Astro 210 Lecture 18 February 28, 2018

Announcements

- HW5 due online in PDF, Friday 5:00 pm
- Office hours: instructor 2:00-3:00 pm today
- TA: 3:30-4:30 pm tomorrow (?)
- Night Observing this week weather permitting Campus Observatory. Wed, Thur 7–9pm bring report form available on Moodle take and submit selfie while there

Craters

Craters caused by meteor/comet impact

- \rightarrow explosion results
- \rightarrow large energy release

Resulting features:

- circular "bowl" cleared out
- in larger craters, central peak ("rebound" of underlying rock)

www: the Moon

 $_{N}$ Q: Why Moon's surface heavily cratered but Earth's not?

Why Moon's surface heavily cratered but Earth's not?

- small meteors burn in E's atmosphere
- ⊳ erosion
- ▷ oceans hide some
- tectonic activity
- volcanos hide some

Some large objects do survive fall

impact on surface

but erosion, geological activity quickly erases evidence

- www: Manicouagan, Canada crater
- www: Clearwater lakes, also Canada
- www: Tunguska, Siberia 1929; exploded in air 1908
- $_{\omega}$ www: Meteor Crater, AZ

Cosmic Calamity!

Mass extinction: dinosaurs died abruptly 65 Myr ago ended Cretaceous era, began Tertiary that continues today

What killed the dinosaurs? Longstanding question

Alvarez+(1980): strong evidence for meteor/comet impact global ~ 1 cm layer at Cretaceous/Tertiary boundary www: K/T boundary found to be (relatively) highly enriched in iridium on Earth: most Ir is in core, binds to Fe in space rocks: not differentiated, higher Ir

Q: what would be the definitive evidence?

T Rex and the Crater of Doom

large impact implies large crater

1990: buried crater found via gravity acceleration anomalies located off Yucatan peninsula, Mexico www: topographical map of Yucatan--note bull's eye

• diameter \sim 180 km – huge!

СЛ

• age (from radioactive 40 K dating): 65 Myrs: when dinos died!

Q: how could a large impactor ruin your day?

K/T Mass Extinction

Lethal Effects

- explosion of impactor and surface material huge numbers of small particles launched into orbit rained down globally, delivering K/T layer but also heat
- fire for hours, heat flux $F_{\rm sky} \approx 10 F_{\odot}$ like pizza oven! global fires. species died if couldn't hide
- **brimstone** raining material sulfur rich \rightarrow acid rain and raised albedo A: skies darkened
- floods tusnamis launched

The Moon

Global Properties

$$\begin{split} M &= 7.3 \times 10^{22} \text{ kg} \\ R &= 1738 \text{ km} \sim 1/4 \text{ } R_{\text{earth}} \\ d_{\text{EM}} &= 3.8 \times 10^5 \text{ km} \sim 60 R_E \end{split}$$

 $ho_{\rm avg} \sim 3000 \ {\rm kg} \ {\rm m}^{-3}$ ightarrow not big metallic core, no magnetism

surface gravity acceleration

$$g_{\text{moon}} = \frac{GM_{\text{Moon}}}{R_{\text{Moon}}^2} = 1.6 \text{ m/s}^2 \simeq \frac{1}{6} g_{\text{earth}}$$

Giant steps are what you take

Walking on the Moon

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– astrophysicists The Police (1979) Reggatta de Blanc

Tides

www: high/low comparison image

www: online data -- pick a beach to visit!

Q: what is tide period: high to high/low to low?

grav. force changes with distance \rightarrow tidal forces compare forces on mass m at different distances

A B

$$F_A = GMm/r^2$$
 $F_B = GMm/(r+d)^2$
 $F_A > F_B$ force tries to pull A and B apart
 \rightarrow tidal force

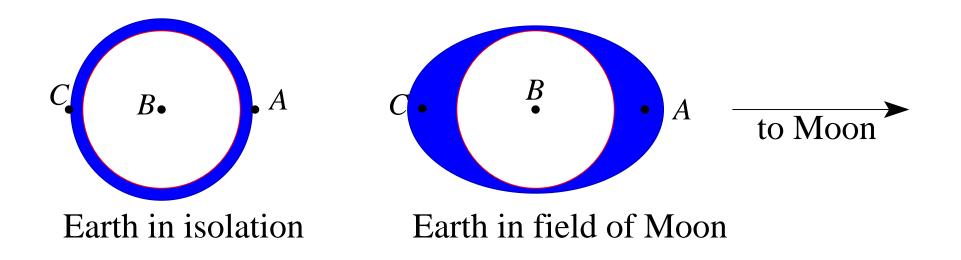
$$F_{\text{tide}} = F_A - F_B \tag{1}$$

$$= GMm\left(\frac{1}{r^{2}} - \frac{1}{(r+d)^{2}}\right)$$
(2)

