

Astro 210
Lecture 24
March 14, 2018

Announcements

- **HW7 due online in PDF, Friday 5:00 pm**
- Office hours: instructor 2:00-3:00 pm today
TA: 3:30-4:30 pm tomorrow (yay!)
- **Night Observing this week**

Campus Observatory. Tonight, tomorrow **8–10 pm**

note later times—solstice approaches!

bring **report form** available on Moodle

take and submit **selfie** while there

Theory of Solar System Origin: Executive Summary

stars born in cold gas & dust clumps: molecular clouds

“gravitational collapse”: runaway contraction

angular momentum: centrifugal barrier to collapse

most matter \rightarrow proto-Sun

high-angular momentum matter: protoplanetary disk around sun

gas ρ , matter state (presence of ices) change with R

water/ice “snow” line at $R_{\text{snow}} \sim 3 \text{ AU}$:

Inner/Outer planet boundary!

Disk Gaps and Planets

As protoplanets/planets arise in a disk
their gravitation influence can clear out disk material
in the disk annulus around the planet's orbit radius

so the protosolar nebular theory predict: a protoplanet/planet
can create a *gap* in the disk

Note: small “shepherd” moons maintain gaps in Saturn's rings!

ω Q: *what observable predictions does Solar Nebula theory make
for young stars, mature planet-bearing stars?*

Testing Solar Nebula Theory

This is a special time!

Now seeing planets, planet formation around other stars

Solar Nebula theory should work generally

→ should apply to these systems too

...though some details might vary *Q: why?*

General Predictions of Solar Nebula Theory

In forming stars (protostars):

1. young protostars have gas disk
2. older protostars have planetesimal disk
3. protoplanets can open gaps in disks

In fully-formed star and planet systems:

1. small planets near star
2. massive planets farther away
3. orbits nearly circular
4. orbit planes near stellar equator

Problem: solar nebula theory built to explain
one data point (SS)! → is the model “fine-tuned”?

iClicker Twofer: Bets on Planet Formation

Vote your conscience!

Which prediction seem most solid to you?

- A young protostars have gas disk
 - B older protostars have planetesimal disk
 - C small planets near star
 - D massive planets farther away
 - E planet orbits nearly circular
-

○ In same list: which prediction seems least solid?

Test I: Young Stars

evidence from direct imaging:

50% – 100% of youngest stars surrounded by gas disk
disks are common and perhaps unavoidable!

www: Orion HST montage

www: protoplanetary disks in Orion

www: Orion disks set of 4

www: Orion disks side view (really disks)

disks thick, blocks light

→ enough material to make planets

→ agrees with Solar Nebula theory!

✓ evidence for disk formation!
Orion objects: “proplyds” = protoplanetary disks

Dusty Disks

Some older protostars and fully-formed have spectrum that has **two** peaks → two temperatures

- optical emission from the hot surface of star, and
- infrared emission from dust in disk!

Recently: can **image** the disks in the infrared

warm dust: small (mm-sized) particles heated by star → thermal IR

www: β Pic disk w/star

www: ALMA: HL Tau, TW Hyd

ALMA in Chile is most powerful mm-wave telescope

has revolutionized study of planet formation

- disks seen with multiple gaps!
- disks seen in binary protostars (each star has disk)

Solar Nebula Scorecard: Midterm Grades

General Predictions of Solar Nebula Theory

In forming stars (protostars):

1. young protostars have gas disk? **check!**
2. older protostars and fully formed stars have particle-bearing disk? **check!**
3. gaps found in some disks? **check!**

Solar Nebula Theory status:

Woo hoo! so far so good!

theory works up through disk formation

how about planets themselves?

recall Solar Nebula predictions:

- giant planets far from stars
- rocky planets close in
- orbits nearly circular
- orbit planes near stellar equator

Test II: Exoplanets

Exoplanets = extra-solar planets = planets around other stars

★ have been sought for centuries!

★ first positive, definitive detection: 1994 (around dead star)

★ first detection around normal star: 1995

What took so long?

Exoplanet detection is a huge technical challenge

Q: Why?

Q: possible workarounds?

Challenges for Planet Hunters

Can't “just look” – glare!

feeble light from planet drowned out by star flux

→ need a more clever workaround

Several detection techniques have been proposed
three of these have already borne planetary fruit!

Successful strategies thus far involve:

- look for planet(s) effect on host star
- get lucky
- both of the above

Planet Effects on Host Stars: Reflex Motion

recall Newton III: since star exerts gravity force on planet
planet *must* exert *same* force on star!

- *both* must accelerate! the star moves (“reflex motion”)
...but $a = F/m \rightarrow a_{\star} \ll a_{\text{planet}}$
- both stars and planet orbit fixed “center of mass”

thus:

- the star moves too!
- what remains fixed is the center of mass
a point on the line connecting the star and planet

consider two objects of equal masses $m_1 = m_2$

↳ Q: *where is center of mass?*

Q: *how do distances r_1, r_2 to COM compare?*

Center of Mass Reminder

Newton II: $a \propto F/m$

+ Newton III: $F_{p \text{ on } \star} = -F_{\star \text{ on } p}$

$\Rightarrow F$ magnitude same, heavier object accelerated less

\Rightarrow star moves slower, nearer to COM

- distances to center of mass:

total star-planet distance: $r_{\star} + r_p = a$

and $m_{\star}r_{\star} = m_p r_p$

so: $r_{\star}/r_p = m_p/m_{\star} \ll 1$

\Rightarrow and so $v_{\star} \ll v_p$ but $\neq 0$!

How to use this?

- in practice, **can't** track star orbit path – too small on sky
- but **can** look for “wobble” in star speed

1995: detection!

Planet Detection by Good Luck: Transits

if very lucky, planet orbit plane seen **edge-on**
so planet sometimes passes in front of star
★ causes partial eclipse of host star!
★ star dimming small but observable

Strategy: monitor light from candidate stars
look for brightness changes
as planet crosses (**“transits”**) star's disk

Q: What is expected “light curve” of flux $F(t)$ vs time t ?

Q: How to verify signal was due to a planet?