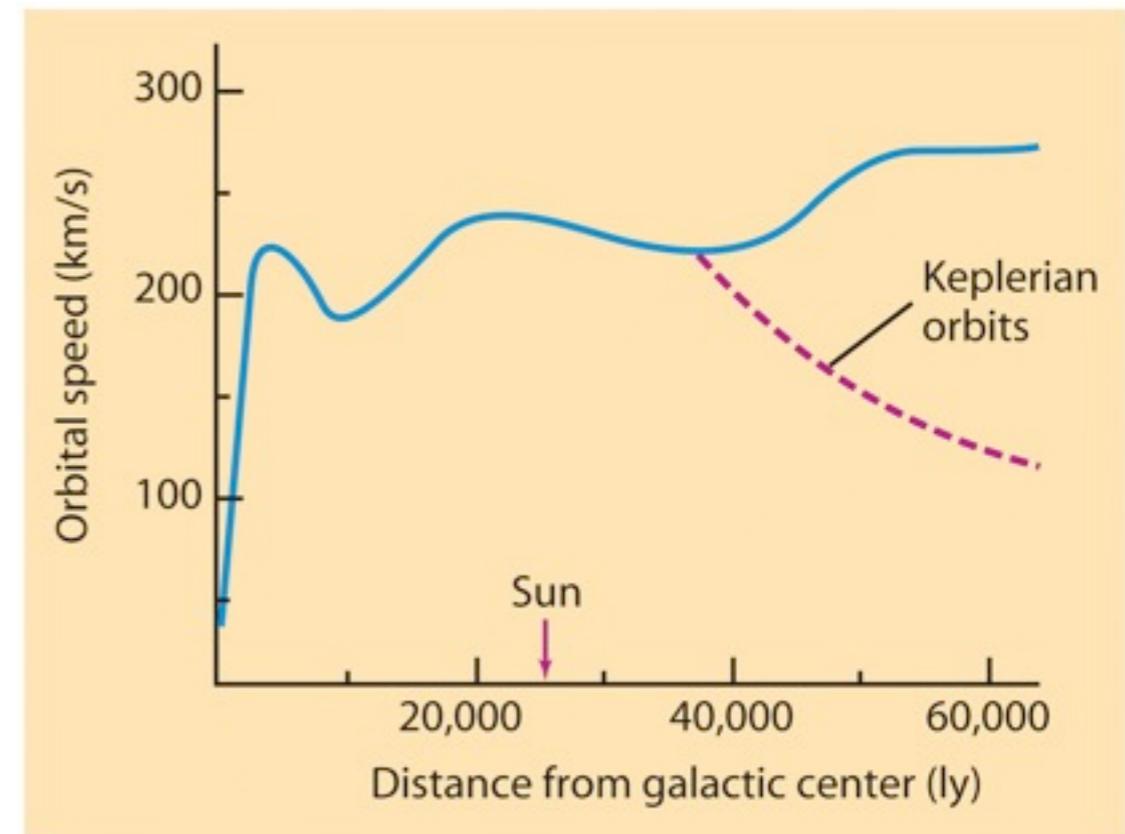


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So: The mass in the outer part of the Galaxy must be something **dark**.



Dark Matter

MW mass estimate (Kepler's laws):

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Revised view of Milky Way structure:

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- ▶ total mass $M_{MW} = 10^{12} M_{sun}$
- ▶ but stars & gas: $M_{stars+gas} = 10^{11} M_{sun}$
- ▶ stars & gas (visible stuff) only **10% of total!**

Forced to conclude:

- ▶ a large amount of mass is unseen!
- ▶ most (90%!) of Galaxy mass is in the form of unseen **dark matter!**

Revised view of Milky Way structure:

- ▶ **disk**: most stars, all gas/dust $R_{disk} \approx 15,000$ pc

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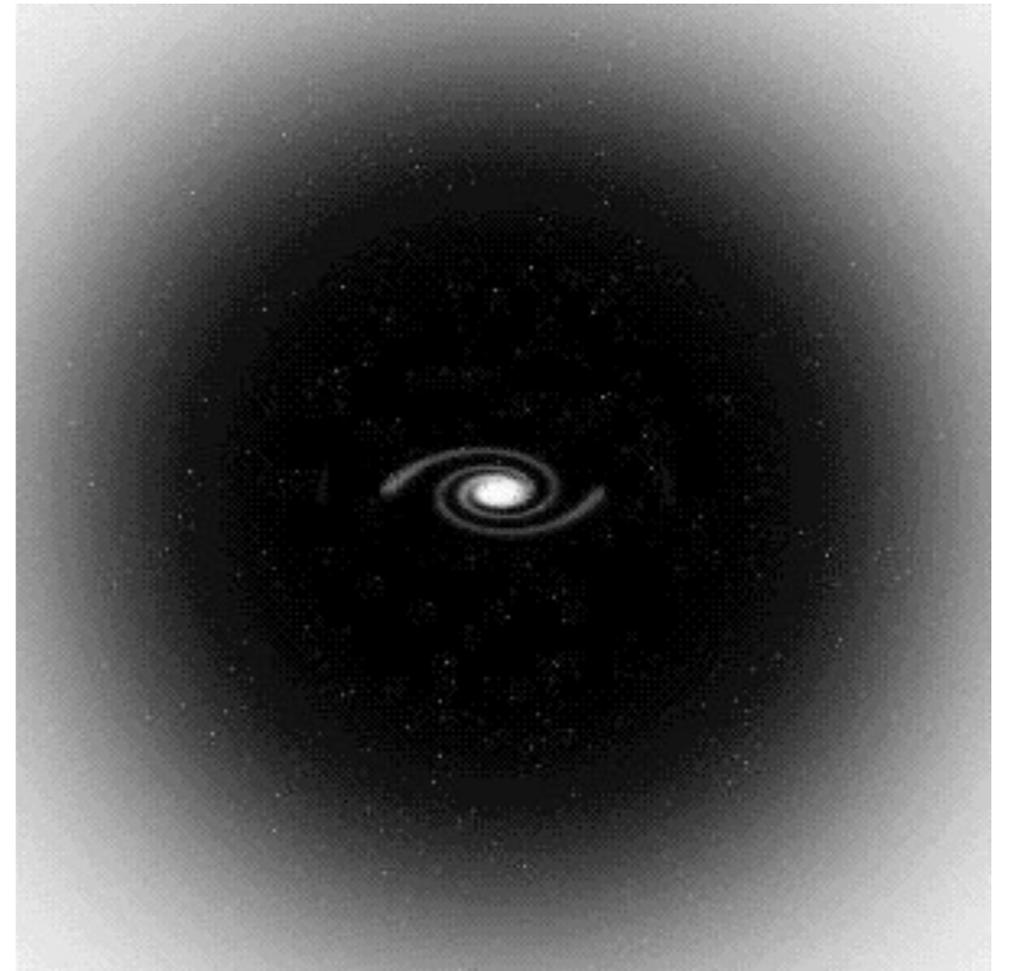
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Milky Way much more massive, larger, than meets the eye!

Dark Matter

The dark matter in the Galaxy is in greatly extended halo

- ▶ **Up to 90% of the Galaxy's mass is dark matter!**
- ▶ **Galaxy may have over a trillion solar masses total!**



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One of biggest questions in science today!

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Will return to dark matter soon....

“Spiral Nebulae”

All stars you can see by eye or with small telescopes are in Milky Way

- ▶ as are stellar remains:
 - white dwarfs, neutron stars, black holes
 - and planetary nebulae

But by 1800's, other dim, diffuse “nebulae” seen

- ▶ with spiral patterns
- ▶ Spiral structures catalogued mid-1800s by Lord Rosse (Ireland)

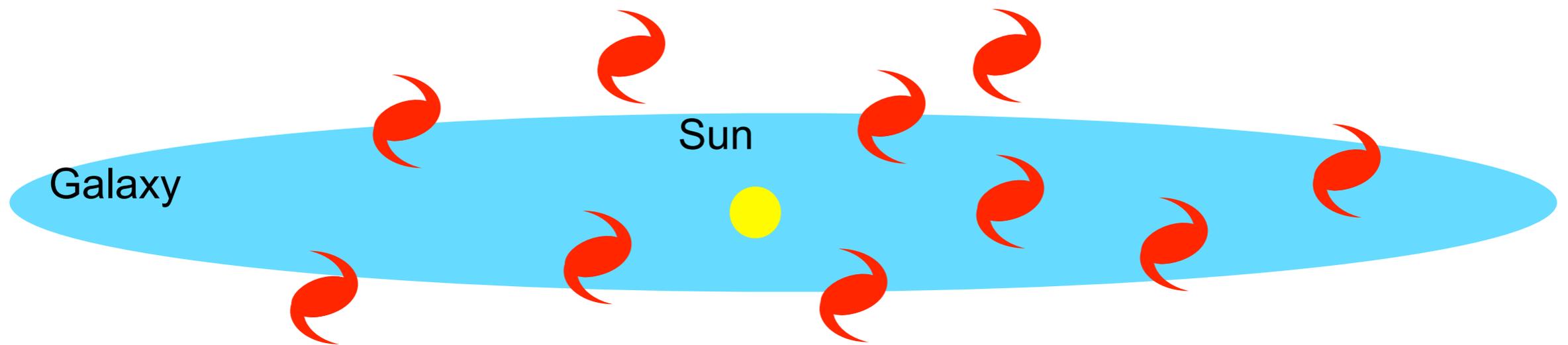


Rosse's M51 sketch



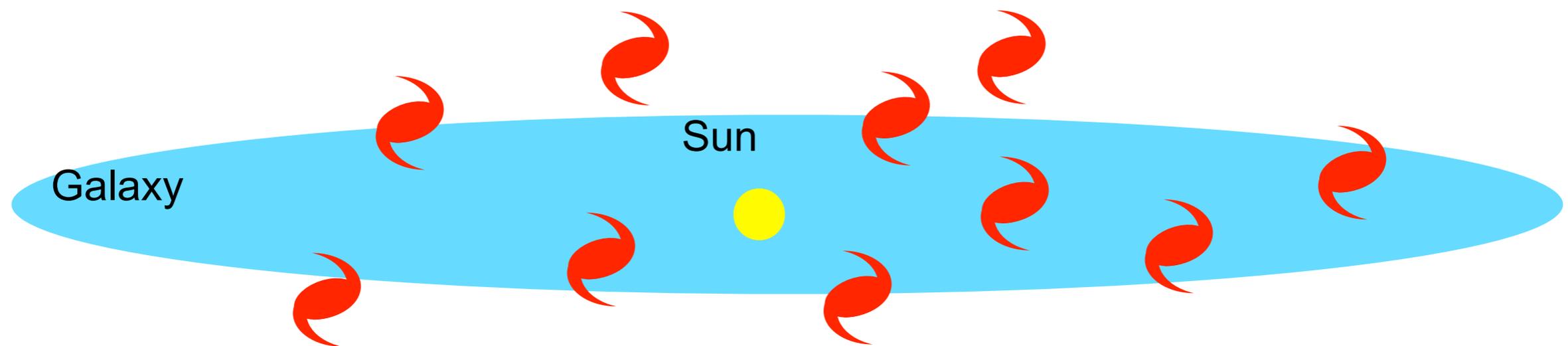
“Leviathan”
1.8 m telescope

Those weird Spiral Nebulae?



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Dim, diffuse, “interstellar” nebulae with spiral structure were seen in the 17th century.

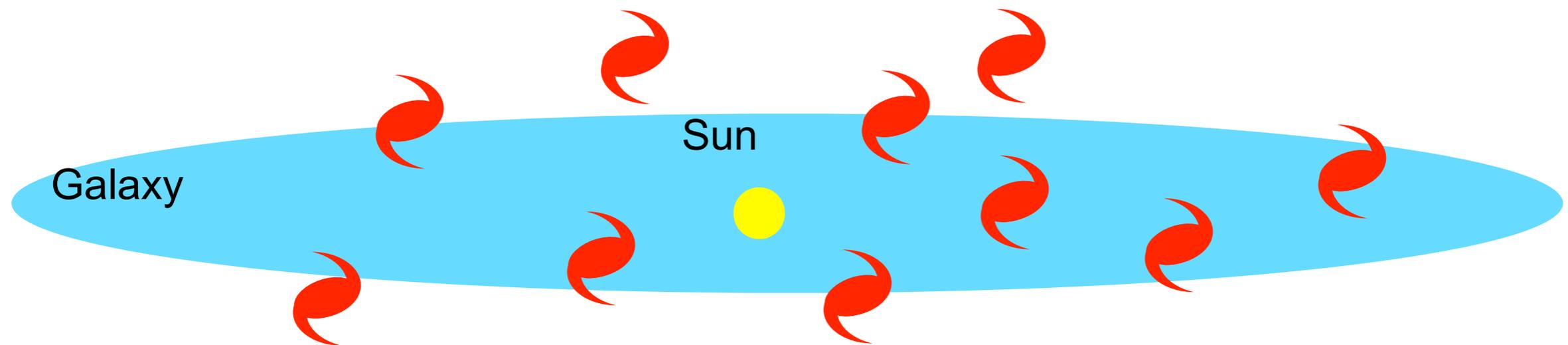


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Some disagreement on what they were.

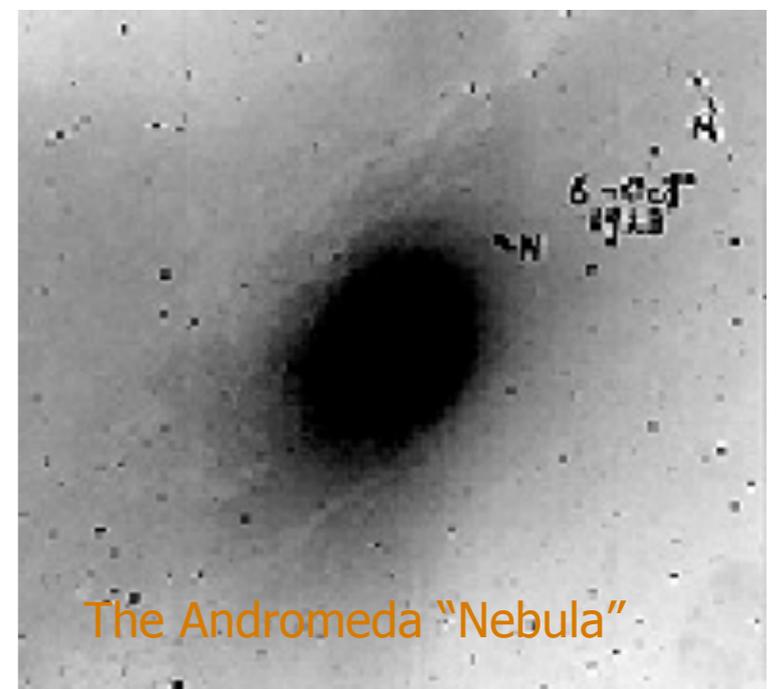
- ▶ Kant: Our galaxy is a spiral “island universe” and the other spiral nebulae are the same and far away
- ▶ Herschel and others: Milky Way is all there is in the Universe, and the spiral nebulae are nearby. More prevalent idea.