$$= GMm \frac{(r+d)^2 - r^2}{r^2(r+d)^2}$$
(3)

$$= GMm \frac{d(2r+d)}{r^2(r+d)^2} = GMm \frac{2dr(1+d/2r)}{r^4(1+d/r)^2}$$
(4)

$$\text{if } d \ll r \Rightarrow \frac{F_{\text{tide}} = 2GMm\frac{d}{r^3}}{F_{\text{tide}}}$$



A feels strongest attractionB feels average attractionC feels weakest attraction

so: gravity acclerations $g_C < g_B < g_A$ relative to average $\Delta g = g - g_B$:

$$\Delta g_C < 0 < \Delta g_A$$

The Moon: Orbit

www: lunation animation: always same face! www: far side

Always same side faces us!

demo: lunar globe

iClicker Poll: The Moon & Spin

The Moon always keeps the same face to us What is the Moon's spin period?

- A zero! no spin!
- B nonzero! spin period < orbit period</p>
- C nonzero! spin period = orbit period



nonzero! spin period > orbit period

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Moon has \omega_{orb} = \omega_{spin} exactly!
"co-rotation"
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Why? Tidal interaction and friction

ex: ball rolling in bowl $F_f \neq 0$ after time: stopped $F_f = 0$ \Rightarrow friction drives a system to a state in which frictional forces are no longer active

Earth & Moon deformed by tidal forces sketch imagine $\omega_{spin} > \omega_{orb}$ Q: What is effect on Moon's surface?

 $\tilde{\omega}$ Q: How will this change the spin & orbit over time?

Tidal stresses on Moon \rightarrow Moon surface constantly deformed Deformed Moon non-spherical: tidal bulges Earth gravity on bulges \rightarrow torque increases Moon orbital angular momentum

repeated stretching/compression \rightarrow friction, heating dissipation \rightarrow evolve to frictionless state: reduces Moon spin angular momentum until $\omega_{spin} = \omega_{orb}$

Note: may take long time! complete for Moon, not for earth!

Earth $\omega_{\text{spin}}^E > \omega_{\text{orb}}$ sketch Earth drags along tidal bulges $F_N > F_F$

Two effects

- 1. slows earth spin (reduces ang. mom.) $dP_{\rm spin}/dt \sim 1.6 \times 10^{-5} {\rm \ s/yr} = 16 {\rm \ s/Myr}$
- 2. adds orbital ang. momentum to moon, (still circular) $(v_c = \sqrt{GM/R} \text{ or } \omega_{\text{orb}} = v_c/R = \sqrt{GM/R^3})$ net effect: earth-moon distance *increases*! $dR/dt \sim 2.3 \text{ cm/yr}$

confirmed by laser ranging measurements! www: laser to Moon

Thus:

- moon recedes!
- Moon closer in past!

The Moon: Surface Features

highlands: lighter in color, heavily cratered
 www: Apollo 17 in highlands (mountains made by impacts)

* maria - "seas" (singular: mare): dark plains
www: Mare Imbrium large scale
www: maria/highlands comparison
smooth: fewer craters, made of volcanic rock Q: how do we
know?
formed by lave flows

formed by lava flows

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* craters
cover surface
occur in all sizes, > 20km to microscopic
www: Mare Oriental
www: maria--overlapping craters
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Apollo: 50th Anniversary

1950s-60s: Cold War "space race" to Moon – US vs USSR

- 1957: Sputnik 1 first artificial Earth satellite
- 1959: Luna 2 probe impacts Moon; Luna 3 orbits, sends images
- 1961: Yuri Gregarin first human in space

NASA Apollo

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goal: land and return astronauts

approach: two spacecraft that dock and undock:

- Command & Service Module (CSM) orbits Moon, returns to Earth
- Lunar Module (LM="lem") lands on Moon, left in orbit

Apollo: 50th Anniversaries

Apollo 8 December 1968

- first humans out of low Earth orbit
- first humans to see Earth globally
- first humans out of Earth's gravity well
- first humans in another gravity well
- first humans to orbit Moon

Apollo 11 July 1969

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humans walk on another world

Mare Tranqillitatis: Sea of Tranquility Neil Armstrong (BS Aero Eng!), Buzz Aldrin, Michael Collins www: Apollo images and video