

Astro 210
Lecture 21
March 7, 2018

Announcements

- **HW6 due online in PDF, Friday 5:00 pm**
- Office hours: instructor 2:00-3:00 pm today
- **Night Observing this week** – weather permitting
Campus Observatory. Thursday 7–9pm
bring **report form** available on Moodle
take and submit **selfie** while there
- **Distinguished Lecture & Extra Credit Opportunity**
Prof. Sara Seager, MIT,
“Exoplanets and the Search for Habitable Worlds”
└ **7-8 pm, Tonight, Lincoln Hall Theater**
Selfie+online report = **bonus points** (“extra credit”)

Solar System: Themes

clear patterns

- similarities among inner/terrestrial planets
- similarities among out/Jovian planets

but also diversity emerges as we learn more

each planet is unique in important ways

dot on sky → new worlds to explore

...and the major planets are not the whole story!

Solar System Debris

in addition to planets

Solar system contains large amount of smaller junk

- **rocky debris: asteroids**

concentrated between Mars and Jupiter

i.e., separates terrestrial & Jovian planets

- **icy debris: comets**

concentrated outside of Neptune's orbit

i.e., beyond Jovian planets

Debris I: Asteroids

Properties

“minor planets” number $\sim 10^5 - 10^6$

masses: total $\sim 10^{-5} M_{\text{Earth}}$

sizes: poorly known, but go up to ~ 300 km

composition: solid (no gasses, ices)

- 5/6 are “C-type” carbon rich
- 1/6 are “S-type” iron rich

Q: how do we know this?

Asteroid Orbits

asteroids orbit Sun → must follow Kepler's laws (of course!)

wide variation in a , e

but average $a \simeq 2.8$ AU

→ between Mars & Jupiter

average eccentricity $e < 0.1$: nearly circular

most orbit planes close to ecliptic

www: Inner Solar System in real time

www: Near-Earth objects

if cross Earth's orbit, enter atmosphere:

friction → heat & light → **meteor**! www: Leonids

www: fireball

if survive the fall: land a meteorite

Note: meteorites on display in Natural History Bldg

extraterrestrial matter! go look!

LSST: will identify $> 90\%$ of near-Earth asteroids > 140 m

Chelyabinsk Impactor

Feb 15, 2013: **meteor fall & impact over Chelyabinsk, Siberia**

- impactor ~ 50 m diameter!
- energy release ~ 500 kTon TNT

Yikes.

Debris recovered over fall area and impact on frozen lake

www: videos of fall

www: videos of sonic boom shockwave

- o shock: broken glass – 1,700 injuries

what do asteroids look like? From ground, see only largest
but now have visited some (on the way to outer SS)

www: Gaspara

www: Ida & Dactyl

Near Earth Asteroid Rendezvous

intercept near-earth asteroid 433 Eros

S-type: stony-iron

large: 35 km long, 14 km wide—Chicago

“potato shaped”

www: NEAR image of 433 Eros

www: NEAR movie

hints of stratification—broken from (much) larger object?

✓

www: Ceres — largest, round!

Debris II: Comets

www: Hale-Bopp, Hyakutake, Ikeya-Seki

last year: brightest comet in decades! www: McNaught

Comet Structure: “dirty snowball” nucleus: ~ 10 km

solid: ices (H_2O , CO_2 , CH_4), embedded dust grains

very elliptical orbits: changing $r \rightarrow$ changing T

far from Sun: completely frozen

as approach: ice \rightarrow vapor (sublimation)

dust, gas released $\rightarrow 10^6$ km coma

www: HST Hale-Bopp: coma & jets, nucleus unresolved

pressure from sunlight & solar “wind” of particles

∞
 \rightarrow tails: Ion, dust

iClicker Poll: Comet Tails

Cometary dust and ions (=ionized atoms)
feel pressure from sunlight, solar wind of particles
⇒ result in comet tail

Where do comet tails point?

- A away from the Sun
- B behind the comet (i.e., opposite comet's velocity vector)
- C tail direction is random

Comet Tails and Structure

ion tail: small, low-momentum particles

→ carried by solar wind

→ points *away from Sun*

dust tail: larger, higher-momentum particles

→ retain \vec{v} component in comet direction

→ non-radial *arc tracing comet path*

ESA (Euro-NASA) Mission: *Rosetta*

rendezvous with comet 67P/ Churyumov-Gerasimenko

- orbiter takes images

- lander *Philae* to take surface data (crashed)

density: $470 \text{ kg/m}^3 \ll \rho_{\text{ice}} \rightarrow$ *porous*

70 – 80% of volume is empty

surface dusty but irregular in shape and composition

when heated: dust storms

Comet Orbits and Populations

“Long Period”: $P > 10^5$ yr

→ $a > 2000$ AU!

all orientations → not just ecliptic

originate in **Oort Cloud**

spherical comet “reservoir” at 3000–100,000 AU

not observed directly!

probably did not form there....

ejected by Jovian planets in early SS?

“Short Period”: $P < 200$ yr

lie in ecliptic

→ not from Oort cloud → **Kuiper Belt**

≡ semimajor axes $a = 30 - 100$ AU

www: Outer solar system sketch

Trans-Neptunian Objects

first Kuiper belt object (KBO) detected in 1992
also known as *Trans-Neptunian* objects; today, tally is hundreds

typically \sim few% – 10% size of Pluto
probably formed where they are now
estimates: 70,000 KBO's
total mass $\sim 0.1M_{\text{Earth}}$

also: some comets strongly deflected,
have orbits with very small perihelion
(i.e., *very* close to Sun).

www: sun-grazing comets

A Strange Visitor

Discovered Oct. 19, 2017 by Pan-STARRS1 telescope in Hawai'i

- huge angular speed on sky: 6 deg/day
- orbit tracked, found speed $v > v_{\text{esc}}$
above Solar System escape speed!
- came close to the Sun, but no dust or gas found
- brightness changes by factor 10 every 7.3 hours
- red coloring, similar to outer solar system objects

Q: what does all of this mean?

Oumuamua

Reading the clues:

- $v > v_{\text{esc}}$: orbit is *hyperbolic* – unbound!
one-time visit: *origin outside of Solar System*
- no gas or dust when near Sun → not cometary → *rocky/metallic*
- brightness changes periodically by factor 10: *spinning*
cross section changes by factor 10: must be *very elongated*
most extreme elongation of any known asteroid
- red color → “*space weathering*” likely over Myr

name: **Oumuamua** www: artist's conception

Hawaiian for “a messenger from afar arriving first”

first known interstellar/extrasolar asteroid

LSST likely to find more!

Pluto

Orbit

- $a = 39.5 \text{ AU}$, $P = 285 \text{ yr}$
has not made a full orbit since discovery in 1930
- $e = 0.25$ – largest for any planet, crosses Neptune orbit

Properties

spherical shape. $R = 1151 \text{ km} \approx R_{\oplus}/6$

$\rho_{\text{avg}} \simeq 2000 \text{ kg/m}^3 \rightarrow \text{ice, rock}$

surface: N_2 and methane ice, coating water ice

atmosphere: very thin, $P = 10^{-5}$ earth

appears at aphelion, freezes and snows out at perihelion

July 2015: *New Horizons* flyby – first closeup look

- Pluto surface: mountains, valleys, plains
- very few craters! \rightarrow tectonic activity, possibly ice volcanos

www: New Horizons multimedia

Jan. 1, 2019 ring in New Year: close flyby of KBO 2014 MU69