Edwin Hubble



Hubble at Mt. Wilson
Observatory



The Andromeda "Nebula"

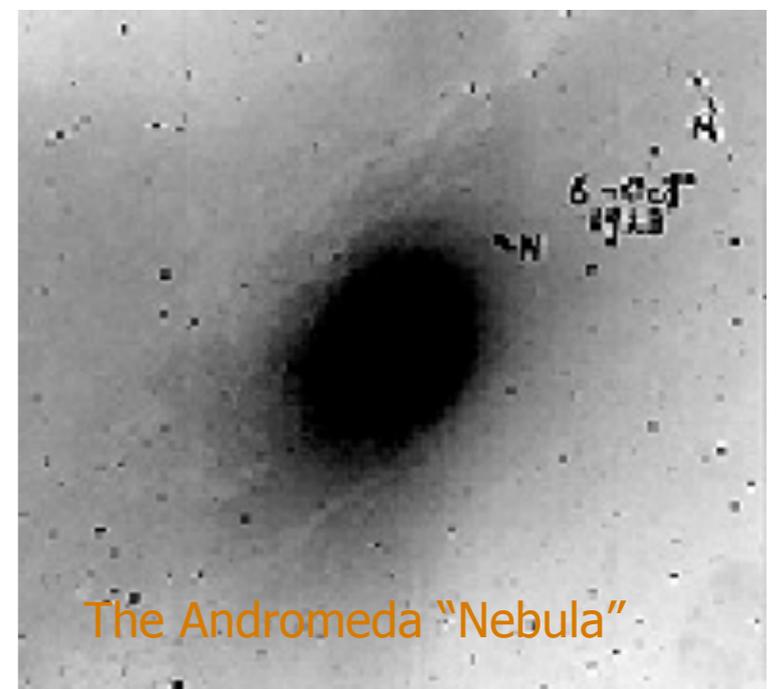
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Finally solved, as it often is in astronomy, with a BIGGER telescope!

- ▶ The old 100 inch trick!



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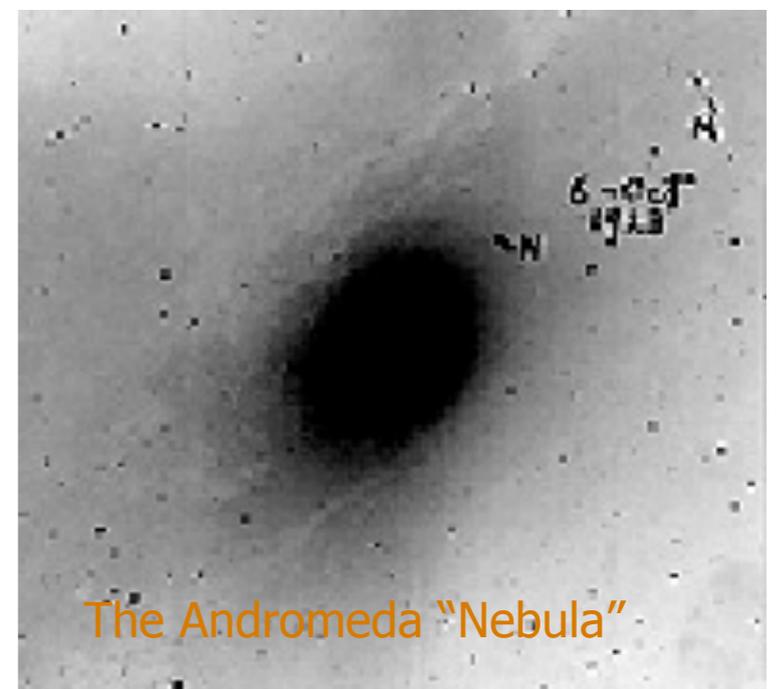
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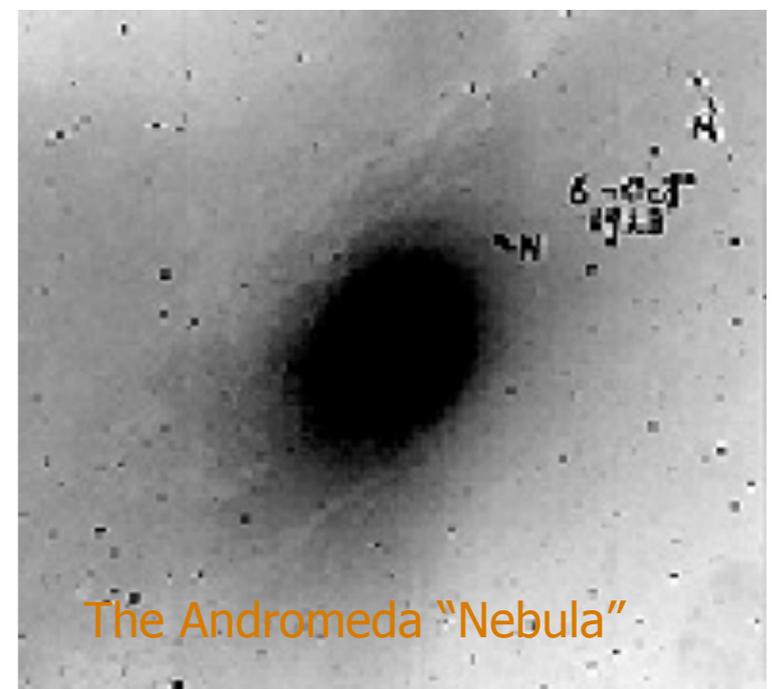
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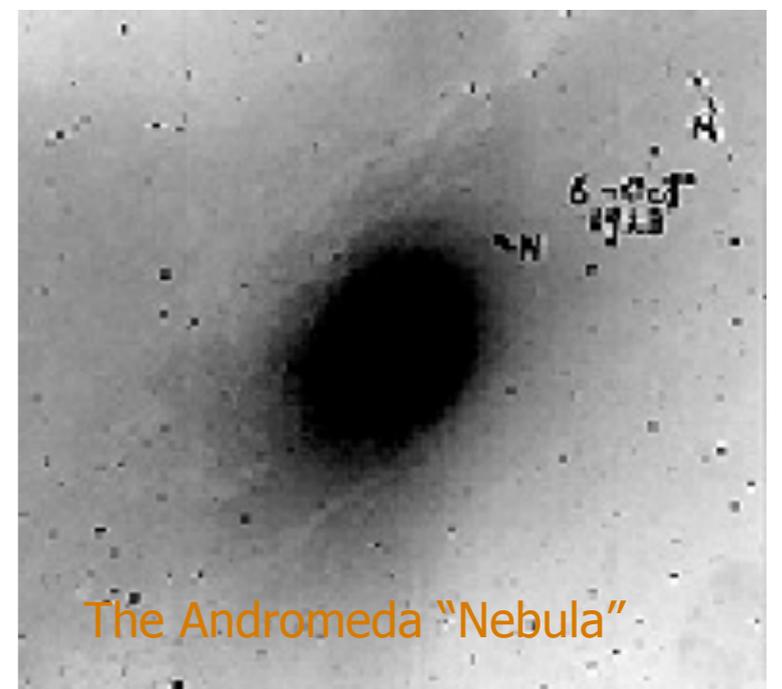
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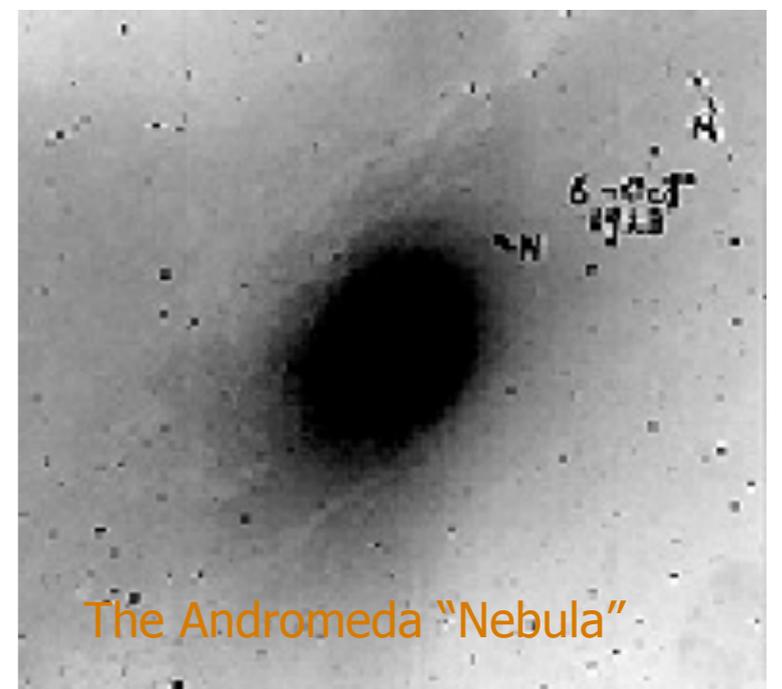
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Andromeda is an “island universe” like our own Galaxy.



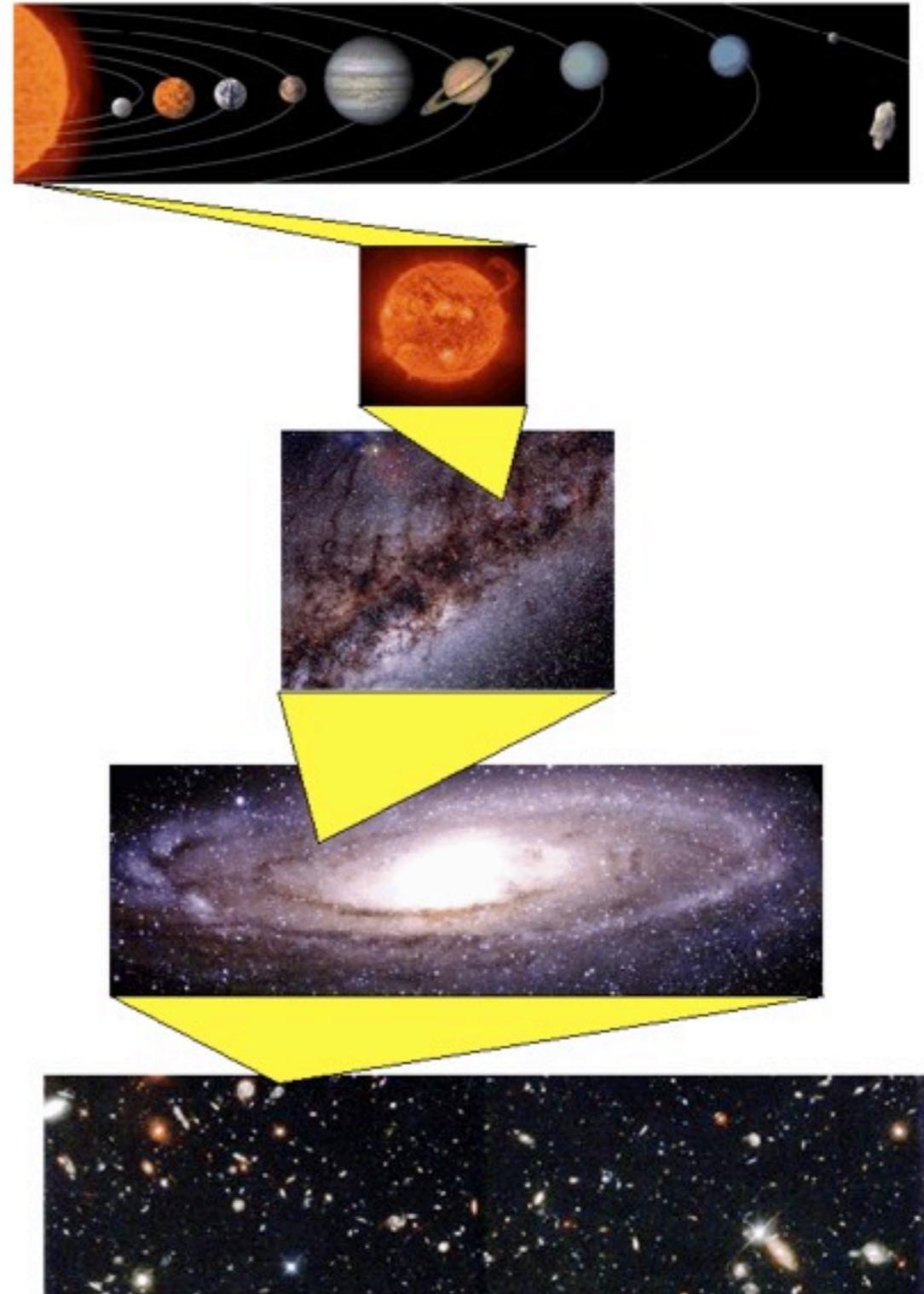
Hubble at Mt. Wilson Observatory



The Andromeda “Nebula”

