

Astro 501: Radiative Processes
Lecture 22
March 8, 2013

Announcements:

- **Problem Set 6** due today 5pm
- **Problem Set 7** out this weekend, due next Friday
- Midterm Exam: bad grading elves! bad!

Last time: synchrotron

Today: Compton scattering

Erratum: Carbon-14

radioactive isotope ^{14}C , $t_{1/2} = 5730 \text{ yr}$

- used in radiocarbon dating
- has its origin in cosmic rays

how? cosmic-ray ions hit Earth's atmosphere
collide with nuclei in air \rightarrow ^{14}C is part of debris
“rains out,” incorporated into bio matter

in class I claimed: $p + ^{14}\text{N} \rightarrow n + ^{14}\text{C}$

in real time, sharp-eyed students saw this is *wrong!*

Q: *why?*

in fact: “radiocarbon” ^{14}C is produced several ways
but the dominant channel is: $n + ^{14}\text{N} \rightarrow p + ^{14}\text{C}$

~ Q: *what questions does this raise?*

www: neutron monitors

Astrophysical Context: Blazars

we met radio galaxies in the context of synchrotron radiation
but there are many beasts in the active galaxy zoo

Blazars

- seen as luminous nuclear region
at center of giant elliptical galaxies
www: optical blazar images (*R*-band)
- but *do not* show the elongated jets seen in radio galaxies
- flux shows rapid and large-amplitude time variability
- subclasses: BL Lacertae objects—weak radio emission
optically violent variables (OVV)—strong radio emission
- demographics: many fewer blazars than other AGN
e.g., Seyfert galaxies
www: AGN demographics plot (INTEGRAL)
- ω • blazar emission spans radio to TeV gamma rays

Q: what does this suggest about the nature of blazars?

Blazars: Staring Down the Jet

AGN “Unification Model” www: unification cartoon

idea: all active galaxies have similar physical conditions

- a supermassive black hole (SMBH)
possibly actively accreting matter
- a surrounding accretion disk, and dusty torus
- a relativistic jet, if SMBH is actively accreting

in unification picture: *blazar = jet pointing directly at us!*

“looking down the barrel of the gun”

emission from small region of jet “tip” → highly variable

blazar spectra www: example

over full EM range, two large features

- power-law rise from radio, peaks near optical
- falls to X-rays, then peak and power-law fall at gamma-ray

⌞

Q: what could be going on?

Blazar Spectra

Power-law rise from radio to \sim optical

- nonthermal
- similarity with radio galaxies suggests *synchrotron origin* from relativistic electrons in jet

Peak and power-law fall in gamma rays

- in non-flare (“quiescent”) state,
gamma-ray energy content similar to synchrotron
- suggests similar origin
→ perhaps a *reprocessing* of the synchrotron photons
- reprocessed how? *by the relativistic electrons themselves!*

Thus: we want to understand how relativistic electrons interact with photons *Q: the name for which is...?*

o

Note: origin of blazar high-energy emission still debated!

Compton Scattering

Thomson Scattering

We already discussed the scattering of light by electrons in the context of *Thomson scattering*

Thomson highlights:

energies of incident photon ϵ and scattered photon ϵ_1 related by

$$\epsilon = \epsilon_1 \quad (1)$$

differential cross section, with $\hat{k} \cdot \hat{k}_1 = \cos \theta$

$$\frac{d\sigma_T}{d\Omega} = \frac{1}{2} r_0^2 (1 + \cos^2 \theta) \quad (2)$$

total cross section, with $r_0 = e^2/m_e c^2$

$$\sigma_T = \frac{8\pi}{3} r_0^2 \quad (3)$$

∞

Q: what assumptions went into this? When will they fail?

Enter the Quantum: Compton Scattering

Thomson scattering derived for classical EM wave

$\nu = \nu_1$ classically

carrying this to photon picture: $h\nu = h\nu_1$

→ coherent or elastic scattering

but really: photon quanta carry momentum and energy

→ and electron will *recoil* and carry away energy

→ expect scattered photon to have less energy,
and to move in a different direction

Compton: treat light as massless particle

for photon incident on electron *at rest*

conservation of energy and momentum implies

$$\epsilon_1 = \frac{\epsilon}{1 + (\epsilon/m_e c^2)(1 - \cos \theta)} \quad (4)$$

scattered photon energy is lower, and direction different

so the wavelength shifts by

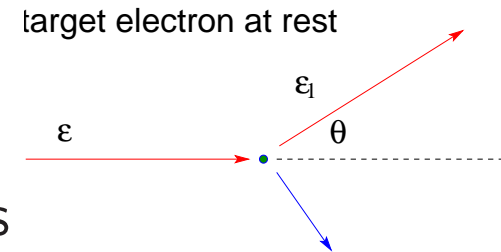
$$\lambda_1 - \lambda = \lambda_c(1 - \cos \theta) \quad (5)$$

where the electron *Compton wavelength* $\lambda_c = h/m_e c = 0.02426 \text{ \AA}$

Q: what energy does a photon with λ_c have?

Q: What region of the spectrum is this?

Q: when is the wavelength shift important? negligible?