One of

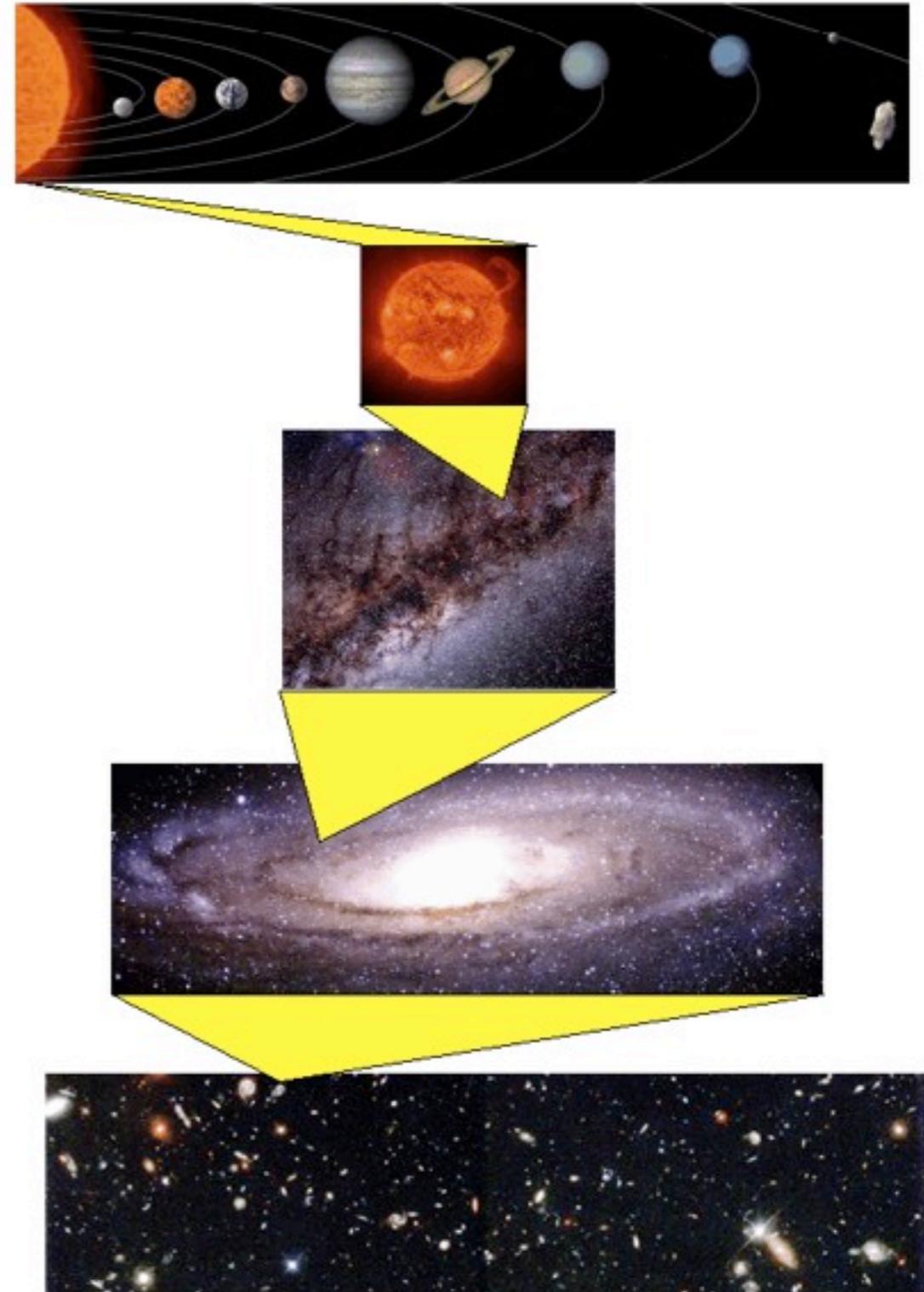
We are:



One of

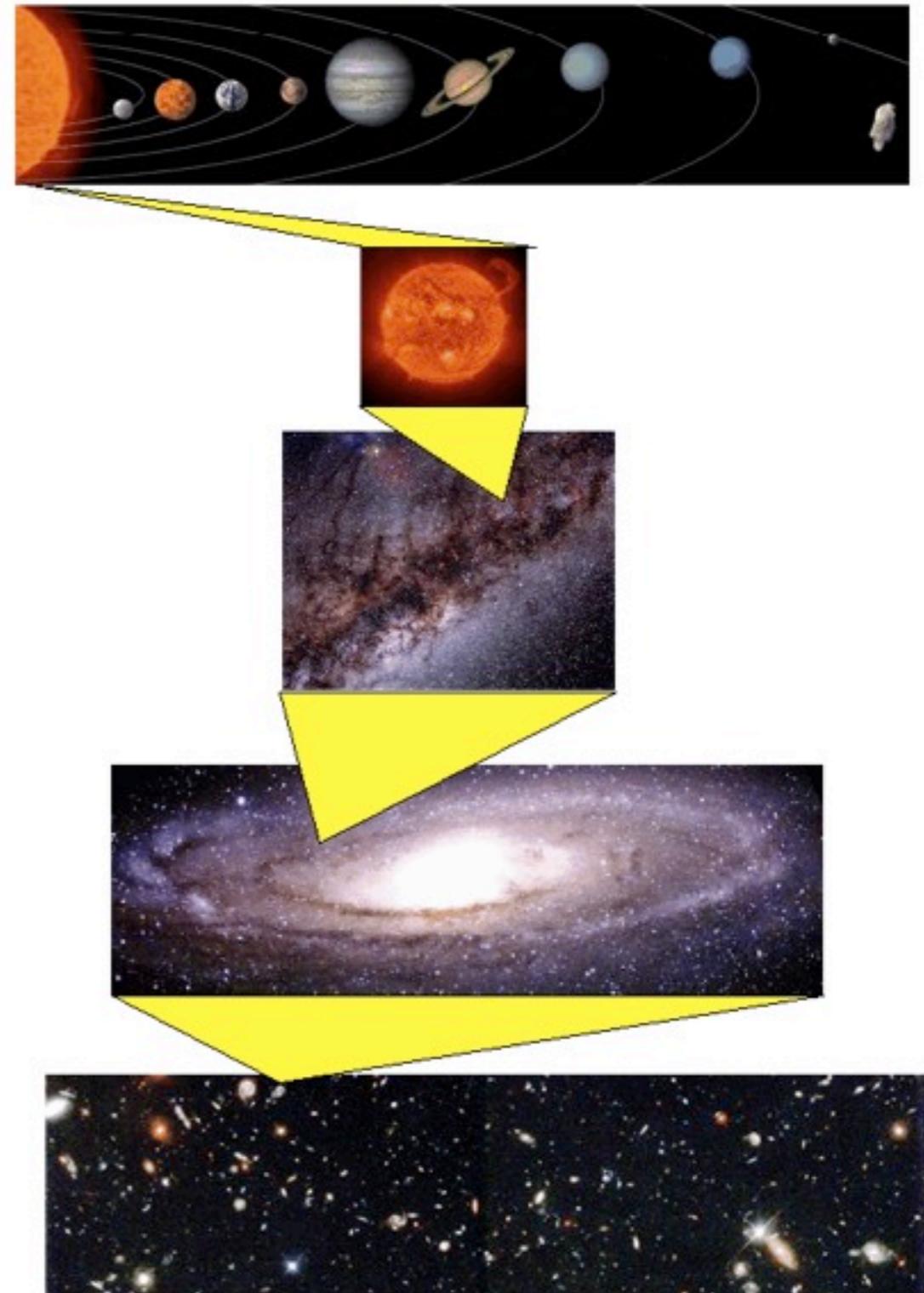
We are:

**1 planet out of 9 in
our solar system.**



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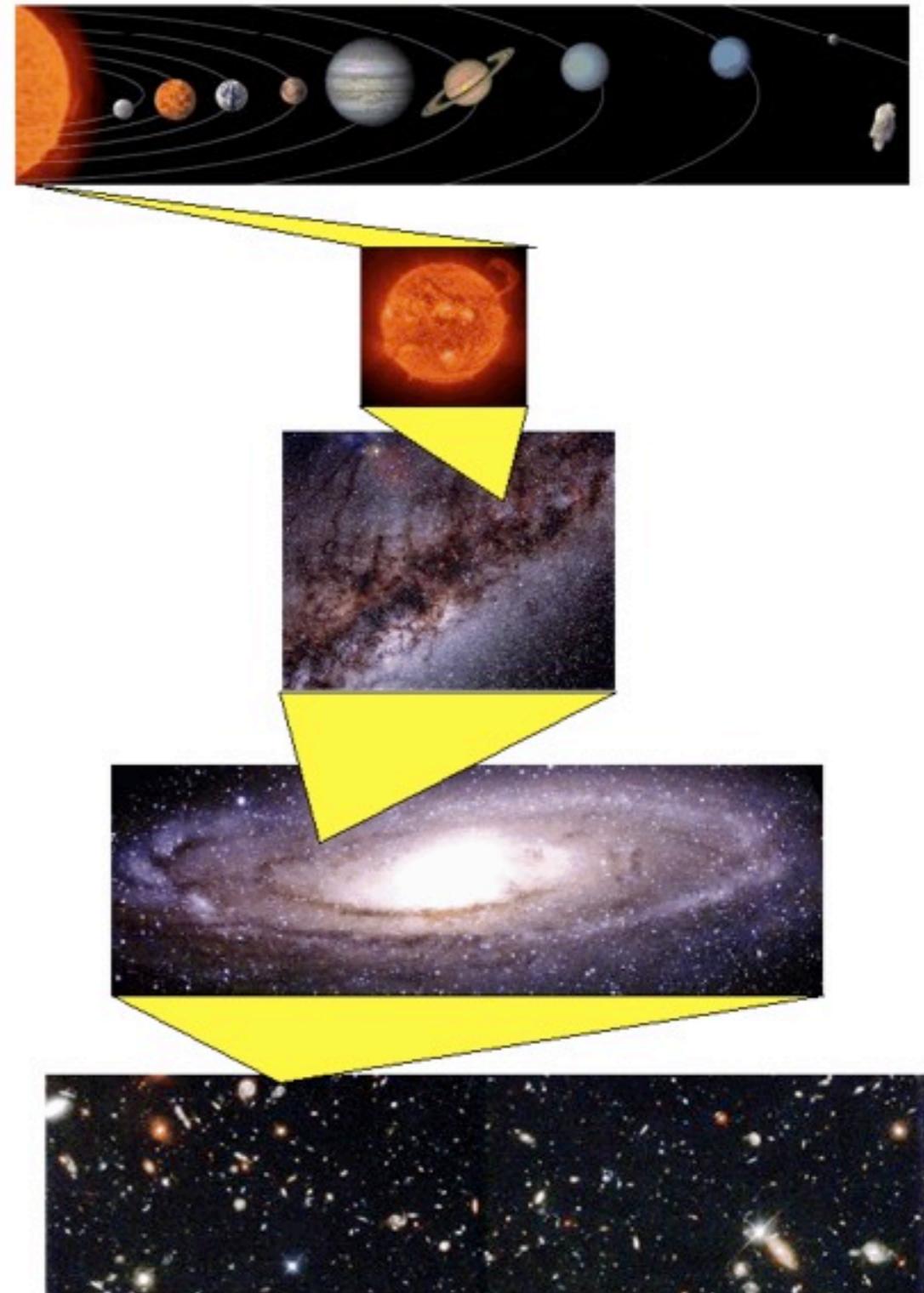


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1 stellar system of
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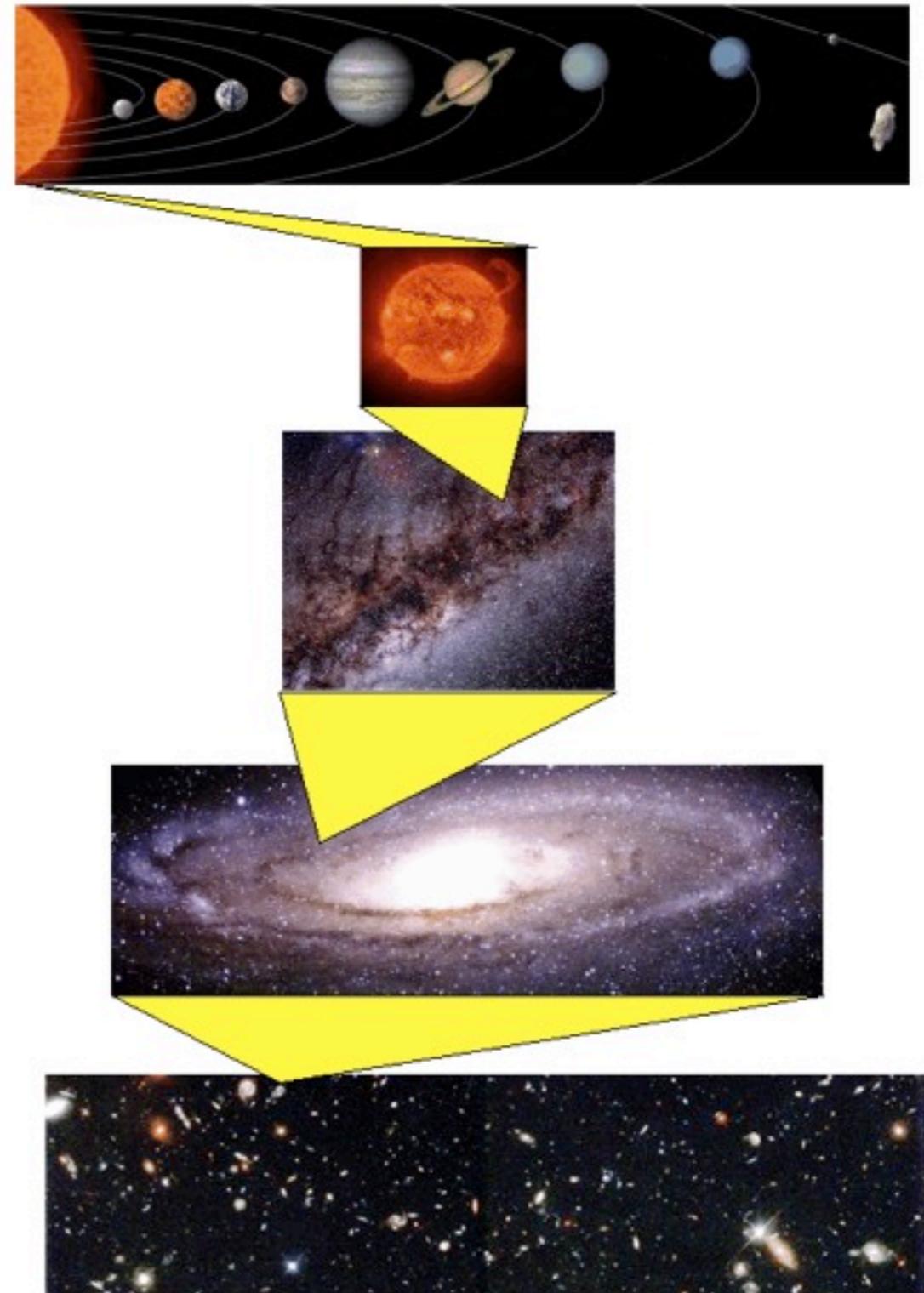
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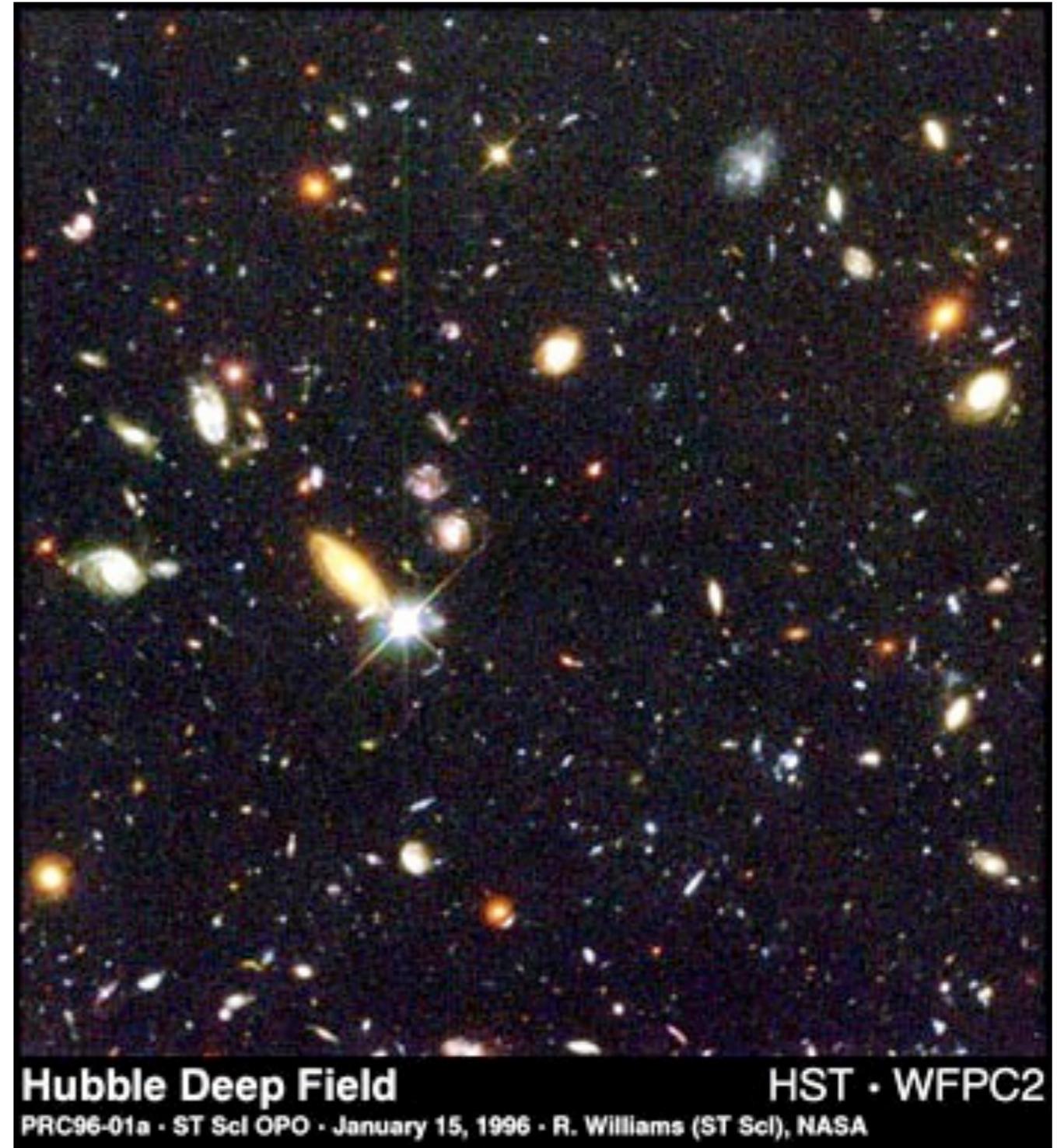
Galaxies – Fundamental “Ecosystems” of the Universe



Hubble Deep Field HST · WFPC2
PRC96-01a · ST ScI OPO · January 15, 1996 · R. Williams (ST ScI), NASA

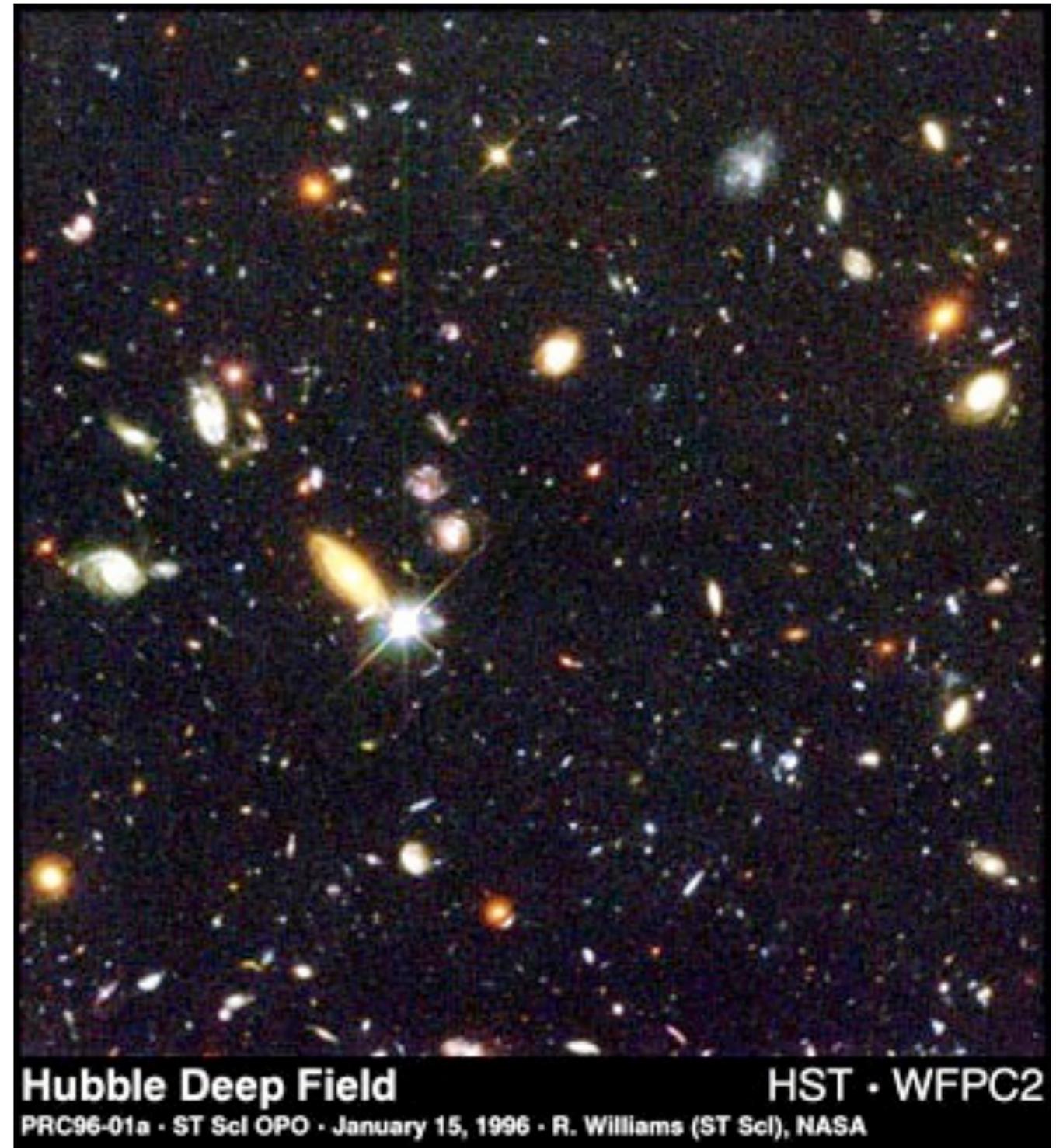
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Galaxies “fill” universe.



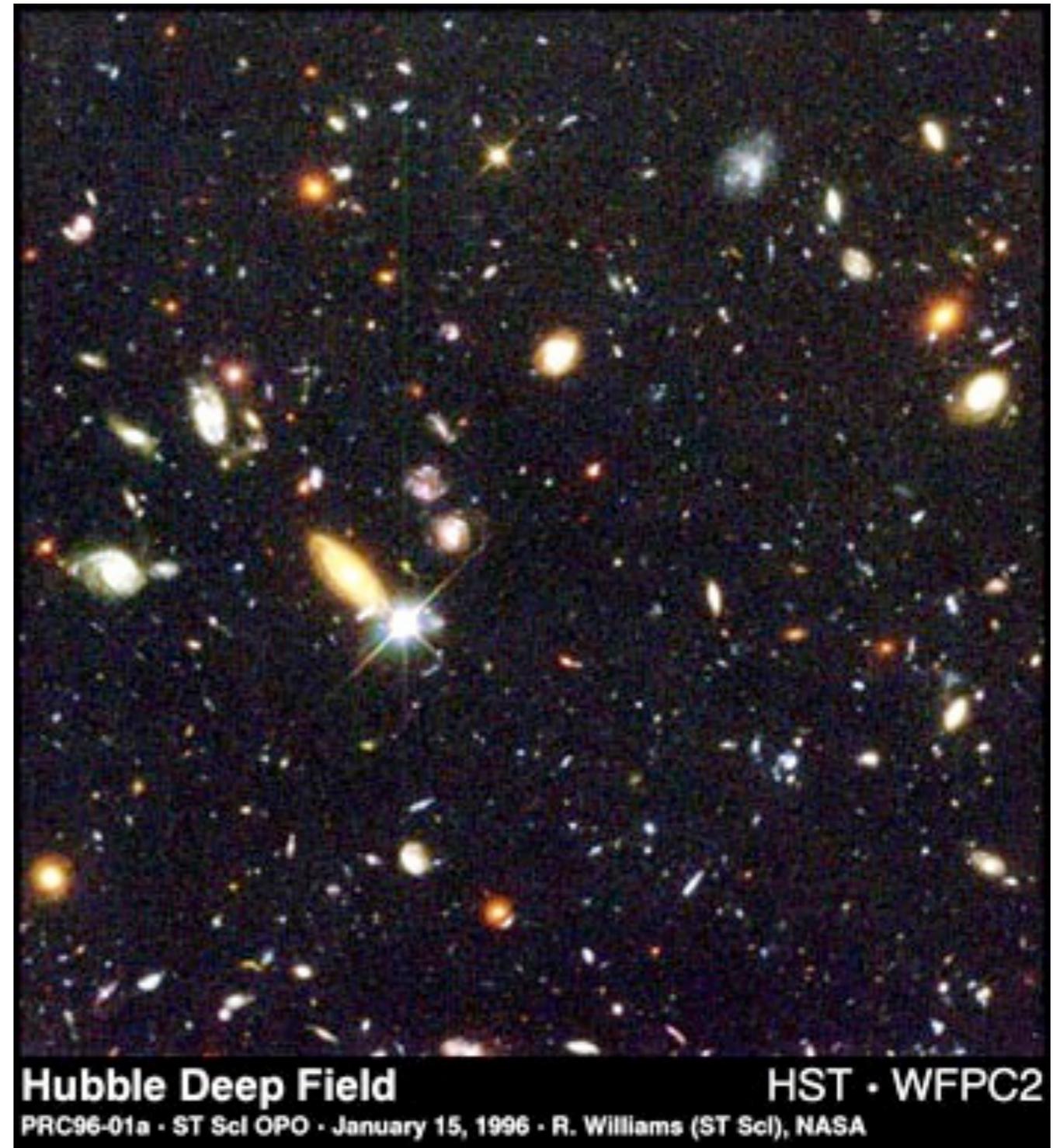
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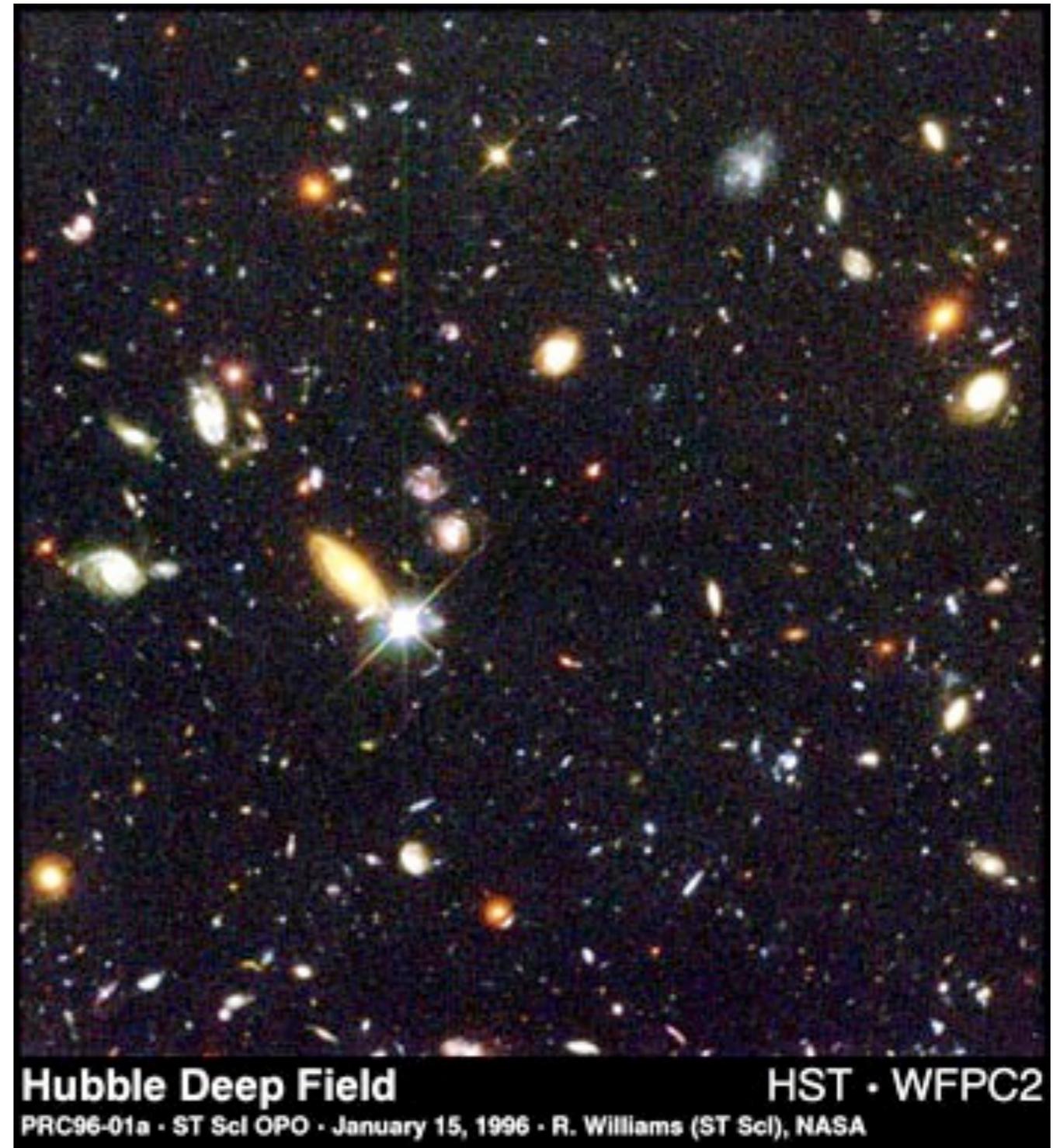
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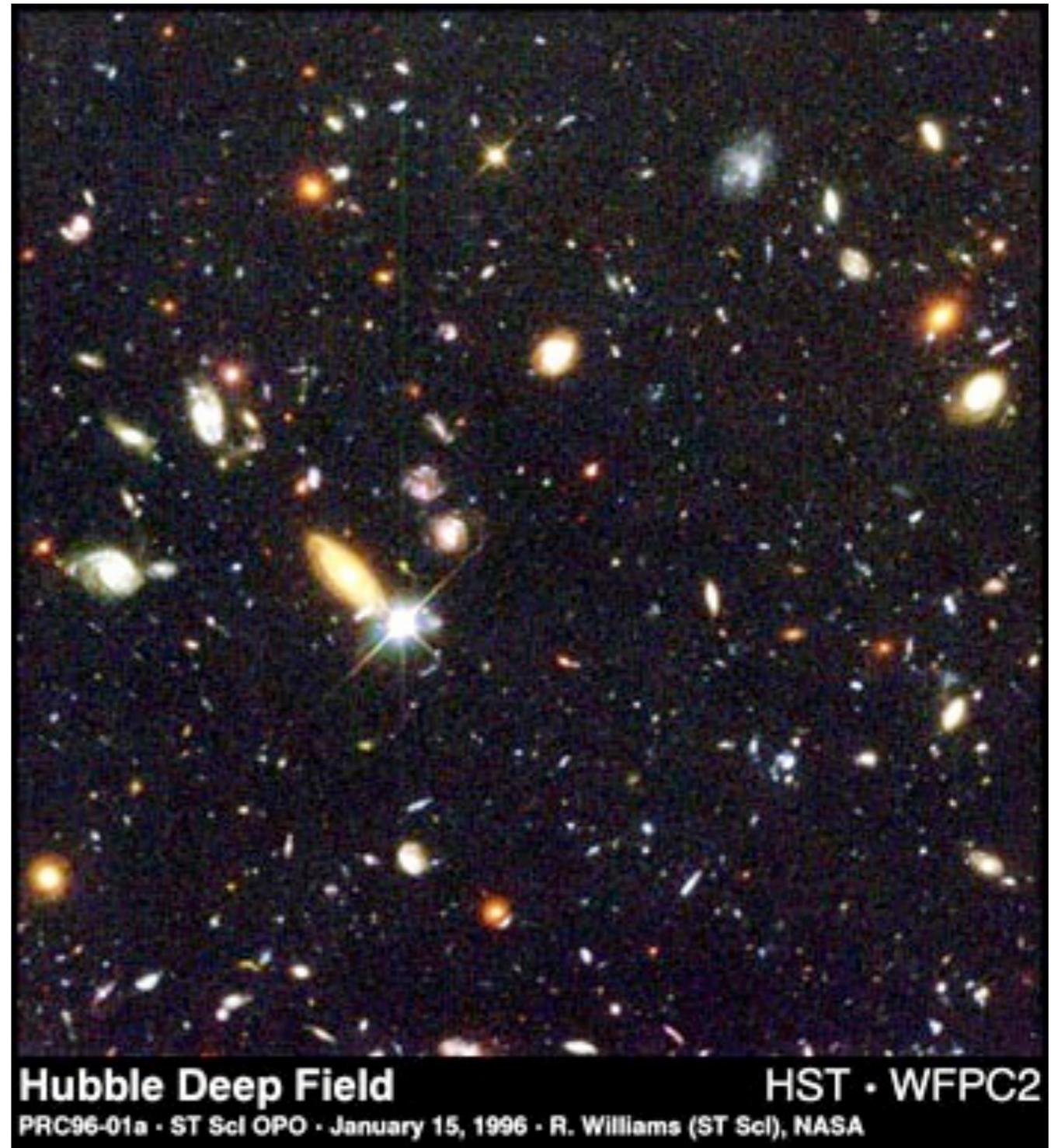
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Range in size from large (MW-like) to small “Dwarf”