Cross Section for Compton Scattering

Compton wavelength shift is $\Delta\lambda \sim \lambda_c$

- *small* if $\lambda \gg \lambda_c$ i.e., $h\nu \ll m_e c^2$
i.e., radio through soft X-rays
- *large* if $h\nu \gg m_e c^2$: hard X-rays, gamma rays

differential cross section: **Klein-Nishina** formula

$$\frac{d\sigma}{d\Omega} = \frac{r_0^2}{2} \frac{\epsilon_1^2}{\epsilon^2} \left(\frac{\epsilon}{\epsilon_1} + \frac{\epsilon_1}{\epsilon} - \sin^2 \theta \right) \quad (6)$$

- classical Thomson expression recovered when $\epsilon \sim \epsilon_1$
- main effect: *smaller* cross section at high energy
- total cross section, with $x = h\nu/m_e c^2$

$$\sigma \rightarrow \begin{cases} \sigma_T (1 - 2x + \dots) & x \ll 1 \\ 3\sigma_T/8 \, x^{-1} (\ln 2x + \dots) & x \gg 1 \end{cases} \quad (7)$$

recall: to understand blazars, we are interested in

- high-energy electrons interacting with ambient photons

Q: why can't we just use the Compton scattering formulae?

Q: how can we use the formulae?

Inverse Compton Scattering

the usual Compton scattering expressions
assume the electron is initially *at rest*
and the *photon loses energy* in scattering
→ “ordinary kinematics”
but this is not the case we are interested in!

in a frame where the electron is relativistic

- then there can be a momentum and energy transfer
and the photon *gains energy*
- “upscattered” to higher frequencies

→ “inverse kinematics” – **inverse Compton scattering**

strategy: use Lorentz transformations *twice*

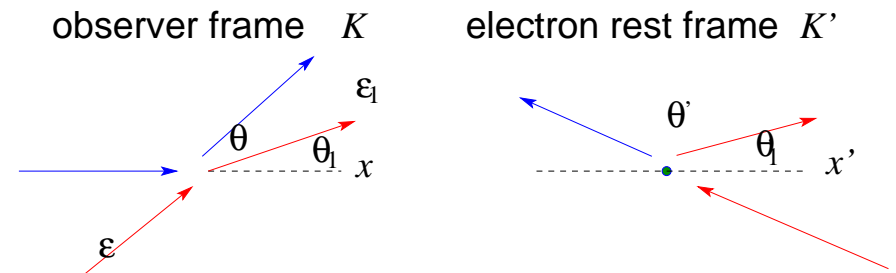
0. start in “lab frame” where e is relativistic
1. **boost** to e rest frame, find scattered photon energy/momentum
2. **boost back** to lab frame to find scattered photon energy

Lab/observer frame K :

- initial electron *relativistic*
- Lorentz factor γ , speed $\beta = v/c$

electron rest frame K' :

- ordinary kinematics



Doppler expression: find photon energies in rest frame

$$\epsilon' = \gamma \epsilon (1 - \beta \cos \theta) \quad (8)$$

$$\epsilon'_1 = \gamma \epsilon_1 (1 - \beta \cos \theta_1) \quad (9)$$

Compton: with $\hat{k}' \cdot \hat{k}'_1 = \cos \Theta$, and assuming $\epsilon' \ll m_e c^2$

$$\epsilon'_1 \approx \epsilon' \left[1 - \frac{\epsilon'}{m_e c^2} (1 - \cos \Theta) \right] \quad (10)$$

if initial lab-frame photon energy is ϵ

Q: initial photon energy in e rest frame, roughly?

Q: scattered photon energy in e rest frame, roughly?

Q: scattered photon energy lab frame, roughly?

Inverse Compton: Order of Magnitude

if initial lab-frame photon energy is ϵ

and lab-frame electron with $\gamma \gg 1$:

in e rest frame: photon energy *boosted*

→ initial energy $\epsilon' \sim \gamma\epsilon$

still in e rest frame:

if $\epsilon' \ll m_e c^2 \rightarrow$ small photon energy change (Thomson)

scattered energy $\epsilon'_1 \sim \epsilon' \sim \gamma\epsilon$

back to lab frame:

scattered photon energy *boosted* to $\epsilon_1 \sim \gamma\epsilon'_1 \sim \gamma^2\epsilon$

15 Bottom line: upscattered photon energy $\epsilon_1 \sim \gamma^2\epsilon$

Q: implications for blazars?

Inverse Compton Power for Single-Electron Scattering

Consider a relativistic electron (γ, β)
incident on an isotropic distribution of ambient photons
→ find power going into inverse Compton

Order of magnitude estimate:

- if typical ambient photon energy is ϵ
then typical *upscattered energy* is $\epsilon_1 \sim \gamma^2 \epsilon$
- if ambient photon number density is n_γ
then *scattering rate per electron* is $\Gamma \sim n_\gamma \sigma_T c$ Q: why?

thus expect power = rate of energy into inverse Compton

$$\frac{dE_1}{dt} \sim \Gamma \epsilon_1 \sim \gamma^2 \epsilon n_\gamma \sigma_T c \sim \gamma^2 \sigma_T c u_\gamma \quad (11)$$

¹⁶ where $u_\gamma = \langle \epsilon \rangle n_\gamma$ is the ambient photon energy density
Q: looks familiar?