- ▶ 1 billion to 500’s billions of stars

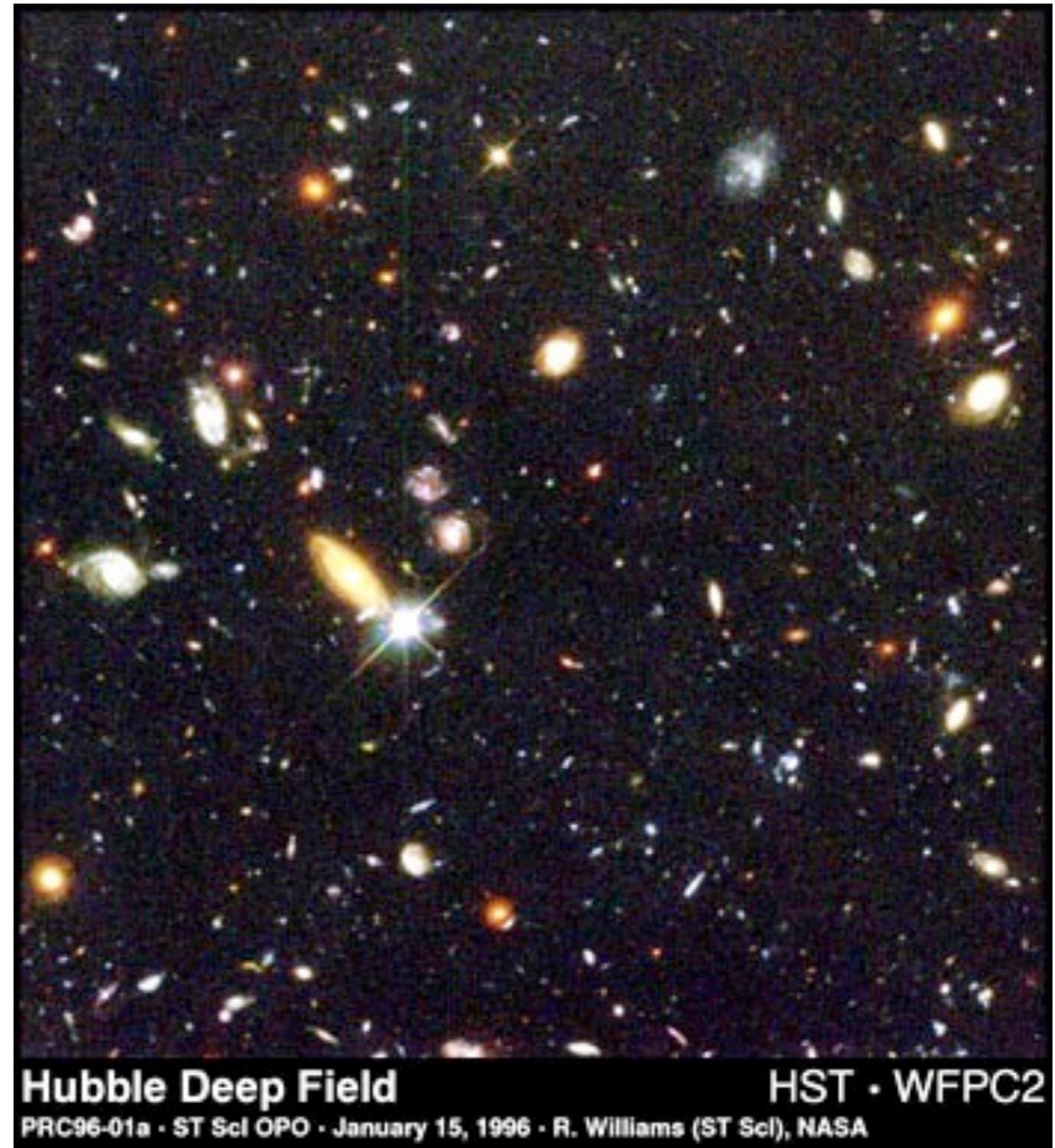


Galaxies – Fundamental “Ecosystems” of the Universe

Galaxies are the cosmic engines that turn gas into stars, then recycles the gas the stars eject back into stars, around and around.

In between galaxies, no star formation occurs – “nothing happens” in intergalactic space.

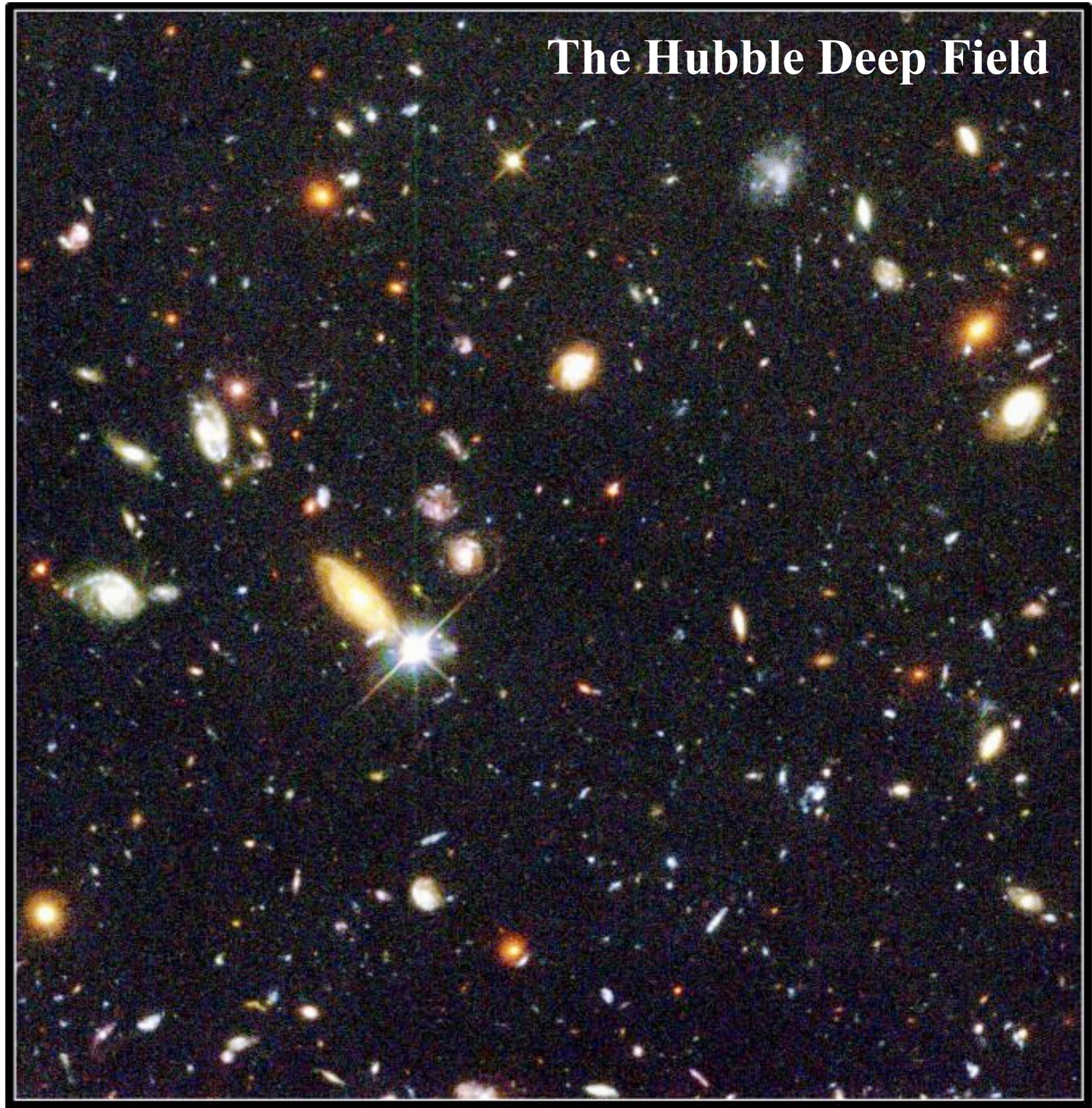
- ▶ Caveat is galaxy clusters

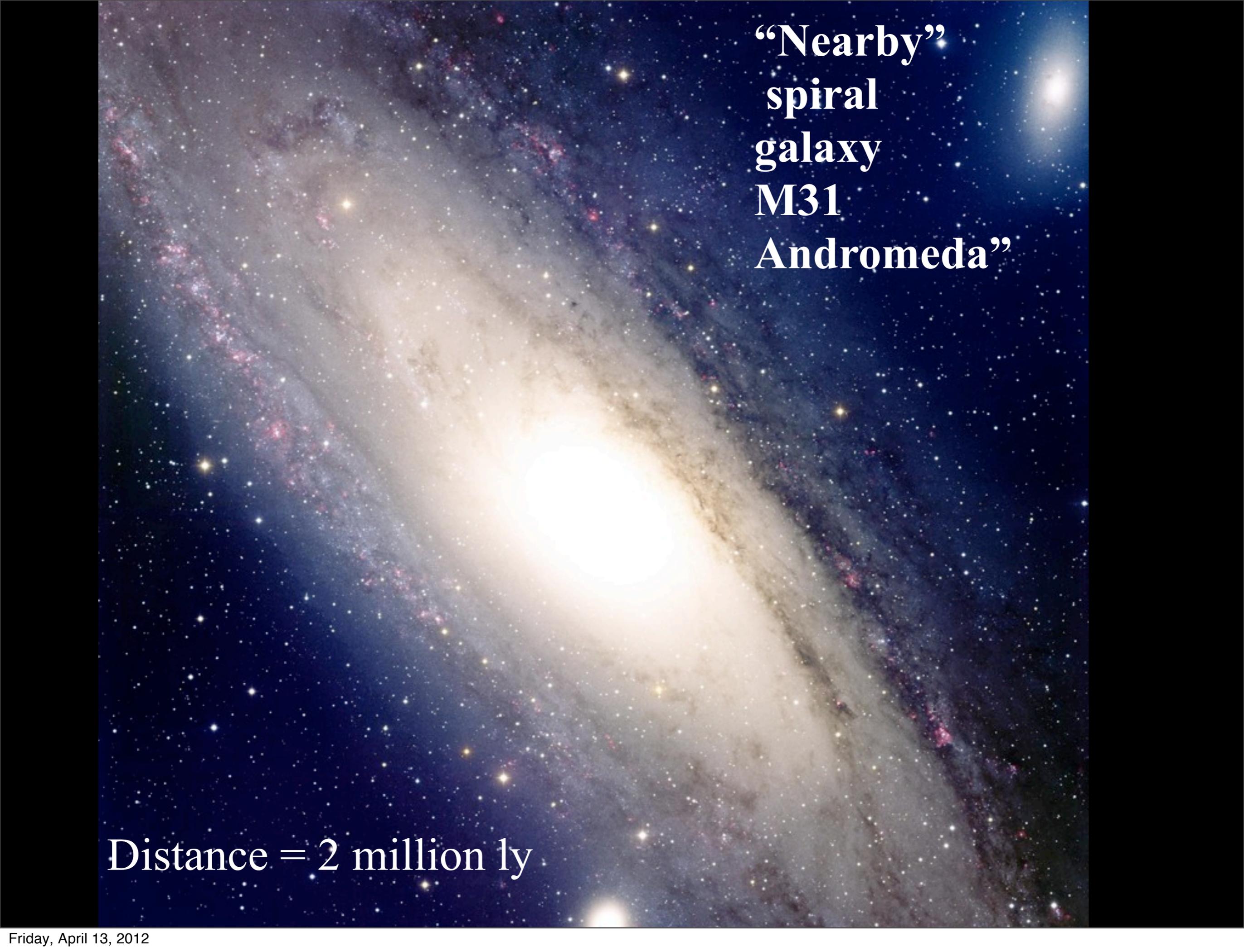


Distant galaxies:

- The deepest optical image of a patch of sky
- Like looking back in time
- ...
- Galaxies as they were, 1 to 10 billion years ago.

The Hubble Deep Field

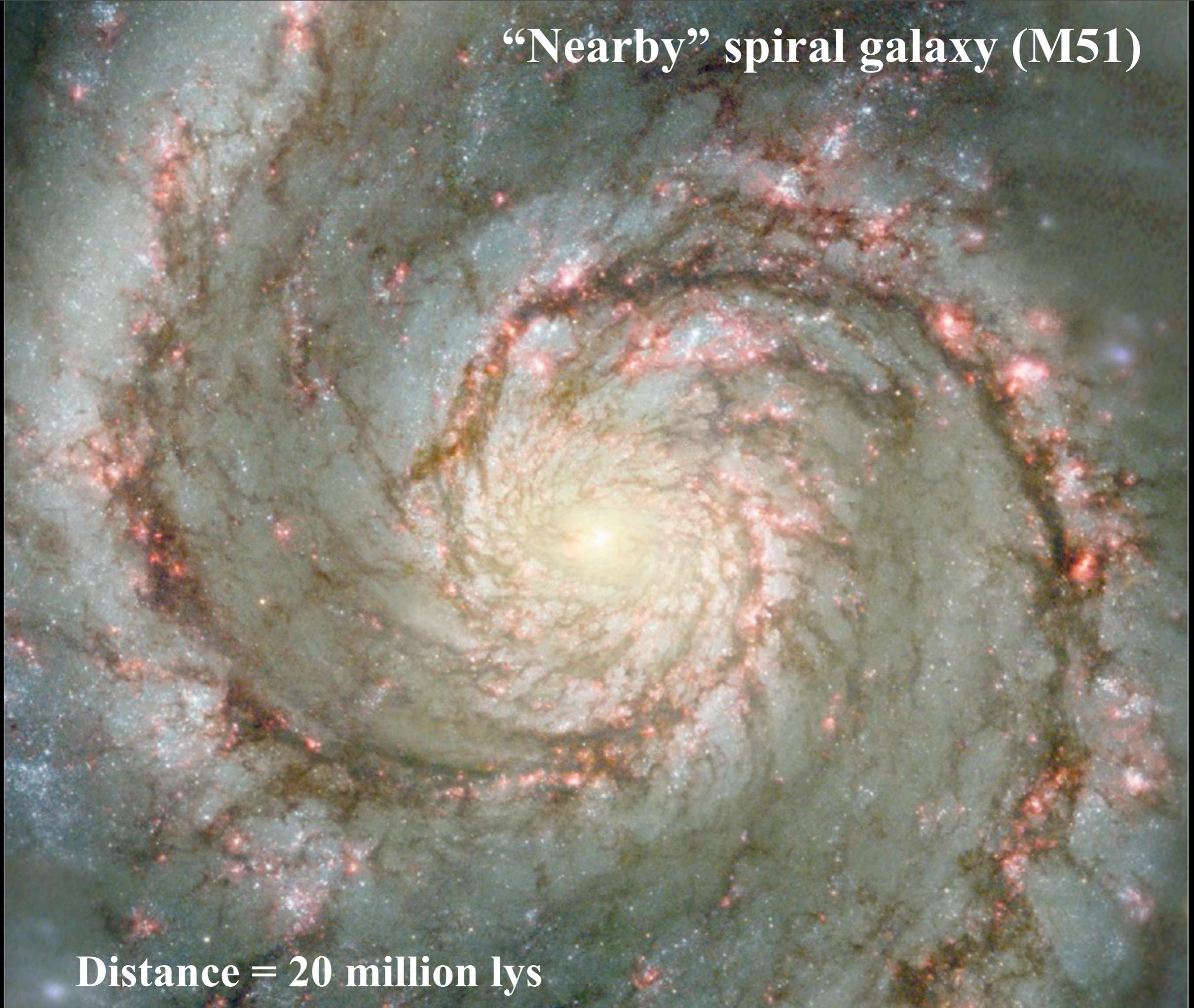




**“Nearby”
spiral
galaxy
M31
Andromeda”**

Distance = 2 million ly

“Nearby” spiral galaxy (M51)



Distance = 20 million lys

Sombrero Galaxy: $\sim 10^{12}$ stars (20 million lyrs)



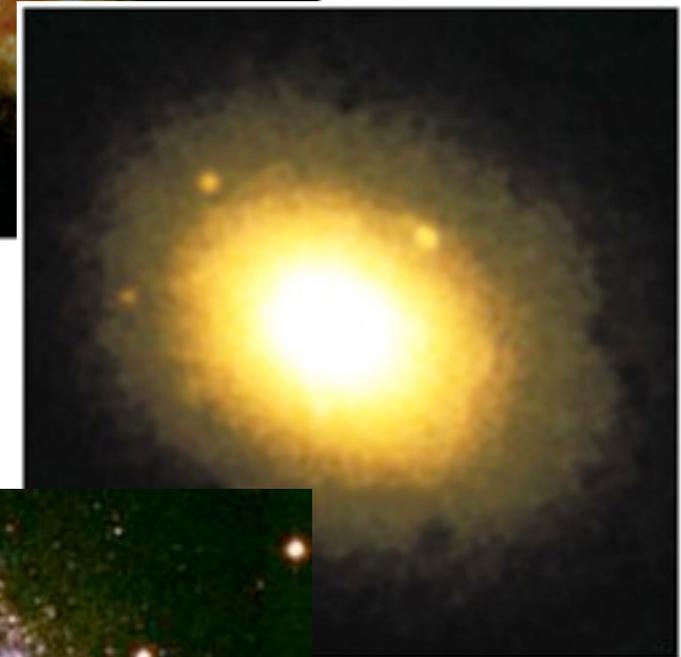
Galaxies are the Fundamental “Ecosystems” of the Universe

Three Main Types of Galaxies:

- Spirals (77%)
- Ellipticals (20%)
- Irregulars (3%)



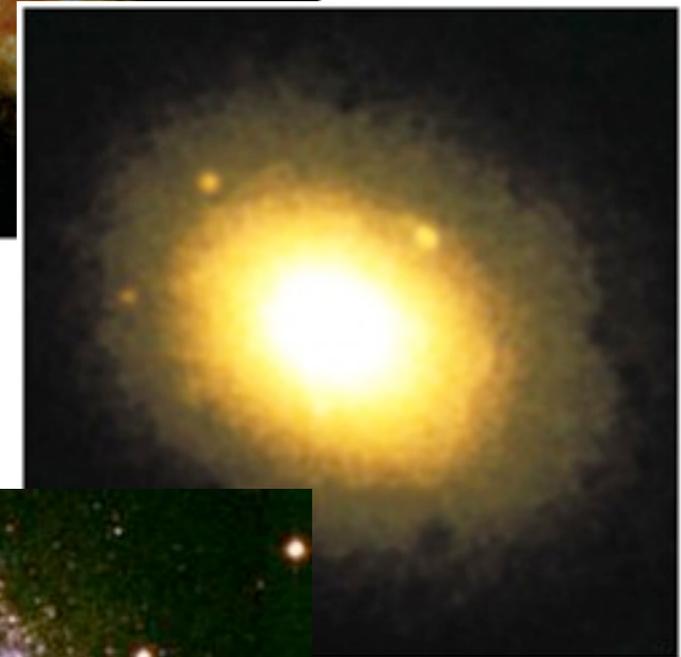
Classes of Galaxies



Classes of Galaxies

Spirals (S)

- ▶ Basic structure: disk and bulge
- ▶ Medium to large galaxies
- ▶ The disk has the young blue stars, while the bulge has older red stars



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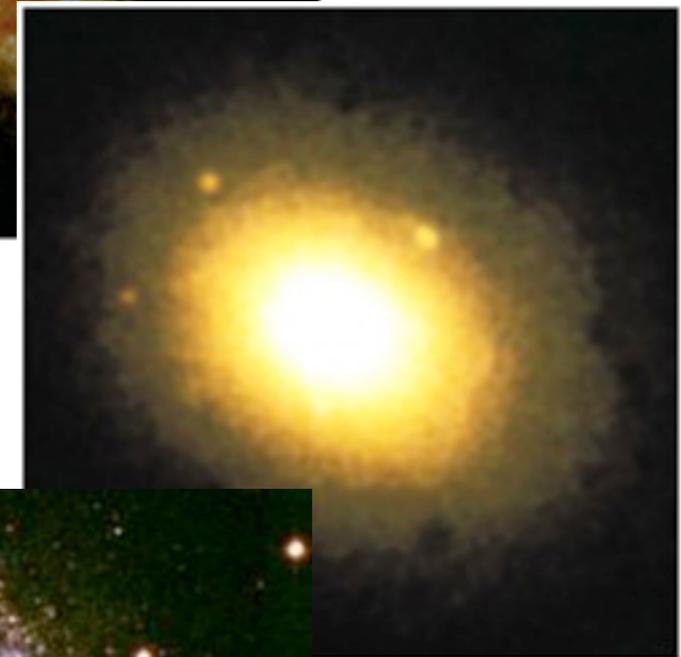
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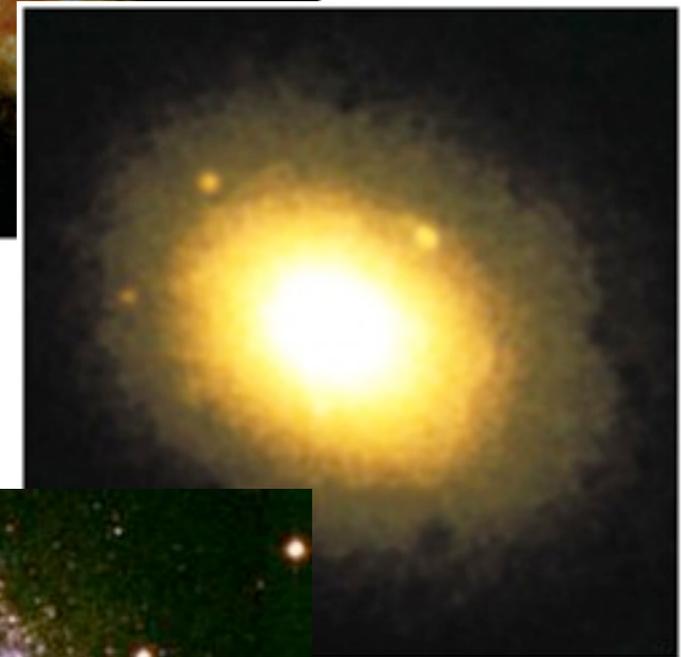
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Irregulars (Ir)

- ▶ Well... odd, irregular structure
- ▶ Smaller galaxies
- ▶ Mostly young blue stars



Spiral Galaxies



Spirals are classified on the amount of bulge component (and how tightly the arms are wound)

These are designated as Sa, Sb, Sc, in order of decreasing bulge

More bulge and tightly wound

More disk and loosely wound

More disk means more ongoing star formation!

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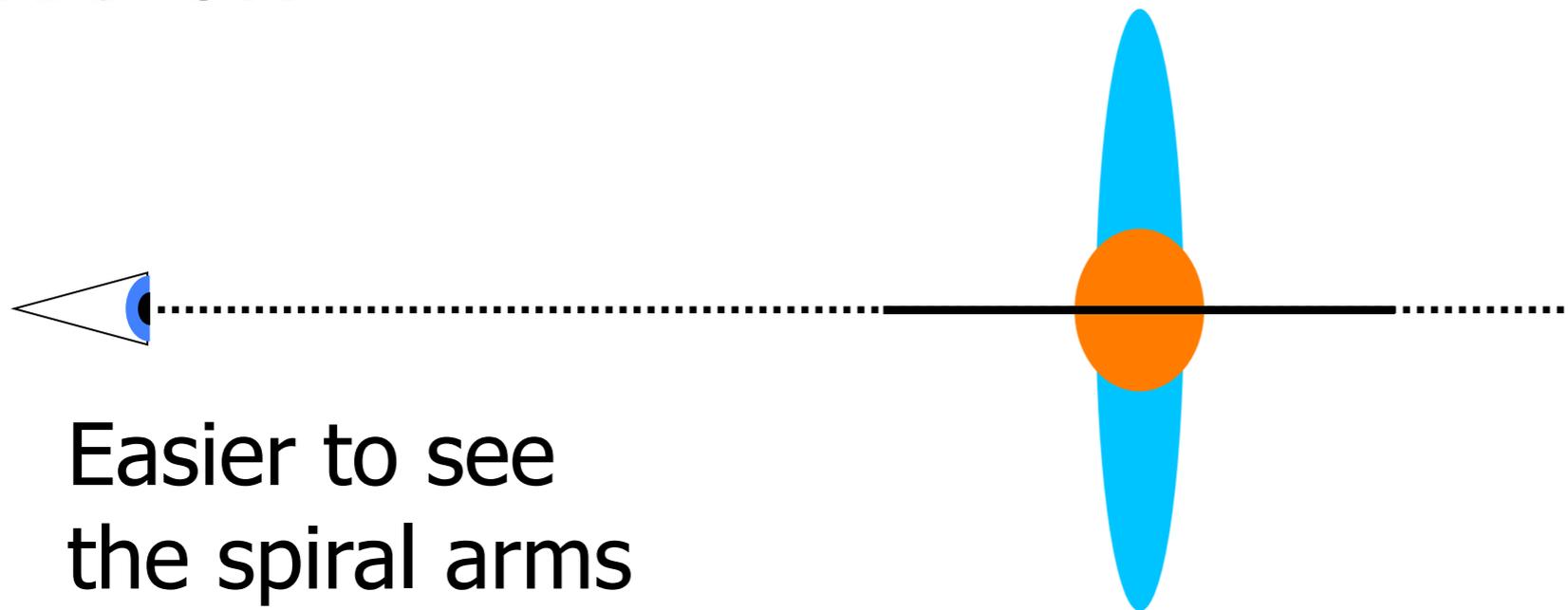
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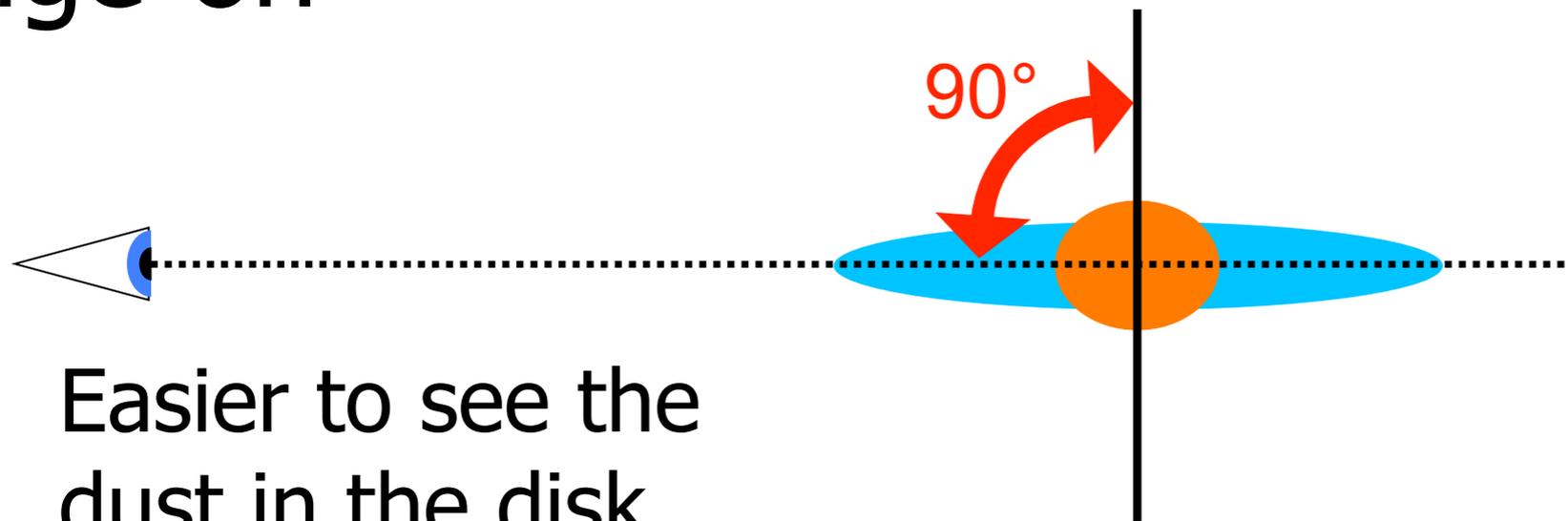
Effect of Viewing Angle

Face-on



Easier to see
the spiral arms

Edge-on



Easier to see the
dust in the disk

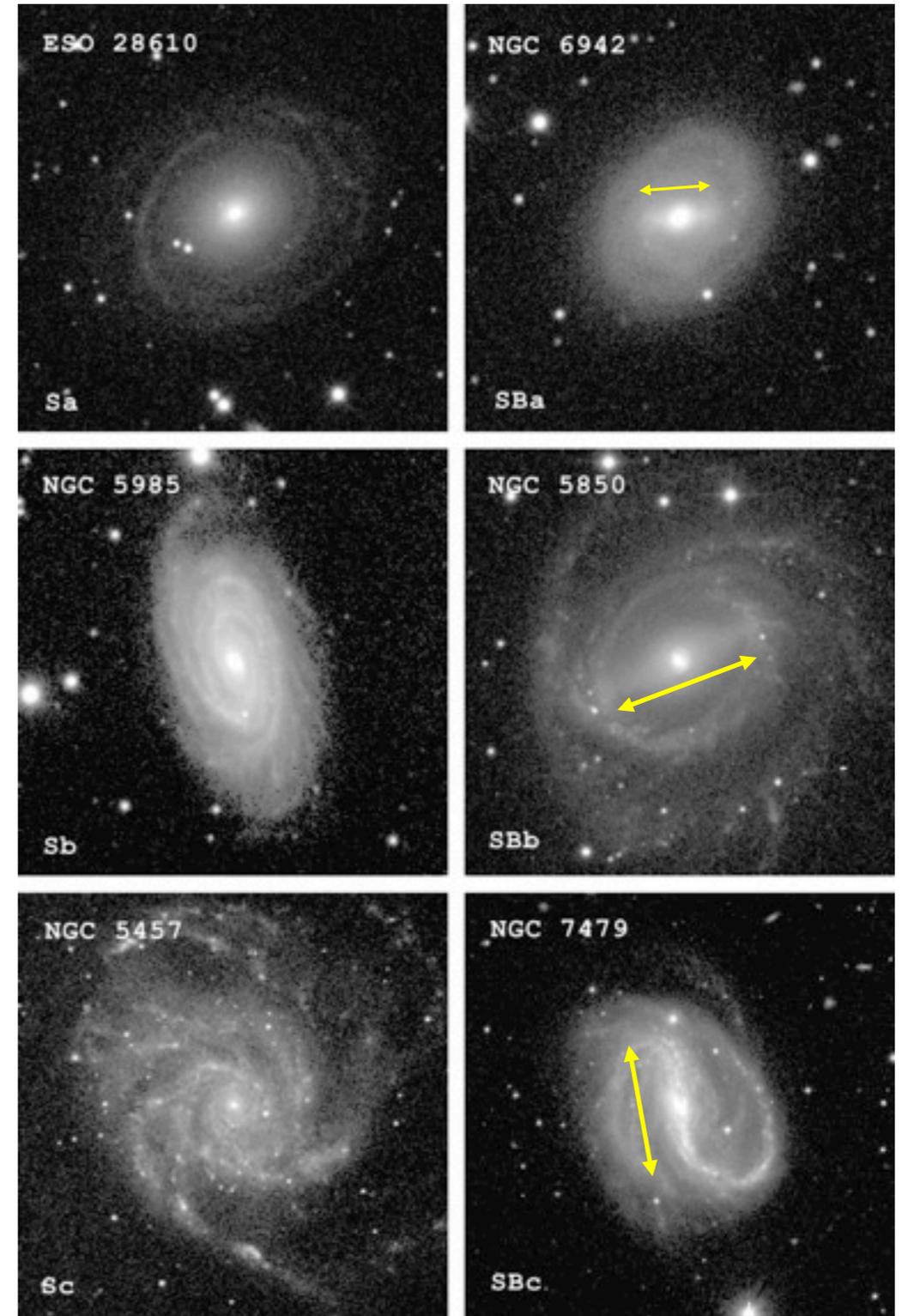
Barred Spirals

About 20% of all spirals are **barred spirals**

The spiral arms branch off from a straight bar of stars that passes through the central bulge

They are designated with an “SB” rather than the usual “S” for spiral galaxies

The classes of barred spirals are SBa, SBb, and SBc



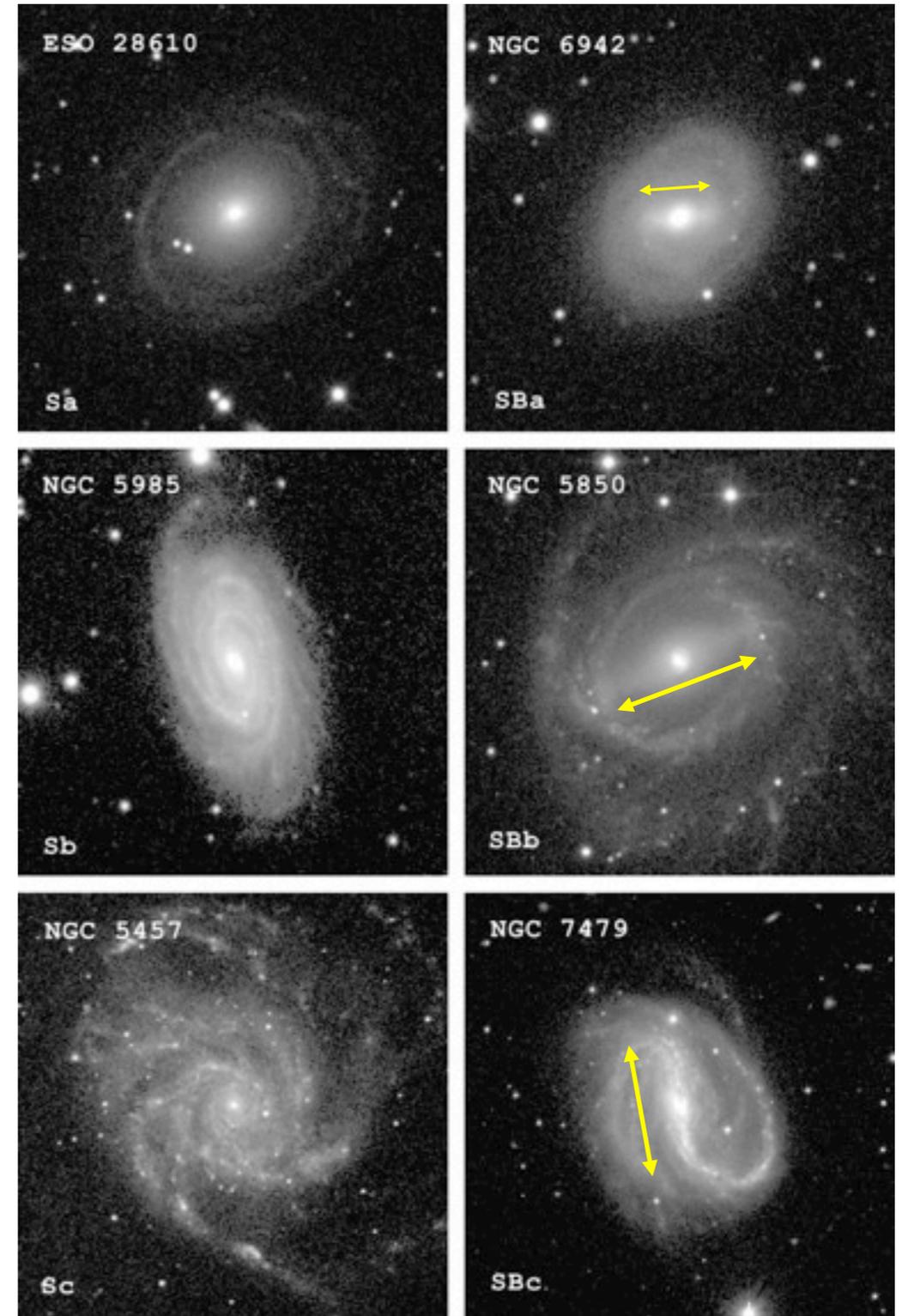
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Elliptical Galaxies

Like a spiral galaxy's bulge

- ▶ Mostly old, redder stars, little gas and dust
- ▶ No disk organization, stars on random orbits

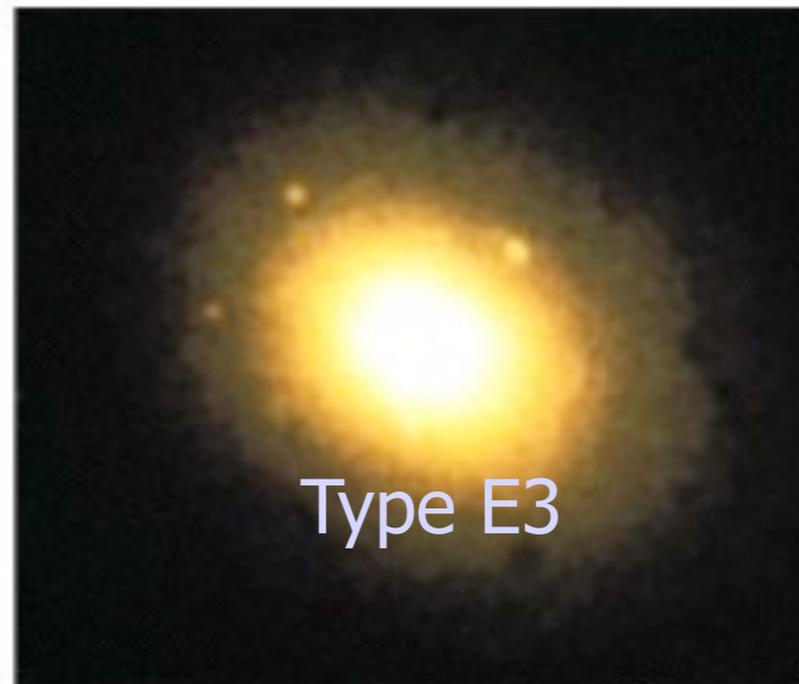
Classified by how elliptical they appear

- ▶ E0 (spherical) to E7 (elongated)



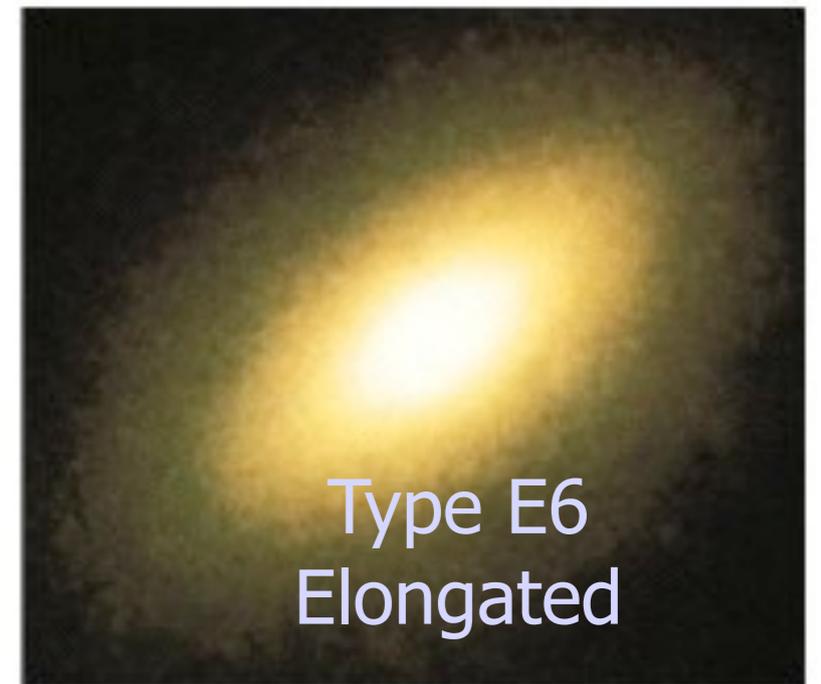
E0

M105



E3

NGC 4365



E6

NGC 3377

Varieties of Elliptical Galaxies



Varieties of Elliptical Galaxies

Ellipticals come in a great range of masses



Varieties of Elliptical Galaxies

Ellipticals come in a great range of masses

The largest are giant ellipticals

- ▶ Up to 100+ times more massive than the Milky Way



Varieties of Elliptical Galaxies

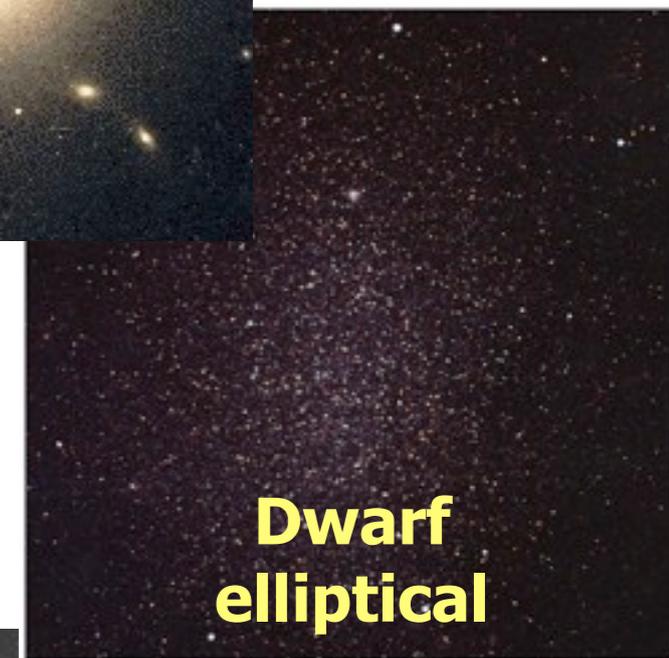
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- ▶ Up to 100+ times more massive than the Milky Way

The smallest are the dwarf ellipticals

- ▶ 10,000 to a million times less massive than the Milky Way
- ▶ Some only a few times larger than a globular cluster!



Varieties of Elliptical Galaxies

Ellipticals come in a great range of masses

The largest are giant ellipticals

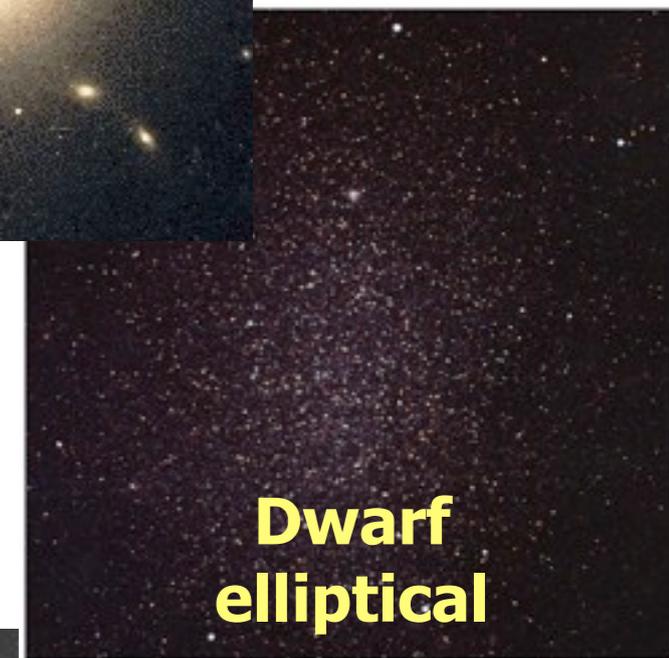
- ▶ Up to 100+ times more massive than the Milky Way

The smallest are the dwarf ellipticals

- ▶ 10,000 to a million times less massive than the Milky Way
- ▶ Some only a few times larger than a globular cluster!

Of course, there are also “garden variety” ellipticals

- ▶ About 100 times smaller than to equal in size to the Milky Way



Irregular Galaxies

Chaotic systems of stars

Prominent examples: The Magellanic Clouds

- ▶ Two of the Milky Way's satellite galaxies

Generally smaller galaxies

- ▶ Thousands to tens of times smaller than the Milky Way

Chaotic systems of stars

- ▶ No disk, no elliptical structure

Dominated by young, blue stars



What Type of Galaxy is the Milky Way?



What Type of Galaxy is the Milky Way?

The Milky Way is a spiral galaxy

▶ Probably type Sb



What Type of Galaxy is the Milky Way?

The Milky Way is a spiral galaxy

▶ Probably type Sb

But is likely a barred spiral!

▶ So, type SBb



Measure other Galaxies Rotation Curves

What do you think we find?

- a) No other galaxies have dark matter**
- b) Some other galaxies have dark matter**
- c) All other galaxies have dark matter**

Masses of Galaxies

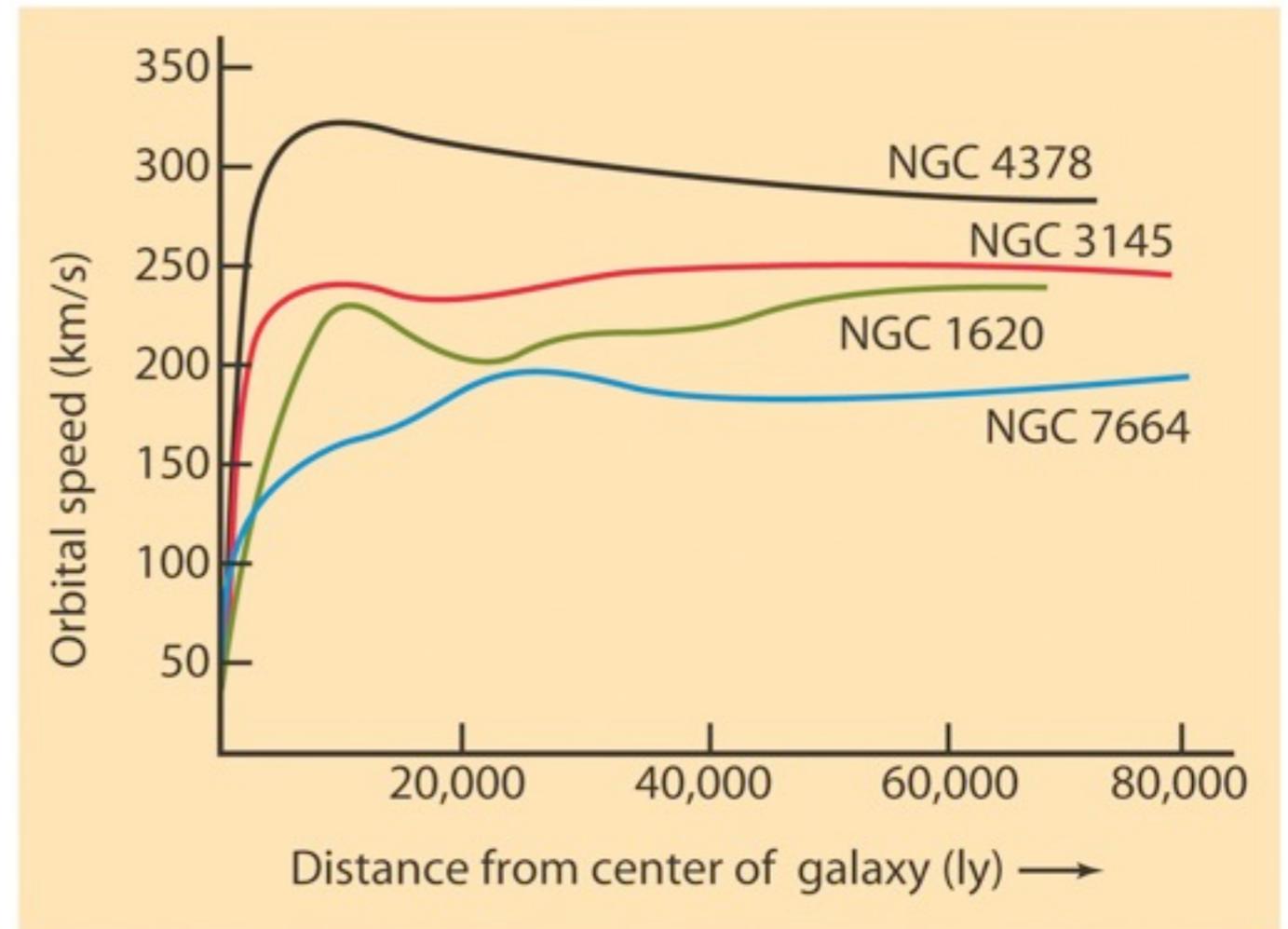
As with the Milky Way, we measure the speed of a galaxy's rotation

Like the Milky Way, other galaxies have a flat rotation curve

Indicates a halo of **dark matter**

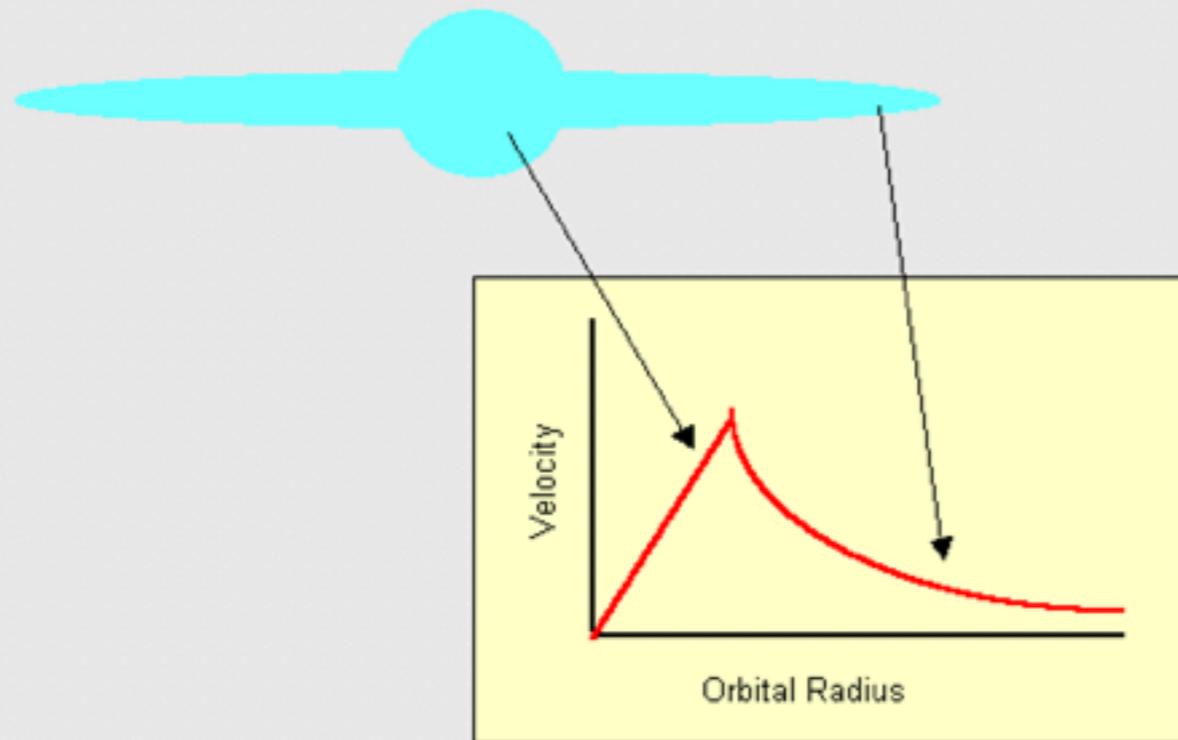
We aren't special that way either.

Dark matter fills the Universe!



Galaxy Rotation Curve

"Expected" Galactic Rotation Speeds



Where's the matter?

